Dynamics of orographic gravity waves observed in the mesosphere over Auckland Islands during the Deep Propagating Gravity Wave Experiment (DEEPWAVE)

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Dynamics of Orographic Gravity Waves Observed in the Mesosphere over the Auckland Islands during the Deep Propagating Gravity Wave Experiment (DEEPWAVE)

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ABSTRACT

On 14 July 2014 during the Deep Propagating Gravity Wave Experiment (DEEPWAVE), aircraft remote sensing instruments detected large-amplitude gravity wave oscillations within mesospheric airglow and sodium layers at altitudes \( z = 78-83 \) km downstream of the Auckland Islands, located \( -1000 \) km south of Christchurch, New Zealand. A high-altitude reanalysis and a three-dimensional Fourier gravity wave model are used to investigate the dynamics of this event. At 0700 UTC when the first observations were made, surface flow across the islands' terrain generated linear three-dimensional wave fields that propagated rapidly to \( z = 78 \) km, where intense breaking occurred in a narrow layer beneath a zero-wind region at \( z = 83 \) km. In the following hours, the altitude of weak winds descended under the influence of a large-amplitude migrating semidiurnal tide, leading to intense breaking of these wave fields in subsequent observations starting at 1000 UTC. The linear Fourier model constrained by upstream reanalysis reproduces the salient aspects of observed wave fields, including horizontal wavelengths, phase orientations, temperature and vertical displacement amplitudes, heights and locations of incipient wave breaking, and momentum fluxes. Wave breaking has huge effects on local circulations, with inferred layer-averaged eastward flow accelerations of \( -350 \) m s\(^{-2}\) h\(^{-1}\) and dynamical heating rates of \( \sim 8 \) K h\(^{-1}\), supporting recent speculation of important impacts of orographic gravity waves from subantarctic islands on the mean circulation and climate of the middle atmosphere during austral winter.

1. Introduction

The Deep Propagating Gravity Wave Experiment (DEEPWAVE) was a field measurement campaign to observe the end-to-end dynamics of gravity waves—generation, propagation, breakdown, and effects on large-scale circulations—at altitudes from the ground to \( \sim 100 \) km. The primary observational platform was the National Science Foundation (NSF)/National Center for Atmospheric Research (NCAR) Gulfstream V research aircraft (NGV: Laursen et al. 2006), which for DEEPWAVE was equipped with a suite of in situ and remote sensing instruments to observe gravity wave dynamics over this deep altitude range. The field phase of DEEPWAVE took place from May to July 2014 from an operating base in Christchurch, New Zealand, which provided regular NGV access to major orographic and nonorographic sources of gravity waves, while the strong vortex-edge winds in austral winter provided a stable propagation channel for deep propagation of these waves into the mesosphere and lower thermosphere (MLT). Fritts et al. (2016) review the planning, execution, and initial results of DEEPWAVE.

One of the many scientific objectives of DEEPWAVE was to acquire gravity wave observations to test recent ideas that gravity waves generated by small island
terrain in the Southern Ocean significantly influence the large-scale momentum budget of the middle atmosphere in austral winter. This idea first arose when Alexander et al. (2009) analyzed radiances acquired by the Atmospheric Infrared Sounder (AIRS) on the Aqua satellite in September 2003 and found three-dimensional gravity waves in the upper stratosphere emanating from the small subantarctic island of South Georgia. Alexander et al. (2009) inferred significant momentum flux deposition from these waves and, since global models typically treat grid cells containing small islands as ocean rather than land, they speculated that omission of gravity wave drag from small subantarctic islands might explain some or all of the stratospheric “cold pole” bias in global models during austral winter (Butchart et al. 2011).

McLandress et al. (2012) tested this idea by inserting artificial subgrid-scale orography into their model’s orographic gravity wave drag parameterization near 60°S. This small change substantially reduced the model’s cold-pole bias in austral winter but somewhat degraded its ozone simulations. Alexander and Grimsdell (2013) later extended the AIRS analysis to other subantarctic islands and different austral winters, finding deep gravity waves to be common over many of the islands they studied, and that estimated wave momentum fluxes could explain some, but perhaps not all, of the cold-pole drag deficit inferred from modeling studies such as McLandress et al. (2012).

Nadir satellite imagers such as AIRS currently provide the only routine observations of deep gravity wave activity from these remote islands. Unfortunately these gravity wave perturbations occur near (and in some cases beyond) the resolution and noise limits of the sensor, so our knowledge of deep gravity wave dynamics over islands remains fragmented. Likewise, global models cannot resolve these waves and presently resort to simple augmentations to existing parameterizations to assess potential effects (McLandress et al. 2012). More recently, high-resolution linear and nonlinear models have been used to model orographic gravity waves from South Georgia, revealing deep propagation of waves into the stratosphere and drag effects due to wave breaking on the stratospheric circulation (Jiang et al. 2014; Vosper 2015). However, similarly extensive high-resolution observations of these island wave fields were not available to validate these simulations.

DEEPWAVE provided an opportunity to use the instrumented NGV to acquire detailed observations of deep gravity wave dynamics over subantarctic islands for the first time and potentially to higher altitudes. Thus, ahead of the mission, islands south of New Zealand were studied as potential measurement candidates. The Auckland Island archipelago received particular scrutiny, because, as shown in Fig. 1, it is located ~1000 km south-southwest of Christchurch, well within flight range of the NGV, and is lined by significant mountain ranges with individual peaks over 600 m high. For simplicity, in this paper we will refer to all the islands of this archipelago, including Auckland Island, Adams Island, and surrounding islets depicted in Fig. 1, collectively as “Auckland Island.”

The AIRS analysis of Alexander and Grimsdell (2013) was unable to distinguish gravity waves emanating from Auckland Island and southern New Zealand, raising doubts about whether Auckland Island produced any deep gravity wave activity to observe. Thus, ahead of DEEPWAVE a 9-yr climatology of stratospheric gravity wave amplitudes over the greater New Zealand region was recomputed from AIRS data using modified analysis and averaging algorithms and a different collection of AIRS thermal radiance channels to improve signal to noise and geographical resolution [following Eckermann and Wu (2012) and Eckermann et al. (2016b, unpublished manuscript)]. The resulting variance maps, shown in Fig. 1a of Fritts et al. (2016), revealed a small but statistically significant peak in AIRS variances immediately downstream of Auckland Island that was distinct from larger peaks observed near New Zealand, identifying Auckland Island as a source of deep gravity wave activity and thus a viable measurement target for the NGV during DEEPWAVE.

On 14 July 2014, DEEPWAVE NGV research flight number 23 (RF23) overflew Auckland Island on four separate occasions and observed orographic gravity waves in the MLT downstream of the island terrain (Fritts et al. 2016; Pautet et al. 2016). Here we conduct a detailed investigation of the dynamics of these observed wave fields. Section 2 reviews the RF23 flight plan, the NGV instruments that observed the MLT, and a 0–100-km
DEEPWAVE reanalysis, which provides backgrounds for a Fourier model for computing three-dimensional linear orographic gravity wave solutions. In section 3 we investigate the dynamics of MLT wave fields observed during the outbound flight leg at \(\sim 0700\) UTC. Section 4 focuses on the time evolution of winds and waves in the MLT over Auckland Island during the \(\sim 7\)-h duration of the flight, particularly the influence of a large-amplitude semidiurnal tide in the MLT reanalysis. This provides context for our investigation in section 5 of observed wave field dynamics during the inbound flight legs that occurred \(\sim 3-4\) h after the initial outbound observations. In section 6 we review our findings and methods, then assess broader ramifications of these observed wave fields for the momentum and energy budgets of the middle atmosphere.

