



Characterizing Noise Trends of the Cross-Track Infrared Sounder (CrIS)

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SDL/15-915



Outline

- CrIS Summary
- Noise characterization process
- Noise trending results
- Recent Changes
- Conclusions

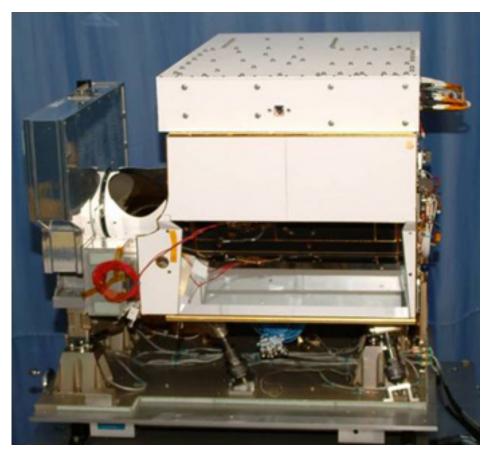


http://www.jpss.noaa.gov/images/media-gallery/gallery-jpss1_1.jpg



Cross-Track Infrared Sounder (CrIS)





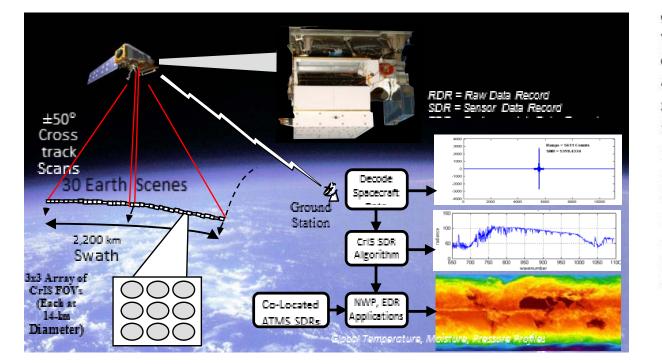
http://www.jpss.noaa.gov/images/instruments-cris.jpg

- The CrIS sensor is an infrared Fourier transform spectrometer
- One CrIS instrument is currently flying on the Suomi National Polar-orbiting Partnership (NPP) spacecraft
- Another CrIS (J1) is preparing for launch in 2017

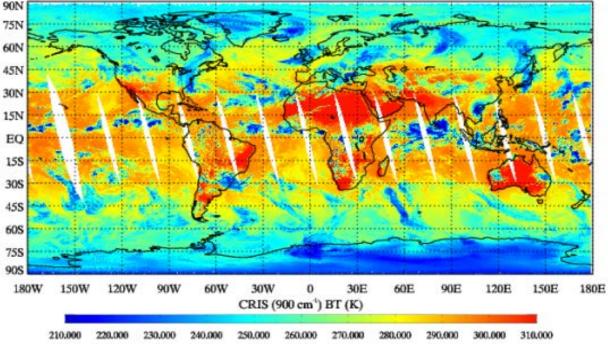


What it does





Ascending_orbits: CRIS (900 cm⁻¹) BT (K) Date: 2012-10-28



• CrIS takes optical readings of the atmosphere





How it works



https://directory.eoportal.org/web/eoportal/satellite-missions/n/npoess

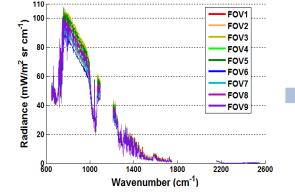
Interferometer

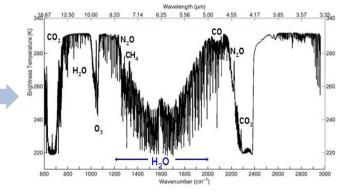
- CrIS is an infrared Michaelson Interferometer
- Its adjustable mirror arrangement causes interference, and the sensor detects varying light intensities

1500 500 -500 -150 -1500

Interferogram

- The magnitude of light intensity is recorded several times as the mirror moves either forward or backward
- Data is sent to earth in packets





Spectra

• From the calibrated Fourier transform, the wavelengths of light that interfered with each other to make the interferogram can be determined

Use

- Data is converted into atmospheric profiles of temperature, pressure, and water vapor
- Information helps predict the weather, document trace gasses, protect sea wildlife*, and monitor climate change

