



Space Dynamics

LABORATORY

Utah State University Research Foundation

Absolute Radiance Recalibration of FIRST

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Outline

- ▶ FIRST
- ▶ FIRST on-board calibration
- ▶ Ground calibration equipment
- ▶ Absolute radiance response calibration
 - Data collection
 - Data processing highlights
 - Non-linearity
 - Results

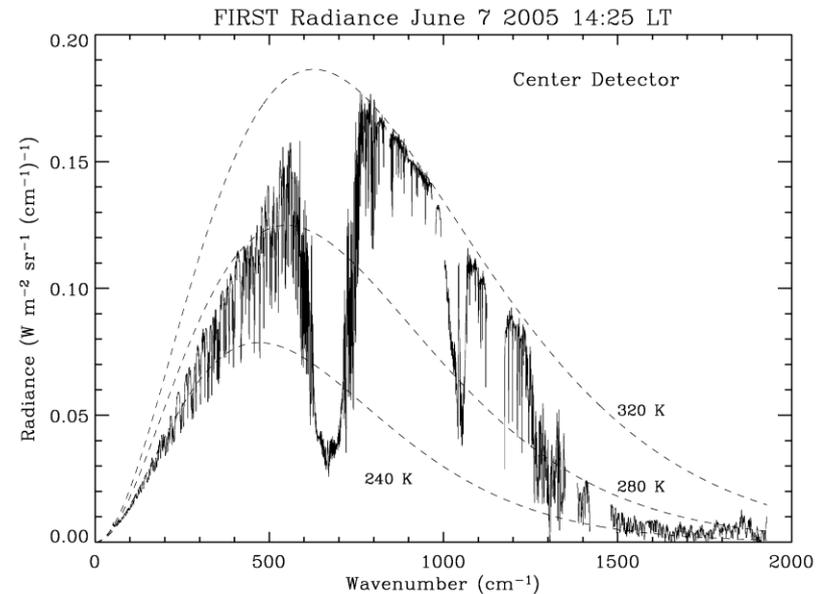


FIRST (Far-IR Spectroscopy of the Troposphere)

- ▶ FIRST is an instrument that measures the Earth's atmospheric radiance in the FAR-IR
- ▶ Has been successfully used since 2005 from high altitude balloons and from the ground
- ▶ FIRST developed under an Instrument Incubator Program
 - Goal of developing technology needed to attain daily global coverage, from low-earth orbit, of the Earth's far-infrared spectrum
 - Technology to be demonstrated with a prototype instrument in a space like environment

Far-IR ($>15 \mu\text{m}$, $<667 \text{ cm}^{-1}$)

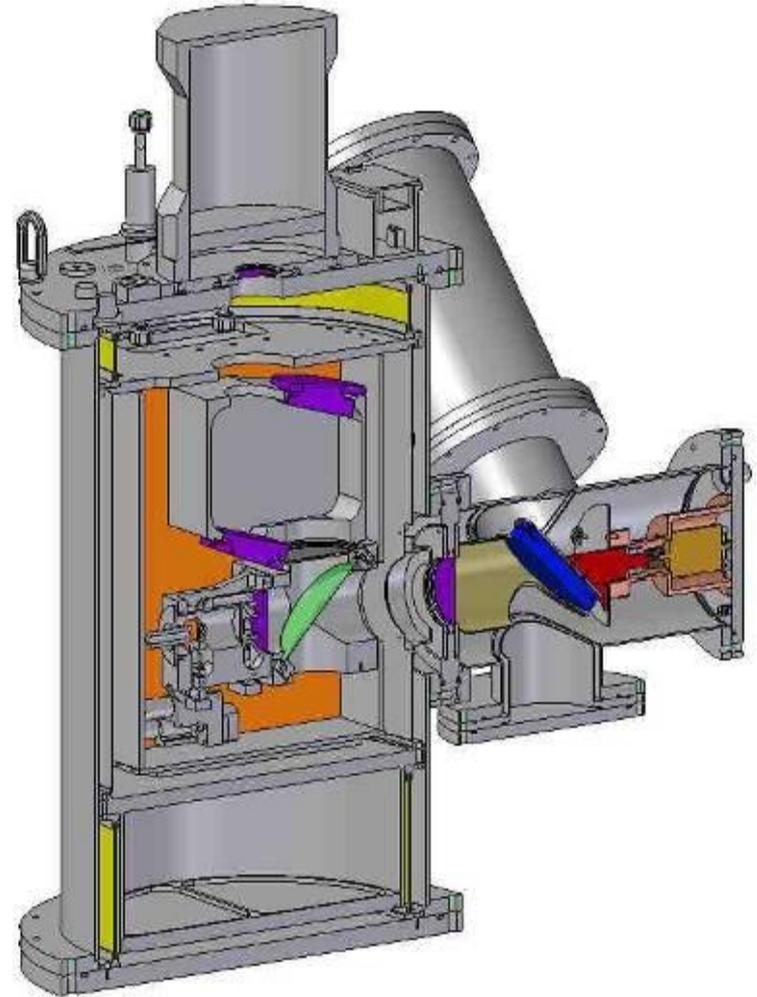
- ▶ Contains half of Earth's outgoing long-wave radiation
- ▶ Is not well observed spectrally