2. Observations, analysis, and modeling tools

a. DEEPWAVE RF23 observations

1) FLIGHT PLAN

Fritts et al. (2016) review the numerical weather prediction (NWP) models and ancillary tools used to inform DEEPWAVE flight planning. In the days prior to 14 July 2014, the NWP models predicted southwesterly near-surface flow upstream of Auckland Island intensifying to \(\sim 20\) m s\(^{-1}\) (see Fig. 2), with wind speeds increasing with height into a strong southwesterly tropospheric jet. High-resolution regional forecasts centered over Auckland Island using the U.S. Naval Research Laboratory (NRL) Coupled Ocean–Atmosphere Mesoscale Prediction System (COAMPS; Hodur 1997; Doyle et al. 2011) and Mountain Wave Forecast Model (Eckermann et al. 2006b) predicted wave generation and penetration of orographic gravity waves into the stratosphere.

Accordingly, the RF23 flight track was devised, among other things, to profile possible deep orographic gravity waves over Auckland Island. Figure 3 provides a three-dimensional representation of the flight track executed by the NGV. After taking off from Christchurch at 0545 UTC and flying south at 40 kft (12.2 km), the first transect of Auckland Island occurred from 0650 to 0715 UTC at \(\sim 12\)-km altitude. The first dropsonde (D1) was deployed from the NGV at 0713 UTC to measure the upstream atmospheric environment using the Airborne Vertical Atmospheric Profiling System (AVAPS; Young et al. 2014). Thereafter the NGV flew southwest to Macquarie Island (not shown), before returning to Auckland Island \(\sim 3\) h later and performing three more transects of the island, first at \(\sim 4\) km, then at \(\sim 7.5\) km, and finally at \(\sim 12\) km, with two additional drop sondes deployed upstream and downstream. Figure 2b reveals that (as planned) NGV transects across Auckland Island were close to parallel to the forecast upstream surface flow.

2) MLT OBSERVATIONS

Gravity waves were observed in the MLT near Auckland Island using two remote sensing instruments on board the NGV. Fritts et al. (2016) and Pautet et al. (2016) review this instrumentation and the MLT gravity wave measurements during RF23, and so only a brief summary is provided here. Later sections investigate the observed MLT gravity wave characteristics in detail.

The Advanced Mesospheric Temperature Mapper (AMTM) measured the temperature-sensitive infrared \((3, 1)\) Meinel band of hydroxyl, which during RF23 emitted from a 7-8-km-thick mesospheric layer peaking at a height \(z \sim 83.5\) km (Pautet et al. 2016). Rotational temperatures were retrieved from airglow brightness acquired by a zenith-viewing imager using the algorithm set forth in Pautet et al. (2014). Two additional “wing” cameras imaged airglow brightness on either side of the aircraft: although temperatures were not retrieved from these observations, for simplicity we refer to all three imagers collectively as the AMTM from this point forward. MLT sodium (Na) abundances were also measured from the NGV using a zenith-pointing 589-nm lidar. Additional details on this lidar system can be found in Bossert et al. (2015).

b. NAVGEM reanalysis

To model the evolution and dynamics of gravity waves observed in the MLT during RF23, accurate wind and temperature profiles are needed from the surface to the MLT upstream of Auckland Island. Global analyses issued by operational centers provide the best estimates of local winds and temperatures owing to their assimilation of available observations, which near Auckland Island come from satellite overpasses in addition to the three NGV AVAPS profiles (see below). These production data assimilation (DA) systems currently extend only to \(\sim 50-70\)-km altitude, leaving a critical analysis gap from \(\sim 60\) to 100 km, where waves can be significantly refracted, filtered, or dissipated.

Thus, a 0–100-km reanalysis project for the entire DEEPWAVE austral winter was undertaken using a high-altitude configuration of the Navy Global Environmental Model (NAVGEM). This system and its DEEPWAVE reanalysis products are the subject of a separate paper (Eckermann et al. 2016a, unpublished manuscript), and so only a brief overview relevant to the current study is provided here. NAVGEM is the U.S. Navy’s operational global NWP system, comprising a forecast model based on a 3-time-level semi-implicit
semi-Lagrangian discretization of the fluid equations on the sphere, coupled to a four-dimensional variational (4DVAR) DA algorithm. Hogan et al. (2014) provide a detailed description of the key model and DA components.

In common with other operational DA systems, the current operational NAVGEM has a rigid upper boundary at 0.04 hPa ($z \sim 70$ km; Hogan et al. 2014). For the DEEPWAVE reanalysis, the forecast model was reconfigured from 60 to 74 levels (L74) with a new upper boundary at $6 \times 10^{-5}$ hPa ($z \sim 115$ km), then augmented with a range of additional physical parameterizations needed to model the MLT. The DA system continued to assimilate its regular extensive suite of archived operationally available observations spanning the 0–50-km-altitude range. Among these were the AVAPS dropsonde data that, after rapid postprocessing in Christchurch, were transmitted on WMO’s Global Telecommunication System (GTS) for access by operational centers. From 50 to 100 km, the system assimilated the following additional observations:

1) temperature-sensitive microwave radiances from the upper-atmosphere sounding (UAS) channels of the

![Figure 2](image-url)
Special Sensor Microwave Imager/Sounder (SSMIS) on four Defense Meteorological Satellite Program (DMSP) polar orbiters (F16–F19 inclusive), using the NAVGEM UAS radiance assimilation procedures described in Hoppel et al. (2013);

2) limb measurements of temperature and water vapor up to 0.002 hPa from version 3.3 retrievals of the Microwave Limb Sounder (MLS) on the Aura satellite and of temperature up to $10^{-4}$ hPa from version 2.0 retrievals of the Sounding of the Atmosphere Using Broadband Emission Radiometry (SABER) instrument on the Thermosphere Ionosphere Meso- sphere Energetics and Dynamics (TIMED) satellite, using assimilation procedures for these instruments described in Eckermann et al. (2009).

Static bias correction profiles for MLS and SABER were updated from those in Eckermann et al. (2009) based on mean biases computed for coincident profiles for the period 2004–13 using version 3.3 MLS and version 2.0 SABER temperatures. At heights above 0.002 hPa SABER temperatures were assumed to be unbiased.

The system was initialized on 23 March 2014, roughly two months prior to the core May–July DEEPWAVE period, and run out to the end of September. The March–April period was treated as a DA spinup phase.

For this study, given the focus on large-scale background conditions controlling wave propagation, we used reanalyses from a low-horizontal-resolution re-analysis run using a T119 forecast model (outer loop) coupled to a T47 tangent linear forecast model (TLM, inner loop) used by the 4DVAR DA algorithm. Comparisons to a companion L74 reanalysis run using T425 outer and T159 inner loops revealed generally close agreement, though with more finescale structure in the T425 analysis at upper levels due to resolved large-scale gravity waves. The reanalyzed winds and temperatures from 0 to 50 km have been compared to other publically available operational analyses and reanalyses, revealing generally close agreement. We expect the NAVGEM 4DVAR DA algorithm to improve MLT tidal assimilation, relative to our earlier high-altitude 3DVAR-based system (Eckermann et al. 2009), by realistically (nonlinearly) propagating MLT increments from the various measurement times to the central analysis time using the TLM. To further improve our representation of MLT tidal effects, the 6-hourly reanalysis fields were filled at intervening hourly intervals using the 1–5-h forecasts from each DA cycle.