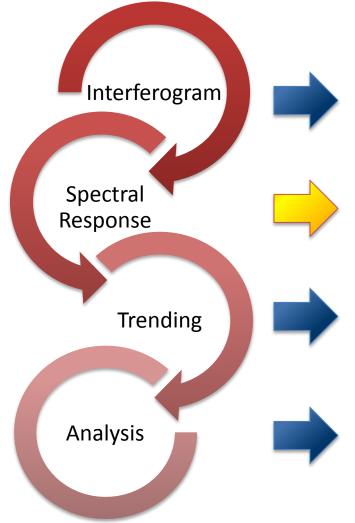
*http://www.jpss.noaa.gov/media. html?story=news-62



• CrIS data helps determine how the atmosphere behaves

What SDL Does





Independently generate spectrograms for cross checking.

Compare deep space and internal calibration, Earth scene, real, imaginary, total, correlated, and random data, daily and trending data.

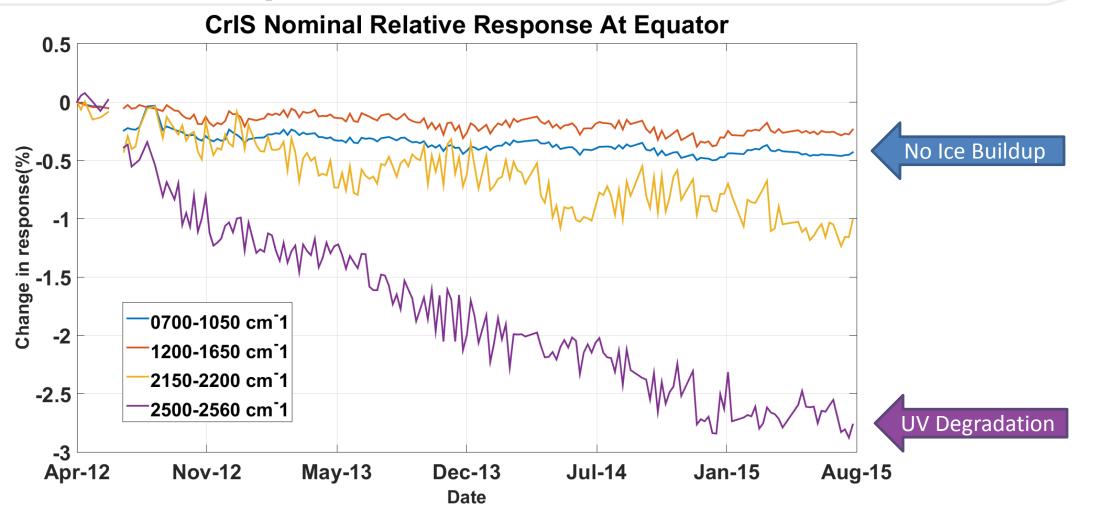
Log maneuvers, radiation events, gain decay, and anomalies.

Compare different algorithms, coefficients, side1/side2 electrical systems, different satellites. Perform PC analysis, instrument optimization, and time trending.