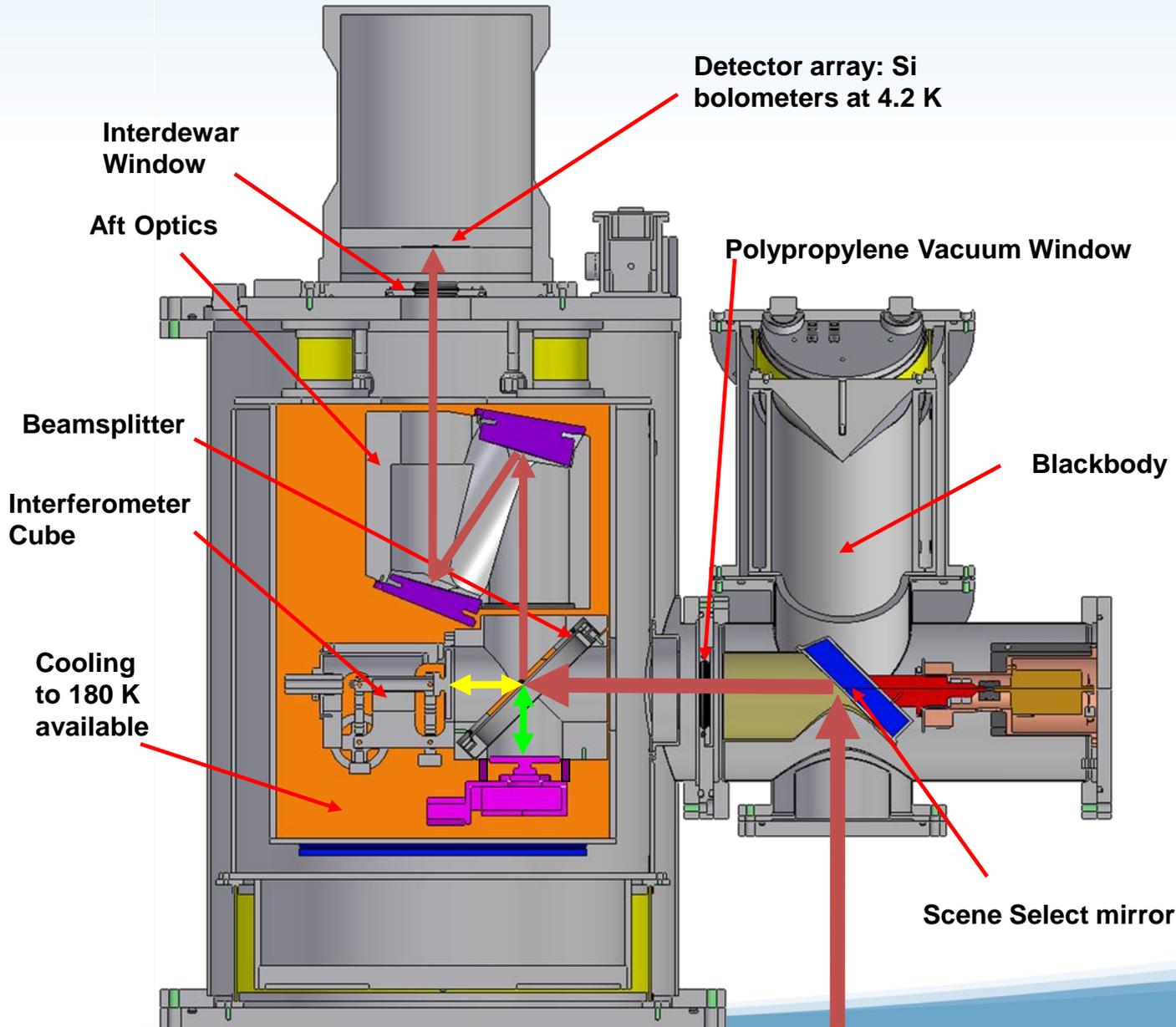
FIRST spectrum from high altitude balloon