Figure 4 displays the NAVGEM reanalyzed horizontal wind vectors $V = (U, V)$ from $z = 0$ to 100 km upstream of Auckland Island at hourly intervals from 0400 to 1200 UTC on 14 July, corresponding to the periods just before, during, and just after RF23. Geometric heights $z$ were approximated using the analyzed geopotential height profiles. The profiles reveal strong...
westerlies from 0 to 70 km, with surface winds near 20 m s$^{-1}$ increasing to a southwesterly tropospheric jet peaking in excess of 50 m s$^{-1}$ near 7 km. At 0400–0500 UTC just prior to RF23, the stratopause jet peaked at over 100 m s$^{-1}$ near 70 km, with rapid diminution of wind speeds above this level due to gravity wave drag. The peak stratopause winds over the island progressively weakened over the next 6–8 h. A detailed analysis of the MLT wind fields is provided in section 4a.

c. Fourier gravity wave model

To model the observed MLT gravity waves, we use a Fourier method that has proven well suited to modeling stationary three-dimensional orographic gravity waves observed over islands: the method was used previously to model three-dimensional orographic wave fields observed in the stratosphere over South Georgia (Alexander et al. 2009; Jiang et al. 2014) and in the troposphere downstream of Jan Mayen (Eckermann et al. 2006a). Preliminary Fourier solutions also revealed striking similarities with the MLT gravity waves observed during RF23 (Fritts et al. 2016; Pautet et al. 2016), motivating a more detailed modeling effort here.

We consider linear nonhydrostatic stationary gravity waves in a height-dependent background atmosphere. For a wavenumber vector $(k, l, m)$ and background wind profile $V(z) = (U, V)$, the intrinsic frequency $\omega(k, l, z) = -kU(z) - lV(z)$. The nonhydrostatic irrotational gravity wave dispersion relation gives the corresponding vertical wavenumber

$$m(k, l, z) = sk_h \left( \frac{N^2}{\omega^2} - 1 \right)^{1/2},$$

where $N(z)$ is the background buoyancy frequency profile, $k_h^2 = k^2 + l^2$, and the sign parameter

$$s(k, l) = \text{sign}[\omega(k, l, 0)]$$

ensures upward group propagation ($c_{gs} = \partial \omega/\partial m > 0$) for each Fourier component.

We express wave field perturbations of a given atmospheric parameter $X$ as the inverse Fourier transform

$$X(x, y, z) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} S(k, l, z) \tilde{X}(k, l, z) e^{ikx + ily} dk dl.$$

A filter function $S(k, l, z)$ is inserted here for later use: until otherwise noted, $S(k, l, z) = 1$.

We evaluate (3) numerically by computing the Fourier solution $\tilde{X}(k, l, z)$ using ray methods, then performing a two-dimensional fast inverse Fourier transform at each height $z$. Although our full Fourier ray solution includes vertical reflections, resonant modes, and evanescent tunneling (e.g., Eckermann et al. 2006a; Broutman et al. 2006, 2009), here we omit those terms since our primary focus in this paper is propagating waves that reach the MLT. Thus at the surface we initialize only those Fourier components $\tilde{X}(k, l, 0)$ that are propagating (i.e., $m^2 > 0$ at $z = 0$). As $z$ increases, many of these initially propagating Fourier components become nonpropagating by reaching either a turning point $z_{tp}$, where $m^2(k, l, z_{tp}) \rightarrow 0 (\omega^2 \rightarrow N^2)$, or a critical level $z_{cl}$, where $m^2(k, l, z_{cl}) \rightarrow \infty (\omega^2 \rightarrow 0)$. Again we simply remove those Fourier components at $z \geq z_{tp}$ and $z \geq z_{cl}$ and ignore downward-propagating waves due to reflection at $z = z_{tp}$.

With these simplifications, the ray solution for vertical displacement $\tilde{\eta}$ is

$$\tilde{\eta}(k, l, z) = \left[ \frac{\rho_0}{\rho(z)} \right]^{1/2} \tilde{h} g \exp \left( i \int_0^z m dz \right),$$

where $\rho(z)$ is background air density, $\rho_0$ is the upstream value at $z = 0$,

$$g(k, l, z) = \left[ \frac{c_{gs}(k, l, 0)N^2(0)\omega(k, l, z)}{c_{gs}(k, l, z)N^2(z)\omega(k, l, 0)} \right]^{1/2}$$

is a wave-action conservation term [cf. Eq. (22) of Eckermann et al. (2015b)], and $h(k, l)$ is the Fourier transform of the terrain heights $h(x, y)$. The latter defines the equivalent Fourier form

$$\tilde{\eta}(k, l, 0) = \tilde{h}(k, l)$$

of the linear lower boundary condition $\eta(x, y, 0) = h(x, y)$. Solutions for other variables $X$ follow from polarization relations (see examples below).

Although the Fourier solution $\tilde{X}(k, l, z)$ represents only the steady-state ($t \rightarrow \infty$) wave field, we can use these solutions to approximate time-evolving wave fields typical of an initial value problem with “switch on” of surface forcing at $t = 0$. Since each harmonic $(k, l)$ has an associated group propagation time

$$t_{prop}(k, l, z) = \int_0^z c_{gs}^{-1}(k, l, z') dz'$$

to reach a given height $z$, then imposing a fixed cutoff time limit $t_{prop}(k, l, z_s) = t_s$ and solving numerically for $z$ in (7) yields a spectrum of corresponding cutoff altitudes $z_s(k, l, t_s)$. We can use this result to approximate the time-dependent wave field solution $X(x, y, z, t)$ at time $t = t_s$ using the inverse Fourier transform (3), as before, but applying a time-dependent filter function $S$ of the form
for any specified \( t = t_c \). Broutman et al. (2006) used a very similar method to model evolution of trapped orographic gravity waves, finding close agreement with direct numerical simulations.

For direct comparisons with the RF23 MLT observations, we convert wave field solutions for vertical displacement \( \eta(x, y, z, t) \) into perturbation temperatures using the adiabatic relation (e.g., Eckermann et al. 1998)

\[
T' = -\left( \frac{T N^2}{g} \right) \eta, \tag{9}
\]

where \( g \) is gravitational acceleration and \( T(z) \) is the background temperature profile. We also present solutions for wave steepness \( \eta_z \), defined as the vertical derivative of \( \eta \) and given by the Fourier integral

\[
\eta_z(x, y, z, t) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} i m S_{\text{prop}} \tilde{\eta} e^{i (k x + ly)} dk dl. \tag{10}
\]

We use these solutions to check for linearity (\( \eta_z < 1 \)) and to diagnose transition to wavebreaking amplitudes (\( \eta_z > 1 \)) versus height and time in the MLT. We do not attempt to parameterize the net nonlinear effects of wave breaking on modeled wave fields.