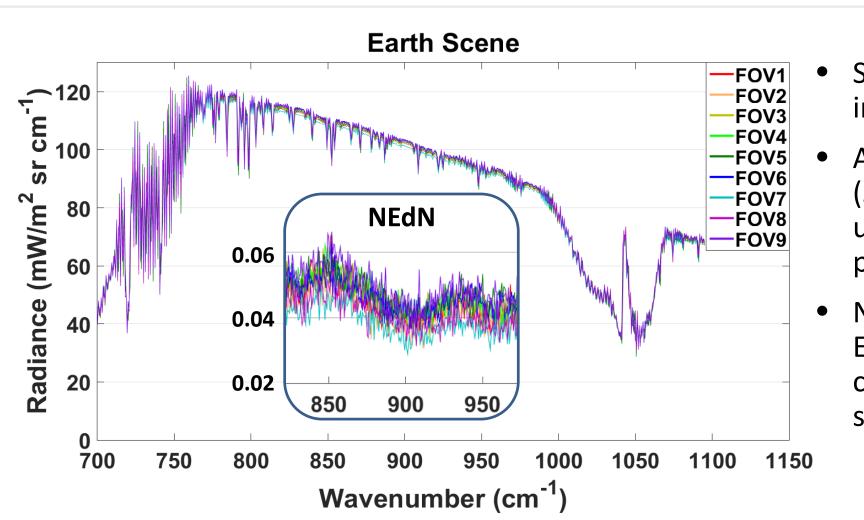
CrIS Relative Response over Time



• CrIS response has only dropped 3 percent in 3 years



Noise





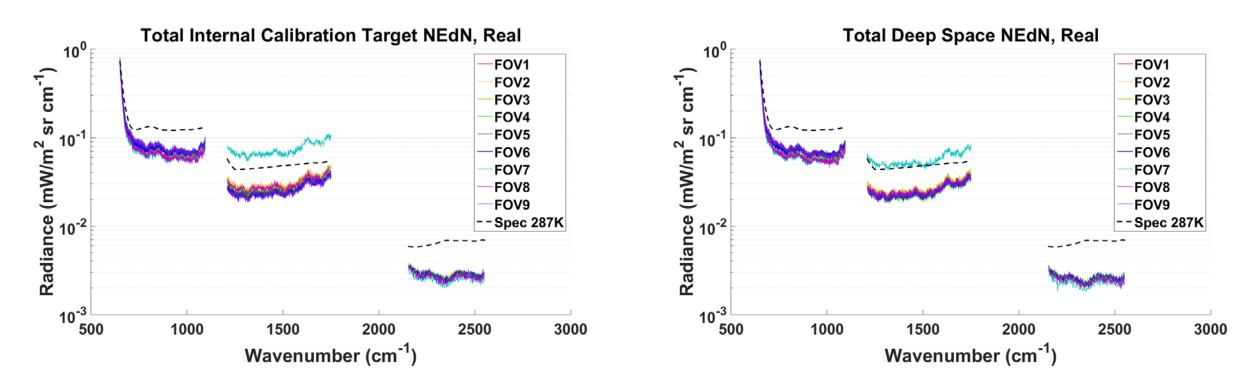
- SDL monitors the instrument's noise trending
 - A higher signal-to-noise ratio (small NEdN) means more useful information can be pulled from the data
 - Noise results reported for
 Earth scene, internal
 calibration target, and deep
 space scenes



• The NEdN is very low for the CrIS instrument

Deep Space and Internal Calibration Target Noise Profiles



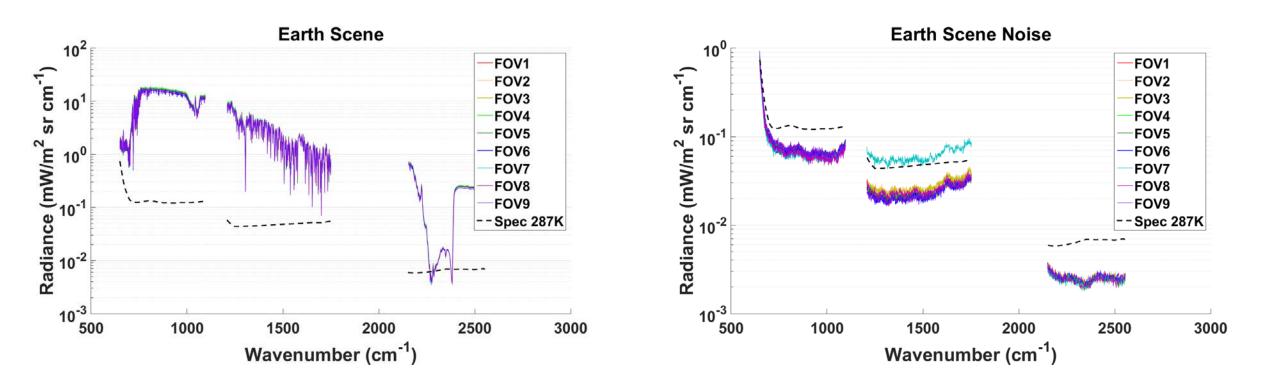


- Noise results broken into total noise, correlated noise, uncorrelated components
- Total noise derived directly from measurement (std-dev of multiple scenes)
- CrIS meets specification in all fields of view (FOV) except FOV 7





Noise Correction for Earth Scene



- Uncorrelated (random) and correlated noise derived from principle component analysis (PCA)
- With PCA, we can see NEdN in Earth scenes