FIRST specs

- ▶ Fourier Transform spectrometer
 - Michelson interferometer
 - Coverage
 - Goal: 100 to 1000 cm^{-1} (100 to 10 μm)
 - Actual: 50 to 2200 cm^{-1} with breaks
- ▶ Spectral Resolution: 0.643 cm^{-1} (unapodized)
- ▶ NE Δ T goals
 - 0.2 K (k=1) 170 to 1000 cm^{-1} @ 230 K
 - 0.5 K (k=1) 100 to 170 cm^{-1} @ 230 K
- ▶ Accuracy goal: equal to NE Δ T
- ▶ On-board blackbodies or blackbody and space view for calibration
- ▶ 7 cm aperture
- ▶ Ability to have 4.4° FOV (~100 km from orbit)
 - 10 detectors in sparsely populated array
- ▶ Liquid He cooled Si bolometers
- ▶ 0.41° IFOV (~10 km from orbit)
- ▶ 24576 points per interferogram
- ▶ 11.5 sec collection time



FIRST



- Simple optics
- 3 sections
- 3 port SSA scene select assembly (SSA)
- SSA rotates
- COTS electronics

FIRST on-board calibration

- ▶ FIRST views both on-board calibration sources during data collection
- ▶ Calibration equation

$$R_{Target} = \frac{S_{Target} - S_{ABB}}{\mathfrak{R}} + P_{\leftarrow ABB} \quad \mathfrak{R} = \frac{S_{WBB} - S_{ABB}}{P_{\leftarrow WBB} - P_{\leftarrow ABB}}$$

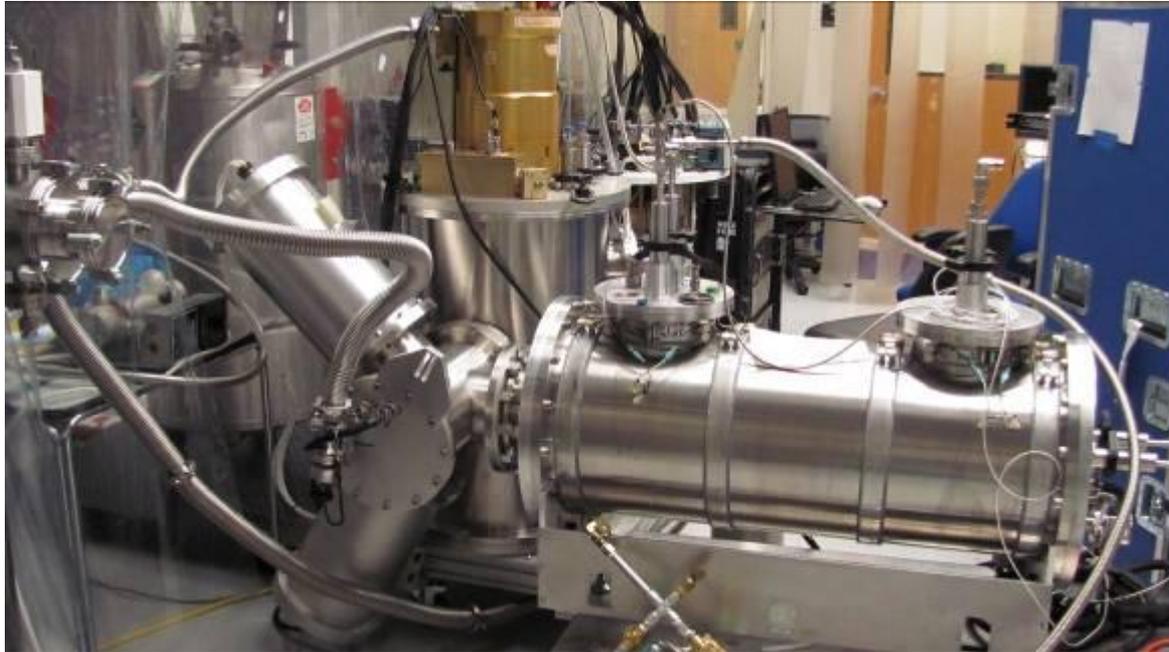
S_{Target} , S_{WBB} , S_{ABB} : Observed signal from target, warm and ambient blackbodies

