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Comparison of Flight and Ground Tests of Environmental Degradation of MISSE-6 SUSpECS Materials

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Abstract

The effects of prolonged exposure to the LEO space environment and charge-enhanced contamination on optical, thermal, and electron emission and transport properties of common spacecraft materials has been investigated by comparing preand post-flight characterization measurements. The State of Utah Space Environment & Contamination Study (SUSpECS) deployed in March 2008 on board the Materials International Space Station Experiment (MISSE-6) payload, was exposed for ~18 months on the exterior of the International Space Station (ISS), before retrieval in September 2009. A total of 165 samples were mounted on three separate SUSpECS panels on the ram and wake sides on the ISS. Electron-, ion-, and photon-induced electron emission yield curves, crossover energies and emission spectra, resistivity, dielectric strength, optical and electron microscopy, UV/VIS/NIR reflection spectroscopy, and emissivity were tested for pre-flight SUSpECS samples in their pristine conditions.

MISSE Objectives

The purpose of MISSE is to characterize the performance of new prospective spacecraft materials when subjected to the synergistic effects of the space environment.

Launch









USU Electron Emission/Resistivity Samples Au [R,WB], AI [R,WB], 316 SS [R,W*], Au/Ni on Cu [R,W*] Dupont Black Kapton [R,WB], Sheldahl Thick Film Black [R,WBI, Aquadag [R,W* PI-Kapton [R,W], PET-Mylar [R,W], PTFE-Teflon [R,W] FR4 [R,W], DC93-900 RTV on Cu [R,W] Ce0₂-doped glass solar cell cover slips [R,W], A1₂0₃-Alumina [R,W] Anodized AI (Cr) [R*,W], Anodized AI (S) [R*,W] Cu [R*,W*], Graphitic amorphous carbon [R*,W*] Nylon [R*,W*], Kevlar [R*,W*I, ITO on Kapton [R*,W*], SiO₂ [R*,W*] UV-coated CC0₂-doped solar cell cover slips [R*,W*] CV-1147 RTV on Cu [R*,W*]

Charge-enhanced Contamination Samples Au [R,WB], AI [R,WB], Dupont Black Kapton [R,WB] Sheldahl Thick Film Black [R,WB]

CRRES Samples PI-Kapton [R,W], PET-Mylar [RW], PTFE-Teflon [R,W] FR4 [R,W], A1₂0₃-Alumina [R,W], Si0₂ [R*,W*]

ISS Samples Au [R,WB], AI [R,WB], 316 SS [R,W*], Dupont Black Kapton [R,WB], PI-Kapton [R,W], DC93-900 RTV on Cu [R,WI, A1₂0₃-Alumina [R,W] Ce0₂-doped glass solar cell cover slips [R,W] Anodized AI (Cr), [R*,W], Anodized AI (S), [R*,W] UV-coated Ce0₂-doped solar cell cover slips [R*,W*] CV-1147 RTV on Cu [R*,W*I

ATK and USU Space Dynamics Lab Composite Samples COIC S400 COIC S200H and COIC S300 Nonoxide Ceramic-Metal Composites SDL/GIFTS Reinforced Carbon Nono-fiber/RS-3 Cyanate Ester Composite [W] SDL PTFE [RWB]

Solar Probe Mission (JHU/APL) Samples Plasma sprayed Al2O3 on Carbon/Carbon Composite [W] Al2O3 [W[, Pyrolytic Boron Nitride on Carbon/Carbon Composite [W] Pyrolytic Boron Nitride [W], Barium Zirconium Phosphate [W], Carbon/Carbon Composite [W]

JWST Samples PTFE [W], Kapton HN [W]. Tefzel [W], Hytrel [W] Kapron XC100 [W], Kapton E [W][R,WB]

Sample Sources

- structure.
- charging and arching. probe for ISS.