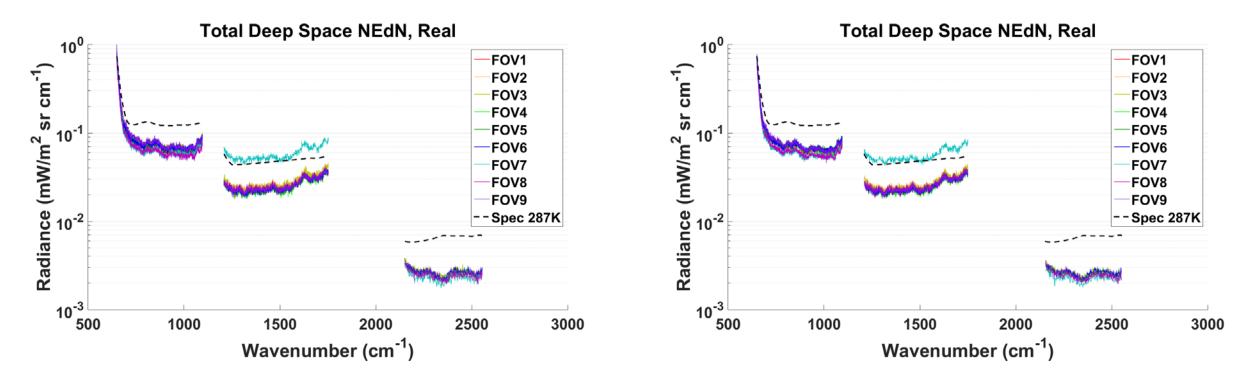




NEdN not Significantly Different between Algorithms

Processed by ADL

Processed by SDR Generator

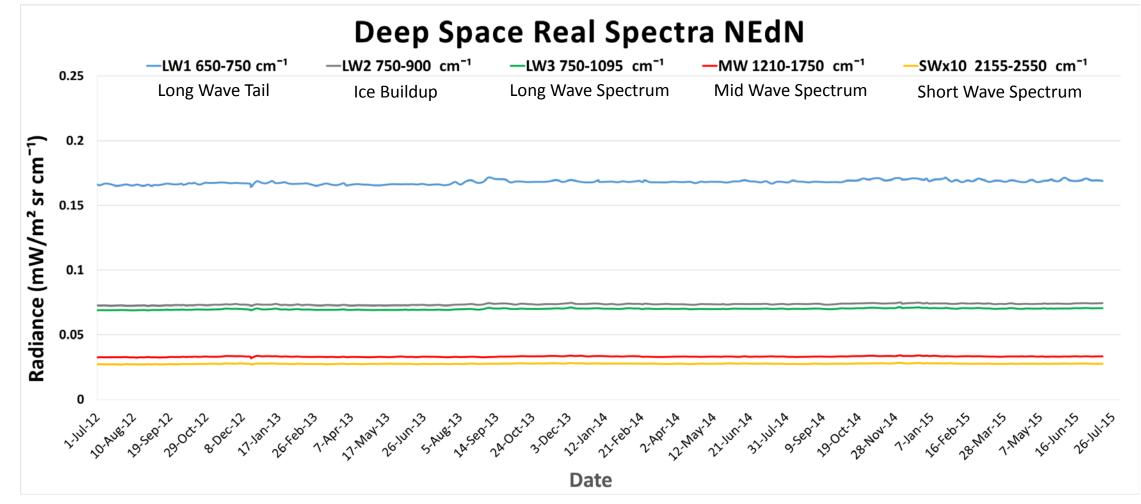


• NEdN is processing independent



Real Trending





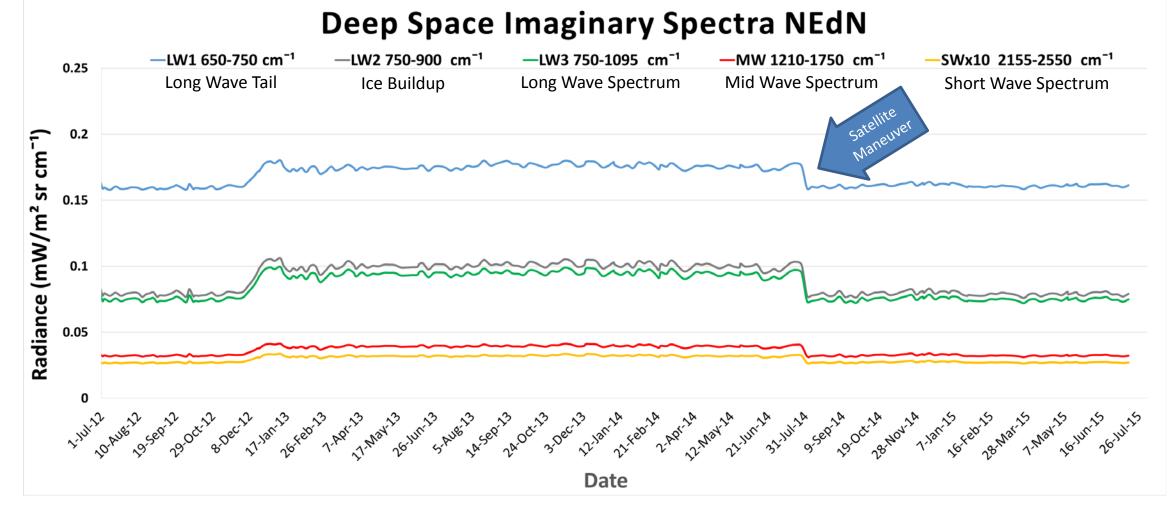
• Worst case, NEdN has increased only 0.4%