T_{WBB} , T_{ABB} : Temperature of warm and ambient blackbodies

- ▶ Used to calculate target radiance
- ▶ Warm, Ambient blackbodies used for ground data
- ▶ Warm blackbody, space view used for balloon data
- ▶ Forward and backward scans are calibrated independently

FIRST ground calibration

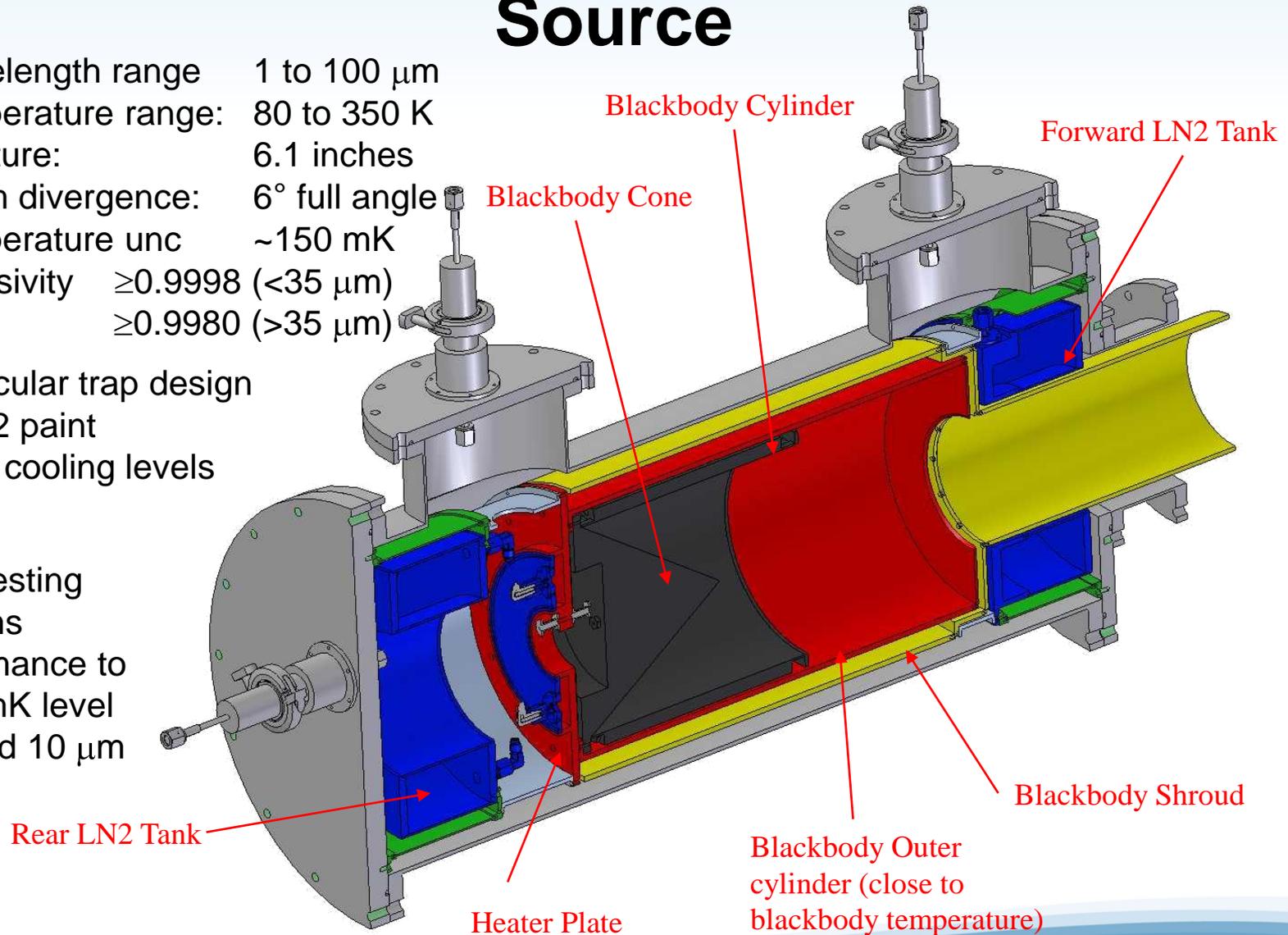
- ▶ FIRST calibrated in lab when built in 2005
- ▶ Re-calibrated in 2012 for absolute radiance response
 - Calibration data is collected by looking at warm blackbody, ambient blackbody, and LWIRCS (calibrator blackbody)