Wake Side

- 13 Grounded Samples 12 Biased Samples: for 3 sets of 4 samples with low current biases for chargeenhanced contamination studies.
- 6 Concealed samples

Sample Holders

• Holder area 5 cm x 15 cm • 9 mm diameter exposed sample area

Ram Side

 25 Exposed samples on each of 3 tiers for varying AO and UV exposure. 23 Concealed samples

Wake Side

- 25 Grounded Samples
- 10 Concealed samples

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MISSE-6 Time Table

Sample selection completed PEC's completed and tested for flight readiness Launch on Space Shuttle ISS-123 **Return of samples from space**



SUSpECS Test Samples and Mounting

Wide array of common spacecraft materials (see left). Basic materials and key contaminants of ISS solar arrays and

Materials from CRRES satellite designed to study environment-induced

• Materials used in Floating Potential Measurement Unit (FPMU) plasma

• Critical thermal control and optical materials for SDL GIFTS payloads. • Composite and ceramic materials of the ATK Thermal Protection Systems (TPS) and the ATK Lightweight Structure Systems (LSS). • James Web Space Telescope Insulator Sample Charging Tests. Solar Probe Mission Heat Shield Insulator Samples tests.



2.25"

SUSpECS III Sample Holder





The second comparison reported here focused on four materials [carbon-loaded polyimide, polyester (Dupont Mylar), Al_2O_3 (sapphire), and SiO_2 (quartz)] that showed some of the most pronounced environmentally-induced changes in optical properties. Samples of each material on the wake and 3-tiered sample panels were exposed to a complex environment during the flight. Identical witness samples were also exposed to a simulated subset of the environment in the Characterization of Combined Orbital Surface Effects (CCOSE) space environment test chamber at the USAF Arnold Engineering Development Center to mimic the space exposure profile. The primary optical characterization methods employed for the comparison were UV/VIS/NIR and FTIR transmission of the sapphire and quartz and UV/VIS/NIR reflectance of the polyimide and polyester. Comparison of pre-flight, post-flight, and simulated exposure samples served two primary purposes: to investigate the validity of simulated environmental testing methods and to help distinguish the effects of specific components of the complex space environment that samples were simultaneously exposed to during the flight.

Analysis of constraints limited the selection of samples to those on SUSpECS II wake panel, primarily due to a maintenance overhaul of the CCOSE atomic oxygen source. In order to verify that all samples did not change in the same way, exhibited a variety of optical property changes on orbit. The final list was as follows: Mylar for having highly varying optical properties, Black Kapton for minor change, and quartz and sapphire for no change. These samples are shown in the Figures at right.

The primary constituents of the ISS orbit are vacuum, solar, atomic oxygen, protons and electrons. Of these, CCOSE could replicate all but the atomic oxygen. For the other components, the environment could be replicated if on-orbit data were available. Fortunately, the equivalent sun hour (ESH) estimates were provided by the Boeing ISS Thermal Analysis group. Utah State provided the temperature time history data from NASA. NASA had Boeing-supplied Lithium Fluoride (LiF) thermoluminescent dosimeters (TLDs) radiation detectors located behind some of the samples on the MISSE-6 mission. These detectors output total radiation dose measurements.

In order to properly set the CCOSE electron source, the CCOSE test profile was to divide up the electron spectrum and map it to the electron gun output. Due to the monoenergetic nature of the electron gun output, two different beam energies were selected to represent the full particle spectrum. (Figure-Top Right). monoenergetic beam of 15 keV electrons was run to simulate the lower portion of the spectrum while a 60 keV beam was used for the upper. This resulted in a run time of 129 hrs of 15 keV and 39 hrs of 60 keV.

Pre-test reflectance and transmittance measurements were performed at atmospheric conditions using both a Jasco UV-VIS-NIR spectrometer and a Bruker FTIR spectrometer. Post-test reflectance and transmittance measurements were also performed with these instruments. The following results are based on initia assessment of the data and will be reviewed in greater detail in a future publication. The initial analysis follows the trend of the MISSE-6 SUSPeCS samples.

 As stated earlier, quartz and sapphire were selected for their resistance to any change. The figure at right shows the transmission measurements made before and after testing. Minimal changes were observed in the UV region with these samples but appeared primarily unchanged.

• Black Kapton had some small changes on the orbital samples but had little change in the ground test (Figure-Middle Right). This was likely due to interaction with AO on orbit and lack of AO in the ground test.

The most noticeable change in both the flight and ground samples occurred on the aluminized Mylar (Figure-Bottom Right). Most of the loss appears within the visible region from 300 to 700 nm. This change was likely due to vacuum ultraviolet radiation severing the polymer bonds, but may have been enhanced by the electron beam as other researchers have found degradation in Mylar under energetic electron bombardment.

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ATK THIOKOL INC.



Space Environment Simulation at ADEC



Acknowledgements





1000 1250 Wavelength, nm

1500