Imaginary Trending

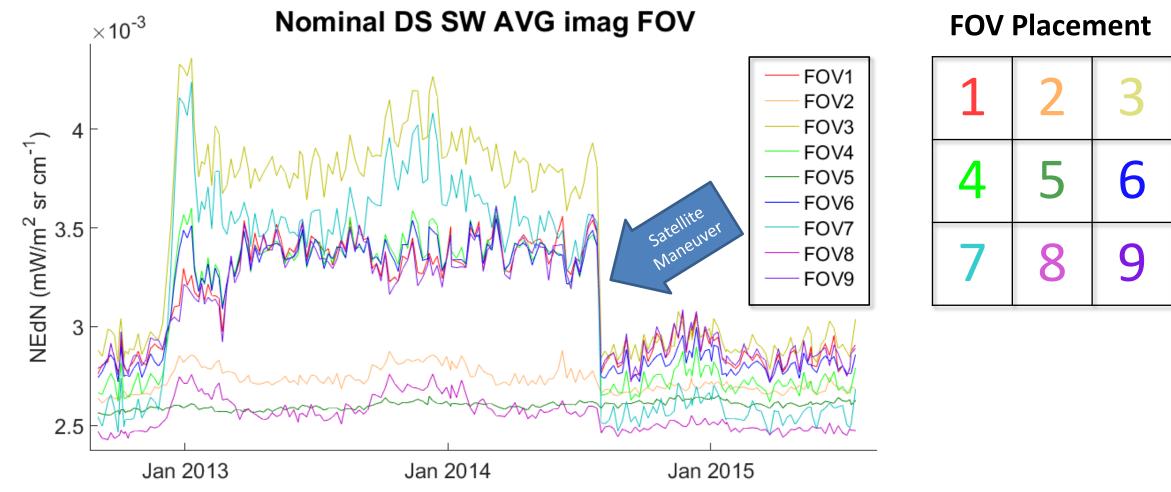


• More variation in imaginary spectra





Independent FOV Trending

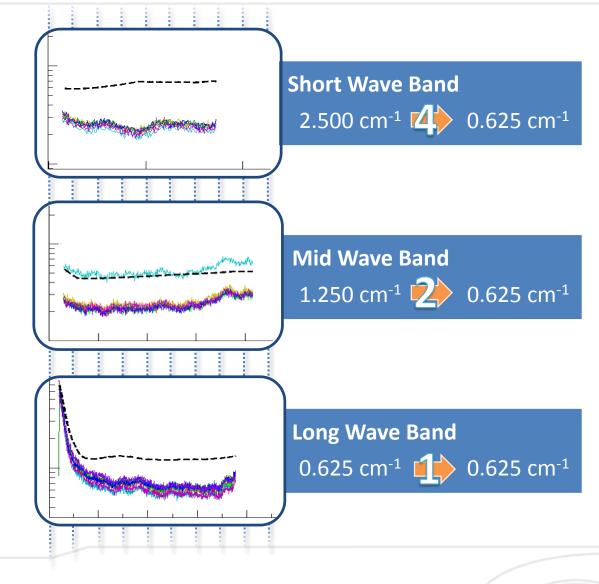


• FOVs not affected equally by satellite jitter/maneuvers



New Developments



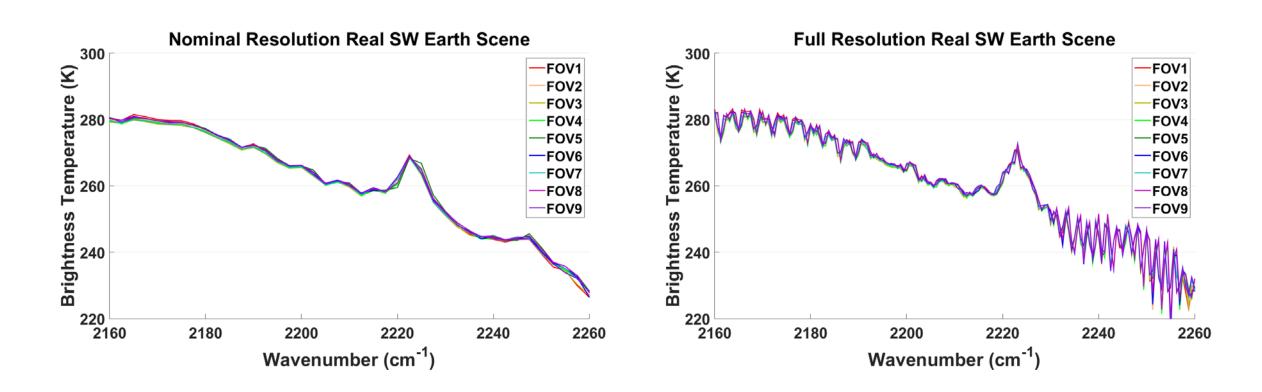


- During the past year, CrIS was reconfigured to enhance its spectral resolution
- Now all three spectral bands CrIS analyzes are sampled at 0.625 cm⁻¹
- CrIS now sends more data