LWIRCS – Long Wave Infrared Calibration Source

- ▶ Wavelength range 1 to 100 μm
- ▶ Temperature range: 80 to 350 K
- ▶ Aperture: 6.1 inches
- ▶ Beam divergence: 6° full angle
- ▶ Temperature unc ~150 mK
- ▶ Emissivity ≥ 0.9998 ($< 35 \mu\text{m}$)
 ≥ 0.9980 ($> 35 \mu\text{m}$)
- ▶ Specular trap design
- ▶ Z302 paint
- ▶ Two cooling levels

NIST testing confirms performance to ~150 mK level at 5 and 10 μm

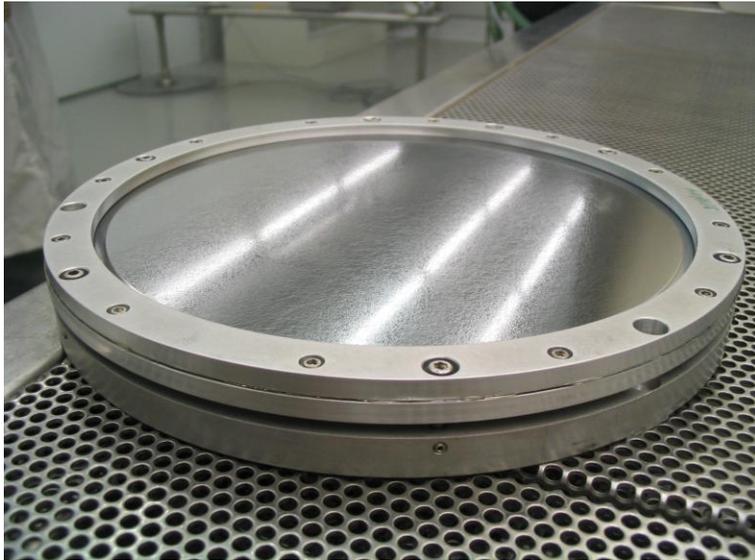


FIRST Response Calibration Data Details

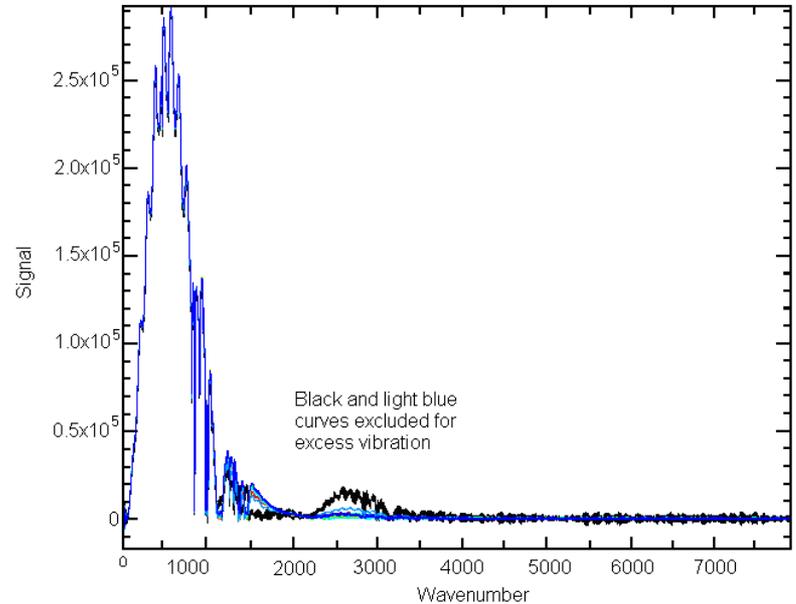
- ▶ Collect data of LWIRCS at a range of temperatures from 169 to 324 K, compare LWIRCS temperature to brightness temperature measured by FIRST
 - 30 min per set
 - Warm BB at ~324.5 K, ambient BB at ~294 K
- ▶ Data processed from interferograms to spectra, spectra averaged by target and scan direction, then spectra calibrated by calibration equation
- ▶ Unusual FIRST data processing
 - Hi and low gain channels
 - Need 20 bit dynamic range
 - Vibration effects
 - Phase drift