We derive vertical fluxes of horizontal momentum per unit mass using the Fourier solutions for wave-induced vector velocity derived from polarization relations: first, vertical velocity \( \tilde{w} = -i \omega \tilde{\eta} \), then zonal and meridional velocities \( \tilde{u} = -kmk_z^2 \tilde{w} \) and \( \tilde{v} = -lmk_z^2 \tilde{w} \), respectively. Given a Fourier array solution \( \tilde{X}(k, l, z) \) in any atmospheric parameter \( X \), we can derive a corresponding complex wave field solution

\[
\tilde{X}(x, y, z, t) = X_A(x, y, z, t) e^{i \psi_X(x, y, z, l)}, \tag{11}
\]

where \( X_A \) is local wave amplitude and \( \psi_X \) is local phase, using

\[
S_{\text{prop}}(k, l, t) = \begin{cases} 1 & z \leq z_c(k, l, t) \left| t_{\text{prop}}(k, l, z) \leq t \right. \\ 0 & z > z_c(k, l, t) \left| t_{\text{prop}}(k, l, z) > t \right. \end{cases}, \tag{8}
\]

\[
\tilde{m} w(x, y, z, t) = \frac{1}{4} \left( \tilde{u} \tilde{\omega}^* + \tilde{\omega} \tilde{u}^* \right) \quad \text{and} \quad \tilde{w} w(x, y, z, t) = \frac{1}{4} \left( \tilde{v} \tilde{\omega}^* + \tilde{\omega} \tilde{v}^* \right), \tag{13,14}
\]

respectively, where asterisks denote complex conjugation (Eckermann et al. 2010).

All Fourier solutions to follow were computed within an \((n_x, \Delta x, n_y, \Delta y, n_z, \Delta z)\) domain with \( n_x = n_y = 4096, n_z = 201, \Delta x = \Delta y = 1 \text{ km}, \) and \( \Delta z = 500 \text{ m}. \) A small amount of vertical diffusive damping \( (K_{zz} = 1 \text{ m}^2 \text{s}^{-1}) \) was used to avoid wraparound interference in the solutions (see appendix A of Eckermann et al. 2015b). In specifying terrain elevations \( h(x, y) \) for Auckland Island at the model’s lower boundary, we encountered considerable disagreement among some commonly used digital terrain databases, which our Fourier solutions proved sensitive to. The terrain adopted in this study, plotted in Fig. 1, was obtained at 200-m horizontal resolution from a global digital elevation model derived from measurements by the Advanced Spaceborne Thermal Emission and Reflection Radiometer (ASTER) on NASA’s Terra satellite then reinterpolated onto the model’s \( 1 \times 1 \text{ km}^2 \) horizontal grid.

3. Outbound NGV flight leg

a. Observations

Figure 5a shows the combined OH airglow brightness imagery observed by the AMTM in the MLT during the outbound transect of Auckland Island at ~0700 UTC. Large-amplitude wavelike banding of airglow brightness downstream of Auckland Island is evident. Rotational temperatures retrieved from the zenith camera observations in Fig. 5b reveal oscillations immediately downstream of Auckland Island with a horizontal wavelength \( \lambda_h \sim 40 \text{ km} \) and amplitude \( T_{\text{AG}} \sim 10 \text{ K} \) (Pautet et al. 2016). For a three-dimensional wave field \( T'(x, y, z) \), these two-dimensional airglow-derived temperature perturbations can be modeled as

\[
T'_{\text{AG}}(x, y) = \int I(z') T'(x, y, z') dz' / \int I(z') dz', \tag{15}
\]

where \( I(z) \) is the vertical variation in airglow radiance intensity through the MLT (Liu and Swenson 2003; Fritts et al. 2014). Using a standard Gaussian fit for \( I(z) \), during RF23 Pautet et al. (2016) inferred a peak at \( z \sim 83.5 \text{ km} \) and a full-width half maximum \( \sigma_{z,\text{AG}} \sim 7.5 \text{ km} \). If the wave has a constant amplitude \( T \) and vertical wave-number \( m \) within the airglow layer, then (15) implies
\[ S_{AG}(m) = \exp(-0.0902m^2\sigma_{AG}^2) \]  

Equations (16) and (17) show that in all likelihood the true wave temperature amplitude \( T \gg \tilde{T}_{AG} = 10 \, \text{K} \), with the actual value depending on the wave’s vertical wavelength \( \lambda_z = 2\pi/|m| \) within the layer.

The minimum vertical wavelength \( \lambda_{z\text{min}} \) detectable by the AMTM is \( \sim 10 \) km owing to the finite vertical width \( \sigma_{z\text{AG}} \) of the emission layer (Liu and Swenson 2003). Rearranging the dispersion relation (1) to yield \( \omega_{\text{min}}^2/N^2 = k_h^2/(k_h^2 + m_{\text{max}}^2) \), then substituting \( k_h = 2\pi/(40 \, \text{km}) \) and \( m_{\text{max}} = 2\pi/\lambda_{z\text{min}} \) yields a minimum intrinsic frequency for the wave of \( |\omega_{\text{min}}| \sim N/4 \), so that \( N/4 \leq |\omega| < N \). Thus the wave field here is non-hydrostatic, consistent with the observed downstream penetration of wave activity in Fig. 5b.

Figure 6 shows the Na mixing ratios that were observed in the MLT by the lidar as the NGV approached the island from the east. The observations end just after 0700 UTC as a result of a laser locking problem. Modeling by Xu et al. (2006) shows that, while chemistry becomes important on the bottom side of the Na layer, sodium remains an accurate dynamical tracer of gravity wave motion at these altitudes so long as mixing ratios rather than Na densities are analyzed. In particular, they showed that overturning of Na mixing ratio isopleths on the bottom side of the layer provides a reliable diagnostic of isentropic overturning due to gravity wave breaking, much like a dye tracer in a laboratory tank experiment.

The Na isopleths in Fig. 6 reveal a large-amplitude wave oscillation of \( \sim 40 \)-km horizontal wavelength on the bottom side of the Na layer, with overturning clearly evident at 78–80 km at 0700 UTC, or \( \sim 30–40 \) km downstream of the island. The AMTM temperature map in Fig. 5b also shows small-scale instability structures within a primary \( \lambda_z \sim 40 \) km wave, as typically occurs during incipient wave breaking (e.g., Andreassen et al. 1998), in a region located \( \sim 30–40 \) km downstream of the island. The steep forward face and rapid reduction in wave amplitude in Fig. 6 resemble tropospheric models of three-dimensional orographic gravity wave breaking in the approach to a critical layer (e.g., Miranda and Valente 1997), but in this case \( \sim 80 \) km above the parent terrain. Our modeling in later sections confirms this initial visual impression.

The Na isopleths permit estimates of wave amplitude \( \tilde{T} \) and vertical wavelength \( \lambda_z \). Tracking the lowest white isopleth along the x axis of Fig. 6 reveals an unperturbed equilibrium altitude of \( \sim 77–78 \) km and a peak upward displacement to \( \sim 81 \) km near 0700 UTC, implying a maximum vertical displacement amplitude \( \tilde{\eta} = 3–4 \) km. Inserting \( \tilde{T} = 210–230 \, \text{K} \) and \( N \sim 0.015–0.018 \, \text{s}^{-1} \) at 78 km, estimated from NAVGEM reanalysis profiles, into (9) yields \( \tilde{T} \sim 18–28 \, \text{K} \).