Improvements – More Information

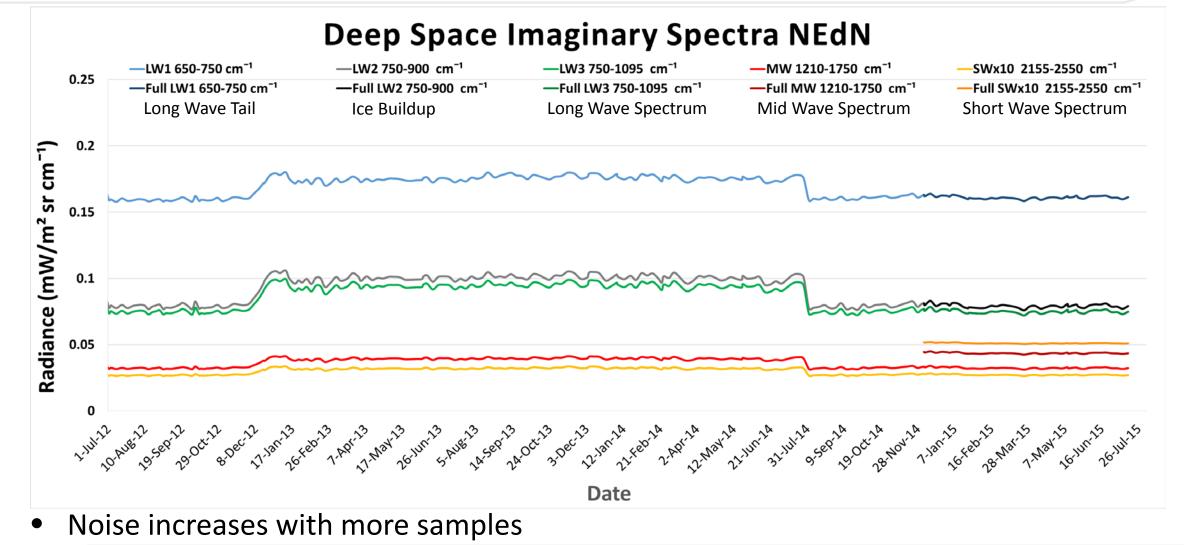


• Full resolution provides more data





Challenges – Increased Noise

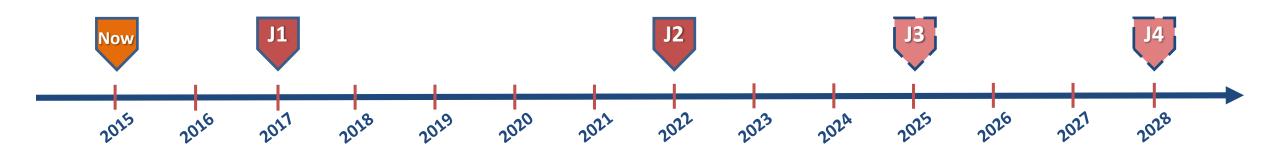






A Look Ahead

- Assess J1 noise on orbit
- Compare with ground measurements to ensure consistency
- Compare NPP and J1 noise
- 2 CrIS instruments in space by 2017







Conclusion

- Response degrades slowly
 - Spectral response will last the life of the instrument.
- NEdN is very low