Vibration issues

- ▶ Beamsplitter is a stretched thin film
 - Ge on polypropylene
 - Can become a drumhead
- ▶ Vibration effects much reduced by sampling every HeNe laser fringe



FIRST Beamsplitter

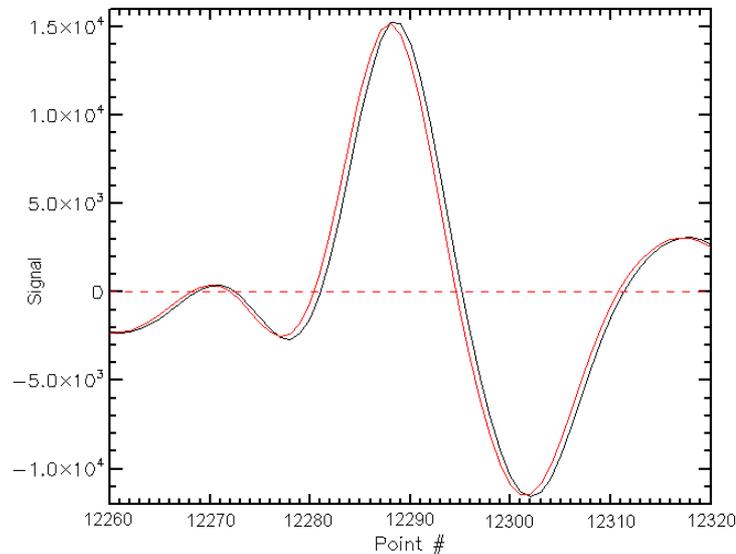


10 spectra from detector 3

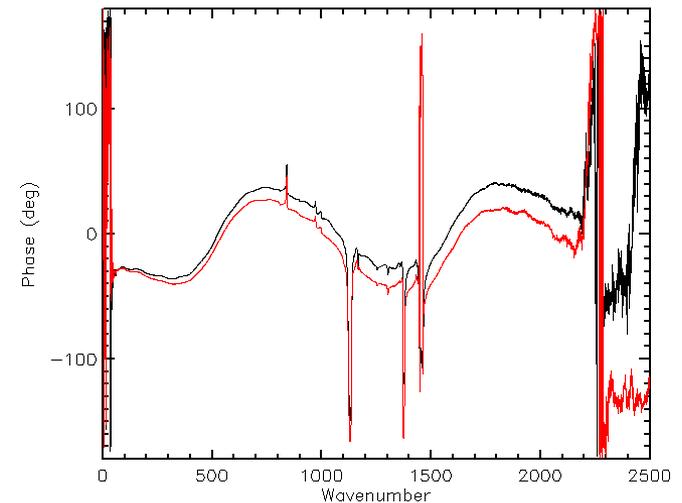
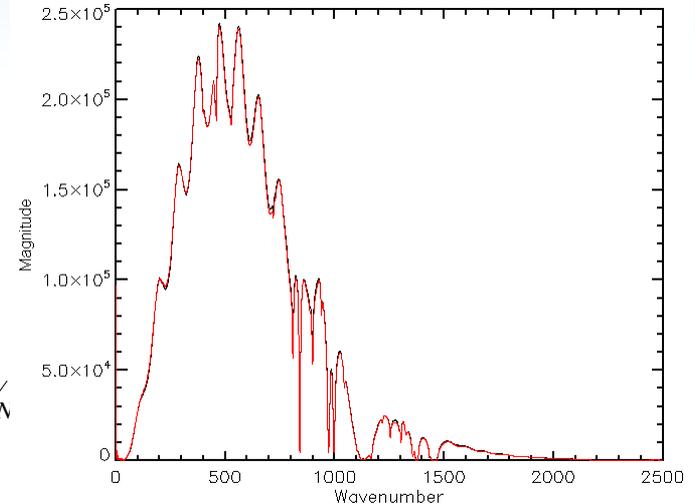
- ▶ Vibration effects appear around 2800 cm^{-1}
- ▶ Occasional noisy spectrum excluded