In linear theory, waves become statically unstable and break when the steepness amplitude

\[ \tilde{\eta}_z = |m|\tilde{\eta} \geq 1, \]

so that the maximum wave amplitude at this wave-breaking threshold is
Since instability in Fig. 6 occurs at $\eta \sim 3–4$ km, (19) suggests $\lambda_z \sim 18–25$ km. Substituting this into (17) yields $T \sim 15–20$ K using $T_{AG} \sim 10$ K. This amplitude is somewhat lower than the $T \sim 18–28$ K inferred from the Na isopleth displacement, but recall (17) is derived assuming constant $T$ and $m$ through the airglow layer, whereas observations in Fig. 6 (as well as later model results) indicate strong vertical gradients in $T$ and $\lambda_z$ as both reduce with height toward a possible overlying critical level.

b. Fourier solutions

To assess linear limits required for application of our Fourier solutions to RF23 wave fields, we first tested the linear lower boundary condition [(6)] by computing the surface Froude number

$$\text{Fr} = \frac{[U^2(0) + V^2(0)]^{1/2}}{N(0) h_{\text{peak}}}.$$  \hspace{1cm} (20)

Surface forcing is linear if Fr $\geq$ Fr$_c$ $\sim$ 1 (where Fr$_c$ is the critical Froude number of $\sim$1) (Smith 1989). Evaluating (20) using analysis fields and RF23 dropsonde data upstream of Auckland Island using $h_{\text{peak}} = 649$ m (see Fig. 1) yielded Fr $\sim 4 \pm 1$ throughout the RF23 period. Modified formulas that account for low-level vertical gradients (Reinecke and Durran 2008) produced similar findings.

Since the surface forcing environment is linear, we computed Fourier wave field solutions from $z = 0$ to 100 km for a range of propagation cutoff times $t_c$ using upstream wind, height, stability, and density profiles from the NAVGEM reanalysis from 0400 to 1200 UTC (see Fig. 4). At every height $z$, we computed the maximum value of the wave field steepness solutions $\eta(z, t_c)$ from $z = 0$ to 65 km, implying linear (nonbreaking) gravity wave dynamics throughout this deep atmospheric layer. At $z \sim 67$ km, $\eta_{z_{\text{max}}}^\text{max}(z, t_c)$ slowly and asymptotically

$$\eta_{z_{\text{max}}}^\text{max}(z, t_c) = \max[\eta(z, x, y, z, t_c)].$$  \hspace{1cm} (21)

of the wave field steepness solutions $\eta_{z_{\text{max}}}^\text{max}(z, x, y, z, t_c)$, as well as the locations $[x_{\text{max}}(z), y_{\text{max}}(z)]$ of those maxima.

Figure 7a shows the resulting $\eta_{z_{\text{max}}}^\text{max}(z, t_c)$ profiles for $t_c$ values ranging from 1 h through to $\infty$ (steady state) based on the NAVGEM upstream profiles at 0600 UTC. The $\eta_{z_{\text{max}}}^\text{max}(z, t_c)$ profiles reveal significant wave activity in the MLT within $t_c = 1.5$ h of forcing at the surface, and by $t_c = 4$ h the profile is already close to the steady-state limit. This is consistent with the high intrinsic frequencies (and hence fast vertical group velocities) inferred earlier for these waves.

Figure 7a also reveals $\eta_{z_{\text{max}}}^\text{max}(z, t_c) \ll 1$ from $z = 0$ to 65 km, implying linear (nonbreaking) gravity wave dynamics throughout this deep atmospheric layer. At $z \sim 67$ km, $\eta_{z_{\text{max}}}^\text{max}(z, t_c)$ slowly and asymptotically
approaches unity as \( t_c \to \infty \), implying a weak slow-developing wave field instability at this altitude. However, as shown in Fig. 8b [see also Broutman et al. (2016), manuscript submitted to J. Geophys. Res.], this occurs in a region of the wave field many hundreds of kilometers downstream and to the south of the island, far away from (and hence largely irrelevant to) the wave fields observed in Fig. 5. Thus, as highlighted in Fig. 7b, the Fourier solutions predict onset of intense wave field instabilities \( \eta_{\text{max}}(z, t_c) \) at a breaking height \( z_b \sim 77.5\text{–}78.5 \text{ km} \)—exactly the altitude where wave breaking is observed to occur in Fig. 6. Moreover, the corresponding locations \( (x_{\text{max}}, y_{\text{max}}) \) of maximum wave field instability at these altitudes, plotted in Fig. 8a, all cluster within a tightly confined region \( \sim 40 \text{ km} \) east of the island, again in agreement with the observed locations of both overturning Na isopleths in Fig. 6 and intrawavelength instability structure in Fig. 5b. Thus, consistent with our diagnostics confirming the validity of linear solutions up to \( z \sim 78 \text{ km} \), those solutions reveal impressive agreement with the salient observed characteristics of the MLT wave fields.

Figure 9a shows the vertical evolution of propagating harmonics in these solutions throughout the stratosphere and MLT. The total number of unique Fourier harmonics is \( n_h n_s / 2 \approx 8.4 \times 10^6 \), yet only \( \sim 10^5 \) (or \( \sim 1\% \)) of those propagate out of the troposphere, and these are then progressively eroded with height to \( \sim 1\text{–}2 \times 10^4 \) by \( z \sim 70\text{–}80 \text{ km} \). Some stratospheric filtering occurs as northeastern wind vectors rotate with height to southeastern in Fig. 4, removing horizontal wave-numbers \( k_h = (k, l) \) orthogonal to \( \mathbf{V} \) via directional critical-level filtering \( \omega = -k_h \cdot \mathbf{V} / \alpha \), as shown in Fig. 9b. Other waves aligned nearly parallel to the intensifying stratospheric winds progressively increase in intrinsic frequency with height, eventually being removed at turning points where \( \omega^2 \sim N^2 \) (see Fig. 9c). Figure 9d shows that many waves removed by the cutoff criterion \( t > t_c \) for small \( t_c \) are eventually removed by critical levels or turning points in the steady-state limit.
Figure 9 also reveals very little wave field filtering above the stratopause jet from $z = 55$ to $80$ km, but that all remaining waves are removed by critical levels within a narrow layer just above $80$ km [accounting also for cessation of $h(z, t_c)$ profiles above $83$ km in Fig. 7a]. This occurs as a result of a rapid reduction in 0600 UTC wind speeds with height in Fig. 4 to a near-zero-wind line at $z = 83$ km.

Figure 10 depicts the vertical evolution of these wave field filtering characteristics in spectral space. The lower boundary condition [(6)] of the Fourier $\tilde{h}$ solution, $\tilde{h}(k, l)$, is shown in Fig. 10a within a $(k, l)$ spectral range of $2\pi (10$ km)$^{-1}$ where most of the Auckland Island terrain variance is concentrated. Smaller panels to the right depict propagating (red) and nonpropagating (other colors) regions of this spectral space as a function of height within the Fourier solutions at $t_c = 4$ h (top row) and $t_c = \infty$ (bottom row). Comparing the two, we again see that regions where waves are removed by $t > t_c$ cutoffs (black areas) in most cases transition to critical levels or turning points in the steady-state limit.

Below $10$ km there are strong northeastward winds and vertical shear (see Fig. 4). The Fourier solutions at $z = 10$ km in Figs. 10b and 10g reveal that $k$ vectors orthogonal to those wind vectors have been removed by directional critical-level filtering (green regions), while shorter $\lambda_h$ waves, particularly those with $k$ nearly parallel to the tropospheric wind vectors, have been removed at turning points (blue regions). This leaves a bowtie-shaped wedge of propagating harmonics (red region) at $z = 10$ km that progressively shrinks from $z = 10$ to $60$ km via additional filtering by directional critical levels and turning points, as shown in Figs. 9b and 9c. At $z = 60$ km the remaining propagating waves in Figs. 10c and 10h all have $\lambda_h \geq 30$ km, with the nearly symmetric diamond shape of the propagating (red) region implying a three-dimensional ship-wave response.