 - Over the life of the instrument, it has increase by only 0.4%
- Noise is algorithm Independent
- Imaginary NEdN helps characterize payload
- Improved resolution provides more data



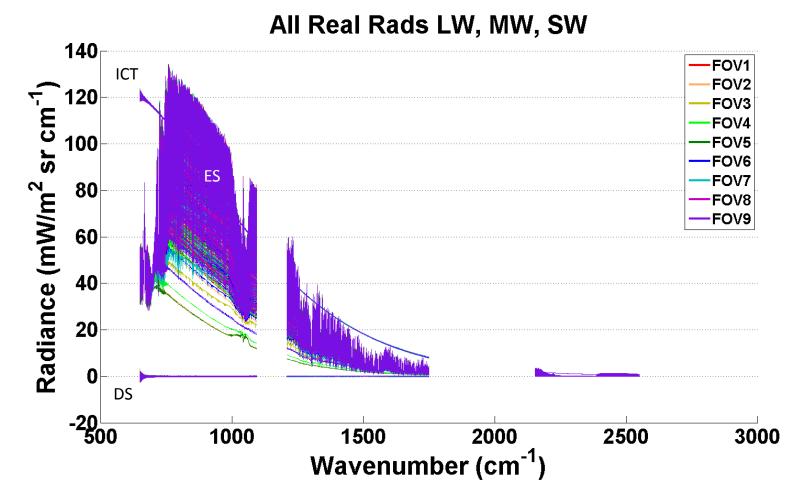
ADDITIONAL SLIDES



20

Deep Space, Earth Scene and Internal Calibration Target



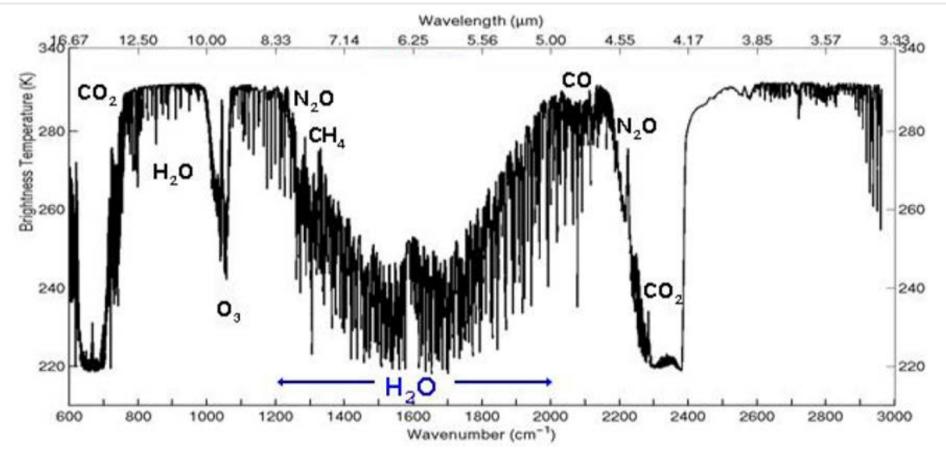


• Earth Scenes are much noisier





Spectral Signatures of Atmospheric Gasses



• Different gasses absorb different wavelengths