Phase shifting

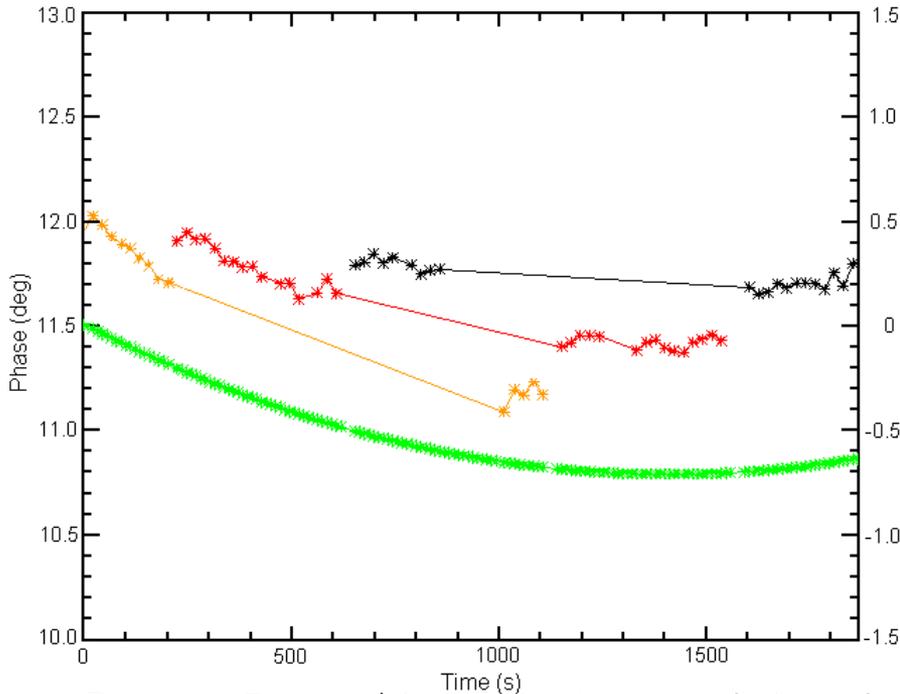
- ▶ FIRST metrology laser does not pass through beamsplitter, uses backside of moving interferometer mirror
- ▶ Thermal expansion changes optical paths differently, results in a shift in the phase of the spectrum



$$f_m = \sum_{n=0}^{N-1} F_n e^{i2\pi mn/N}$$



Phase alignment

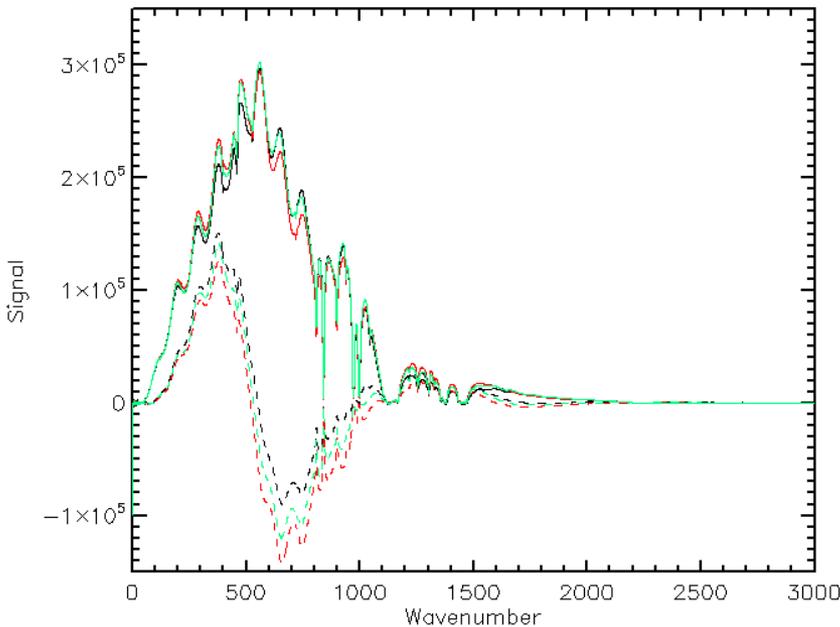


Phase at 514 cm^{-1} for all good spectra (left axis) for detector 3 in the forward direction of ABB (orange), LWIRCS (red) and WBB (black). Green curve is the phase drift curve (right axis)