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**FIG. 9.** Spectral filtering of 0600 UTC Fourier solutions vs height for a range of cutoff times $t_c$ depicted by colored curves (color key at far right): (a) number of remaining propagating harmonics $(k, l)$ at each height, and number of harmonics removed at each height by (b) critical levels, (c) turning points, and (d) propagation times $t_{prop} > t_c$.

**FIG. 10.** (a) Normalized spectral terrain amplitude $|\tilde{h}(k, l)|/|\tilde{h}(0, 0)|$ for $k$ and $l$ in the range $2\pi (10$ km)$^{-1}$. (b)–(k) Same region of spectral space, with colors representing propagation status by wavenumber of the Fourier solutions at 0600 UTC at altitudes $z$ of (left)–(right) 10, 60, 82, 83, and 84 km. Red regions contain propagating waves. In all other regions, waves have been removed owing to critical levels (green), turning points (blue), or exceeding the cutoff time $t_c$ for wave propagation (black). Results for (top) $t_c = 4$ h solutions and (bottom) $t_c = \infty$ (steady state) solutions.
As shown in Fig. 9, little filtering occurs from ~60 to 80 km, such that at $z = 60$ km (Figs. 10d and 10i) the regimes are largely unchanged from those at $z = 60$ km. By contrast, above 82 km the entire remaining propagation space is removed via critical levels within a 2-km vertical layer, so that at $z = 82$ km (Figs. 10f and 10k) no single propagating harmonic remains. Again this result is in agreement with the observations in Fig. 6, which show large amplitudes at $z \sim 78$ km then rapid amplitude attenuation with height such that little or no wave activity is observed above 84 km. We also noted earlier that the steep forward face of the near-overturning wave crest at $z \sim 78$ and ~40 km east of the island terrain in Fig. 6 resembled tropospheric models of three-dimensional orographic gravity wave breaking below a zero-wind line (e.g., Miranda and Valente 1997). Our work confirms that this is in fact the case here: that is, that the outbound RF23 transect observed orographic gravity wave breaking just below a critical level in the MLT.

Next we conduct a detailed comparison of our Fourier solutions to the waves imaged in the airglow. Figure 11a shows the $T'(x, y, z, t_c)$ wave field after $t_c = 4$ h at $z = 78$ km, where solutions are still nearly linear and observed wave amplitudes in Fig. 6 peak. These solutions used the 0700 UTC NAVGEM profiles, which are closest in time to the outbound airglow measurements. Aspects of the modeled wave field show impressive quantitative agreement with the airglow imagery in Fig. 5, including (i) $\lambda_h \sim 40$ km waves downstream of the island, (ii) a cold phase line of longer $\lambda_h$ immediately above the island, and (iii) a diverging wake of waves to the northeast. The nonhydrostatic dispersion relation (1) plays a pivotal role. The corresponding hydrostatic solution (not shown) has no $\lambda_h \sim 40$ km downstream waves: wave fields are concentrated near the island and in diverging wakes, have shorter $\lambda_h$ and larger amplitudes, and become unstable at much lower altitudes. Thus, nonhydrostatic downstream group propagation and wave filtering at turning points exert first-order impacts on the observed MLT wave fields.

Aspects of Fig. 11a that differ from observed wave fields in Fig. 5a are (i) a diverging wake to the southeast that is not observed and (ii) peak amplitudes significantly larger than those in Fig. 5. One key aspect ignored to this point is the vertical averaging of the observed wave response via (15) due to the finite vertical width of the airglow emission profile $I(z)$. Since $I(z)$ extends to altitudes above 90 km, accurate numerical evaluation of (15) requires an accurate model of $T_0(x, y, z, t_c)$ throughout the MLT, whereas our linear solutions begin to break down near 78 km (see Fig. 7). To gauge the effects more simply, we instead apply the spectral filter $S_{AG}(m)$ in (17) as an additional filter function multiplier $S$ when inverting the Fourier solution $\tilde{T}(k, l, z)$ using (3) at $z = 78$ km for $t_c = 4$ h. The result is plotted in Fig. 11b. The two major discrepancies with the observations are largely eliminated: the southeastward diverging wake is now heavily suppressed in amplitude (since $\lambda_z$ values

![Fig. 11](image_url)
here are short), and peak wave field amplitudes are now near the $T_{AG} = 10 \text{ K}$ observed in Fig. 5b.

The vertical fluxes of zonal and meridional momentum per unit mass, $\overline{uw}(x, y, z, t_c)$ and $\overline{vw}(x, y, z, t_c)$ at $z = 78 \text{ km}$ after $t_c = 4 \text{ h}$, are shown in Fig. 12 based on the 0600 UTC NAVGEM profiles. Gray lines show regions where $\eta_z > 1$ in the corresponding $\eta_z(x, y, z, t_c)$ wave field solution. Figure 12a exhibits strong concentration of zonal momentum flux $\sim 40 \text{ km}$ downstream of the mountain, with westward fluxes in the 300–600 m$^2$ s$^{-2}$ range. These values are similar to (but somewhat larger than) the $\sim 320 \text{ m}^2 \text{ s}^{-2}$ estimated by Pautet et al. (2016) using observed properties in Figs. 5b and 6 and linear gravity wave theory. Peak meridional fluxes in Fig. 12b are nearly an order of magnitude smaller and of roughly equal and opposite value in the northern and southern diverging wake regions, such that the area-averaged value is even smaller.

4. Temporal evolution of MLT winds and wave fields

a. Semidiurnal tidal effects in the MLT

Figure 13 shows hourly time series of analyzed wind vectors $\mathbf{V}(z)$ and zonal winds $U(z)$ in the MLT from $\sim 70–100 \text{ km}$ over Auckland Island during RF23. They reveal a zero-wind line at $z \sim 86–87 \text{ km}$ at 0400–0500 UTC that systematically descends over the following hours to $\sim 80 \text{ km}$ by 1000 UTC. Descending structure is also evident in buoyancy frequencies $N(z)$ (not shown).

This descending wind and temperature structure is produced by a large-amplitude migrating semidiurnal tide in the MLT reanalysis. A Hovmöller diagram of analyzed zonal and meridional winds at 0.015 hPa ($z \sim 75 \text{ km}$) and 50.4$^\circ$S for a 5-day period centered on RF23 is shown in Fig. 14. RF23 is depicted as the segment of the white horizontal line at 166$^\circ$E (longitude of Auckland Island) between the two vertical black lines depicting RF23 takeoff and landing times. The plot reveals a wavenumber-2 (migrating) semidiurnal tide in the MLT over Auckland Island with an amplitude in both the zonal and meridional components of $\sim 20–30 \text{ m s}^{-1}$. As the tide propagates rapidly to the west, peak eastward tidal winds just before takeoff at $\sim 0400–0500 \text{ UTC}$ (see Fig. 4) are replaced by peak westward tidal winds nearer the end of the flight at $\sim 1000–1100 \text{ UTC}$. Despite these tidal modulations, winds over Auckland Island remain eastward at this altitude throughout the flight.

Tidal phase varies in the vertical, as evident in Fig. 15, which shows the corresponding Hovmöller diagram at 0.0064 hPa ($z \sim 81 \text{ km}$). Tidal wind amplitudes are even larger at this altitude and tidal phase noticeably leads in time relative to tidal phase at 75 km. This phase...
difference manifests locally above Auckland Island as phase descent of wind regimes with time as seen in Fig. 13. At takeoff zonal winds at 81 km in Fig. 15a are near the tidal node but are soon replaced by the tidal zonal-wind trough, causing $U(z)$ to transition from weakly eastward at the start of the flight at $\sim$0600 UTC, to weakly westward by midflight at $\sim$0900 UTC, then back to weakly eastward by landing time at $\sim$1200 UTC.

Although there are no local MLT wind observations to validate these analyzed tides over Auckland Island, MLT winds have been measured at these latitudes using radars stationed in and around the southern tip of South America. Those measurements during previous austral winters revealed large-amplitude semidiurnal tidal winds with amplitudes of up to 80 m s$^{-1}$ as well as weak diurnal tidal wind amplitudes (Fritts et al. 2010, 2012), consistent with the NAVGEM analyzed MLT wind fields over Auckland Island during RF23.

To summarize, a large-amplitude semidiurnal tide is responsible for descent of a zero-zonal-wind line over Auckland Island during RF23, from $\sim$87 km just before takeoff to $\sim$80 km just before landing, as shown in Fig. 13b.

5. Inbound NGV flight legs

a. Observations

Immediately after the first transect of Auckland Island at 0700 UTC, the NGV flew south to Macquarie Island, returning to Auckland Island at a lower flight altitude $\sim$3 h later. As shown in Fig. 3, three additional transects of Auckland Island were performed during these inbound flight legs in an ascending “stair step” pattern.

Figure 18 shows the AMTM airglow and temperature imagery acquired during each of these transects. Much of the MLT wave structure seen at 0700 UTC in Fig. 5 is still evident at 1000 UTC (Figs. 18a and 18d), but wave amplitudes are weaker. For the subsequent overpass at $\sim$1030 UTC (Figs. 18b and 18e), wave structure immediately downstream of the island has now disappeared, owing presumably to further breakdown of waves into turbulence. However waves in the diverging wake to the northeast are still present in the wing camera imagery. By $\sim$1115 UTC (Fig. 18d) nearly all wave structure has
disappeared, with northeastern wave structure weakly visible but showing evidence of breakdown into smaller-scale instability structures. Correlative Na lidar observations were not available for these inbound transects.

**b. Fourier solutions**

The 1000 UTC wave field in Fig. 17d predicts intense wave breaking throughout the eastward and southeastward regions of the wave field, consistent with wave breakdown observed in these regions in Fig. 18. The surviving portion of the wave field is confined to the northeast in the wing camera imagery of Fig. 18b. Likewise, the $T'(x, y, z, t_c)$ solution in Fig. 17d reveals wave structure to the northeast of similar appearance to that observed and no regions of diagnosed wave breaking ($\eta_z \approx 1$) either.

As for the outbound data, a more quantitative comparison with the inbound AMTM data requires vertical integration of the wave response through the airglow emission layer $I(z)$. Figure 19b shows the 1000 UTC solution after applying the spectral airglow filter $\delta_{AG}(m)$ as before, revealing wave structure concentrated in the northeast quadrant of very similar appearance to that observed. Inferred amplitudes cannot be compared since temperatures are not retrieved from the wing camera observations.

To understand how filtering controls these changes in the MLT wave field structure at these later observation times, Fig. 20 shows propagating and filtered waves within the $(k, l)$ spectral solution space at various heights in the MLT for solutions using the 0900 UTC analysis profiles. The region of propagating waves at $z = 76$ km is similar to that seen at ~0600 UTC (e.g., Fig. 10c). However, strong filtering of this propagating region occurs at lower MLT heights than at 0600 UTC owing to descending westward shear and zero-zonal-wind lines due to the semidiurnal tide (Fig. 13b), which in this case do not completely remove all propagating waves. Instead a narrow wedge of propagating waves with $k$ vectors aligned northwest/southeast remains at $z = 84$ km and above (Figs. 20c and 20f). This transmission is facilitated by stronger southward meridional MLT winds at this time, which prevent complete stagnation of the vector wind. It is these surviving waves that produce the wave activity to the northeast both observed in Figs. 18a and 18b and reproduced in the Fourier solutions in Fig. 19b.

### 6. Discussion and conclusions

MLT airglow imagery acquired during NGV transects of Auckland Island during DEEPWAVE RF23 (Figs. 5
and 18) appear to reveal a three-dimensional orographic gravity wave response to flow over island terrain at unexpectedly high altitudes (Fritts et al. 2016; Pautet et al. 2016). The present work has thoroughly investigated the dynamics of this event from the ground to the MLT. We showed that the upstream flow impinging on Auckland Island terrain satisfied conditions for linear surface forcing (Fr ≫ 1). Our ensuing linear Fourier solutions, derived using analyzed upstream wind, temperature, stability, and density profiles and realistic island terrain, remained stable (i.e., linear: $h_z \sim 1$) at all points in the wave field up to $z \sim 70–80$ km (see Fig. 16). Beyond merely establishing the validity of linear approximations over this deep vertical layer of the atmosphere, our linear solutions accurately reproduced (and provided explanations for) all the salient observed features of these MLT wave fields—wave-breaking heights, time evolution of three-dimensional wave field structure, downstream horizontal wavelengths, airglow temperature amplitudes, lidar-derived vertical displacement amplitudes, and inferred momentum fluxes per unit mass. In summary, we can now confidently classify this observed event as a remarkable validated case study of three-dimensional nonhydrostatic orographic gravity waves from an island evolving linearly from the surface to very high altitudes right up to the point of observed incipient wave breaking in the MLT.

Why do we find linear orographic gravity waves up to the MLT to be remarkable? Linear hydrostatic wave equations are used routinely to parameterize subgrid-scale orographic gravity wave breaking and drag within NWP and climate models [see, e.g., section 2a of Eckermann et al. (2015b)]. In the absence of shear, those equations yield wave amplitude increases with height of

![Diagram of NAVGEM Reanalysis Wind on 6.4x10^{-3} hPa (~81 km)](image)

**FIG. 15.** As in Fig. 14, but at 0.0064 hPa ($z \sim 81$ km) and 50.4°S. Note also change in dynamical range of color bars relative to Fig. 14.

![Graph showing breaking heights $z_b$ as a function of time (UTC hours on 14 Jul 2014)](image)

**FIG. 16.** Breaking heights $z_b$ as a function of time (UTC hours on 14 Jul 2014) derived by identifying the height $z$ where Fourier wave field solutions $\eta(x, y, z, t)$ over Auckland Island first exceed unity. Gray curve shows mean $z_b$, computed over all $t$ values at a given universal time.
$d(z) = \left[ \rho / \rho(z) \right]^{1/2}$. At $z \approx 78$ km, where our observations and modeling indicate RF23 wave fields first break (Fig. 6), $d(z) = 300$, so that, for a representative surface wave amplitude of $\eta_0 = 500$ m (Fig. 1), an amplitude $d(z)\eta_0 \approx 150$ km at $z = 78$ km would be predicted. This unrealistically large value implies $z_b \ll 78$ km, and indeed parameterized midlatitude orographic gravity wave breaking typically occurs first in the troposphere or lower stratosphere (e.g., McFarlane 1987).

So what processes operate here to keep wave fields linear up to such unexpectedly high altitudes? At least three stabilizing dynamical effects are important in our Fourier solutions. First, refraction by background winds affects wave amplitudes via the spectral $g(k, l, z)$ term [Eq. (5)], which, in the hydrostatic limit for purely zonal winds $U(z)$ and constant $N$, reduces to a height profile $g(z) = [U(0)/U(z)]^{1/2}$. The change in $U(z)$ from $\sim -20$ m s$^{-1}$ at the surface to $\sim -100$ m s$^{-1}$ at $z \approx 55$ km (see Fig. 4) reduces amplitudes by a factor $g(z) \approx 0.45$ over this altitude range. Second, the horizontal spreading of wave fields into progressively larger horizontal areas away from Auckland Island (e.g., Figs. 17a–17d) leads to corresponding decreases in local wave field amplitudes to conserve wave action (Eckermann et al. 2015b,a; Broutman et al. 2016, manuscript submitted to J. Geophys. Res.). Finally, the progressive erosion with height of amplitude contributions at specific harmonics $(k, l)$ via directional critical levels and turning points (e.g., Figs. 9 and 10) further reduces the vertical growth of local wave field amplitudes with height. These three processes acting together allow linear conditions to persist deep into the mesosphere. Nonhydrostatic effects, specifically downstream group propagation accompanied by enhanced horizontal geometrical spreading (Broutman et al. 2016, manuscript submitted to J. Geophys. Res.) and turning-point filtering have large impacts on our linear solutions, but are typically not included in orographic gravity wave drag parameterizations in NWP and climate models.

A potential weakness of our Fourier solutions is in our treatment of time dependence, which involves imposition of propagation cutoffs $t_c$ to the steady-state Fourier solutions. Solutions initialized at 0400 UTC, for example, can be visualized as waves emanating from the bottom-left corner of Fig. 4 as many oblique ray paths for each $(k, l)$, each meandering upward and to the right through this $(z, t)$ space. Although our spectral $t_c$ cutoff incorporates this group propagation,
our solutions do not account for the time-varying wind fields that each wave group encounters in Fig. 4 as it propagates forward in time. Instead, our solutions assume the wind profile at initialization \((t = 0)\) remains the same.

The effects of time-dependent winds on linear wave fields can be gauged using a ray equation that, after simplifying here to zonally aligned winds and waves and ignoring horizontal gradients so that \(k\) remains constant, can be expressed as

![Diagram](image)

Fig. 18. Presentation of AMTM (a)–(c) airglow and (d)–(f) temperature imagery as in Fig. 5, but showing results for Auckland Island transects from (left) 0945–1010, (center) 1015–1050, and (right) 1105–1125 UTC.

![Diagram](image)

Fig. 19. Temperature perturbation presentation as in Fig. 11, but for the 1000 UTC Fourier solutions \((t_c = 4\) h) at \(z = 78\) km.
Here \( c \) is zonal phase speed along a ray path, which we assume to be initially stationary \((c_0 = 0)\). In the MLT, semidiurnal wind amplitudes range from \( \sim 20 \) to \( 50 \) m \( s^{-1} \) (Figs. 14 and 15) so \( |\partial U/\partial t| \) can be large locally, potentially accelerating \( c \) and affecting when and where waves break (e.g., Eckermann and Marks 1996). The issue here comes down to the wave propagation time \( \Delta t \) in (22), which we estimate as the time required to propagate from \( z = 60 \) km, where tidal amplitudes first become significant, to \( z = 78 \) km where waves initially break: we evaluate \( \Delta t \) using (7) with the above height integration limits. Representative calculations for the \( \lambda_h \sim 40 \) km waves to the east of Auckland Island using reanalysis winds yield \( \Delta t \sim 10\text{–}20 \) min. So, notwithstanding large \( |\partial U/\partial t| \), waves propagate so rapidly through MLT tidal winds that little change in \( c \) occurs via (22), explaining why our Fourier solutions that omit such effects compare so closely to observations. That said, waves in the southeastern diverging wake propagate more slowly in the vertical and might be affected more by tidal accelerations, including tidal \( \partial V/\partial t \) accelerations.

We conclude by estimating effects on the MLT of the local deposition of these wave field momentum and energy fluxes through wave breaking. Irreversible flow accelerations due to wave momentum flux deposition are given by

\[
\frac{\partial U}{\partial t} = a(z) = -\frac{1}{\rho} \frac{\partial}{\partial z} (\rho \bar{w} \bar{w}).
\]  

Since our Fourier solutions do not model wave breaking, here we perform a separate calculation. We specify a stationary \((c = 0)\) wave field at \( z = 78 \) km (just prior to wave breaking) of \( \lambda_h = 40 \) km and \( \bar{w} \bar{w} = -300 \) m \( s^{-2} \), values derived directly from observations by Pautet et al. (2016), who noted that these fluxes are \( \sim 10\text{–}100 \) times larger than typical \( \bar{w} \bar{w} \) values observed in the MLT. Our model results in Figs. 12a and 17h–17k reproduce similar or larger local flux values. We use linear saturation theory (e.g., Holton 1983; McFarlane 1987) to model breaking and deposition of these fluxes through analyzed 0600 UTC zonal winds, which, as shown in Fig. 21a,
decrease with height to a critical level at \( z \approx 83 \) km. These decreasing winds cause all the wave momentum flux to be deposited within a narrow vertical layer beneath this zero-wind line (see Fig. 21b), largely consistent with the rapid attenuation with height of wave amplitudes above 78 km evident in the lidar data in Fig. 6. This results in huge flow accelerations in Fig. 21c, with a density-weighted layer-averaged value from \( z \approx 79 \) to 83 km of around \( 2350 \) m s\(^{-1}\) h\(^{-1}\). Such intense localized forcing likely triggers an intense surrounding ageostrophic circulation and secondary wave generation. This would imply a remarkably rapid evolution from linear wave field dynamics to strongly nonlinear local dynamics due to rapid and intense wave breaking beneath a critical level.

Corresponding dynamical heating rates are given by

\[
\frac{\partial T}{\partial t} = -\frac{U}{C_p} a + \frac{T}{(1 + Pr)} \frac{\partial}{\partial p}(U \rho a),
\]

(24)

where \( p \) is pressure and \( C_p \) is mass specific heat at constant pressure (Medvedev and Klaassen 2003). The quantity \( H_2(z) \) is a differential heating term due to the wave’s vertical heat-flux divergence. We neglect it here since its form depends sensitively on assumption about wavebreaking dynamics [e.g., effective Prandtl numbers (Pr) Akmaev 2007] and on the vertical curvature of the momentum flux profile, which the simple calculations here do not accurately constrain. Furthermore, the density-weighted layer average of \( H_2(z) \) vanishes (Medvedev and Klaassen 2003).

The heating rates \( H_1(z) \) due to frictional dissipation of total wave energy density, plotted in Fig. 21d, are large, with a corresponding density-weighted layer average from \( z \approx 79 \) to 83 km of \( \sim 8 \) K h\(^{-1}\). Figures 18d–18f provide possible observational support for these large inferred heating rates, with mean temperatures averaged through the airglow layer increasing by \( \sim 10–15 \) K over the \( \sim 1–1.5 \) h of measurements during three separate NGV transects, when these waves are observed to be breaking vigorously and depositing their energy and momentum locally in the MLT.

Since such large upper-level MLT drag will also significantly affect deep underlying atmospheric layers through downward control [see, e.g., Garcia and Boville (1994)], the above calculations provide direct observational support for recent conjectures that orographic gravity wave drag due to subantarctic islands contributes significantly to the overall momentum and energy budget controlling the middle-atmospheric circulation and climate during austral winter (Alexander et al. 2009; McLandress et al. 2012).

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