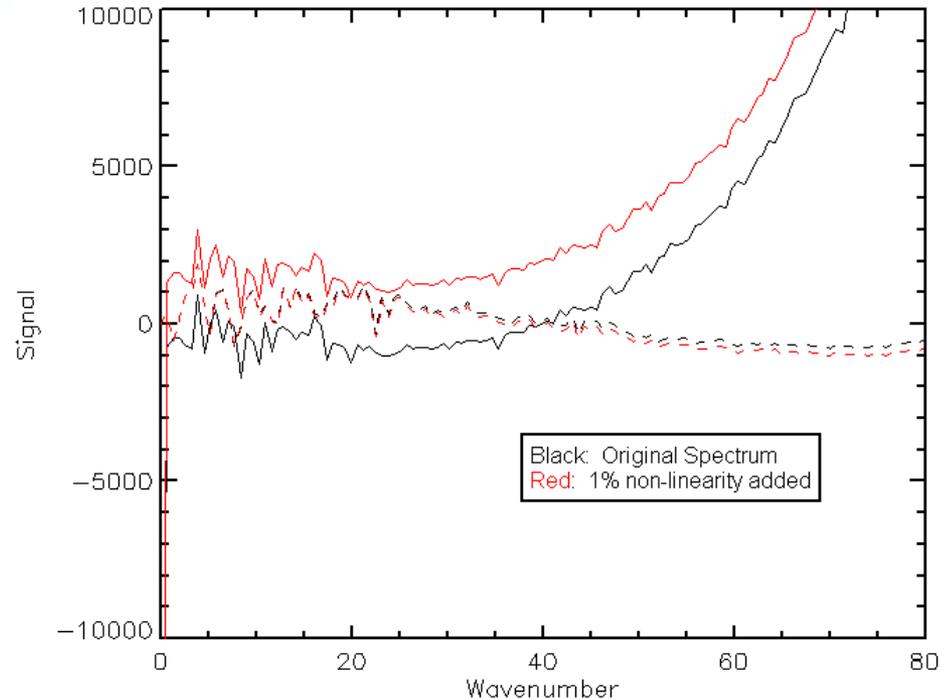
- ▶ FIRST has out of phase light, CANNOT phase correct each spectrum individually
- ▶ Adjust phase by fitting observed drift in WBB, ABB phase and adjusting phase of all spectra by amount of drift curve
- ▶ After this step, average all spectra by target and direction

Non-linearity

- Spectra are complex numbers



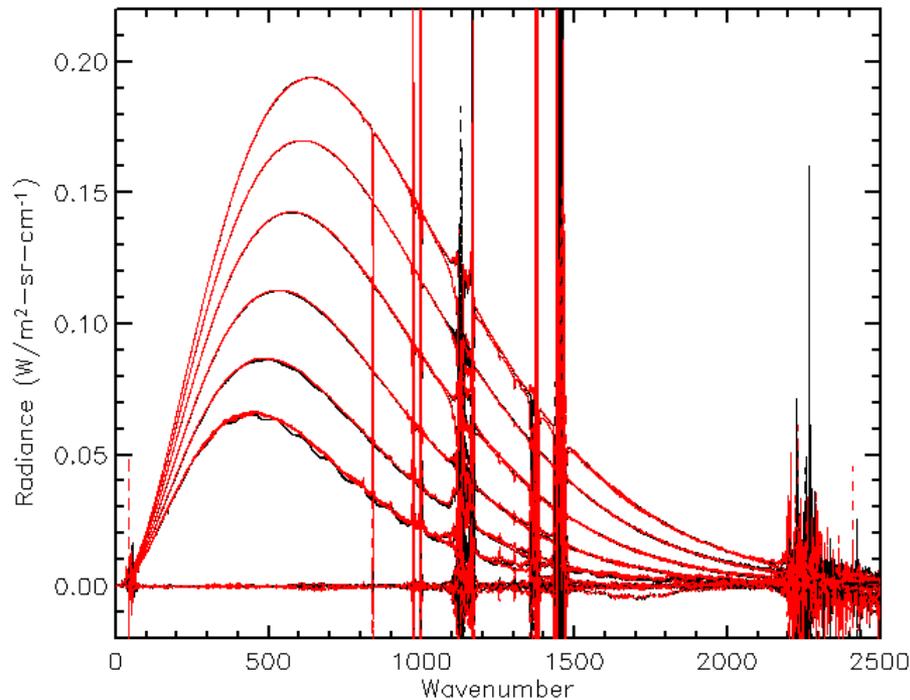
Three average spectra of the WBB in the forward direction from detector 3. Real part is solid line, imaginary part is dashed



- Non-linearity visible at low wavenumber end if present
- Non-linearity less than $\sim 0.3\%$ here

LWIRCS radiance

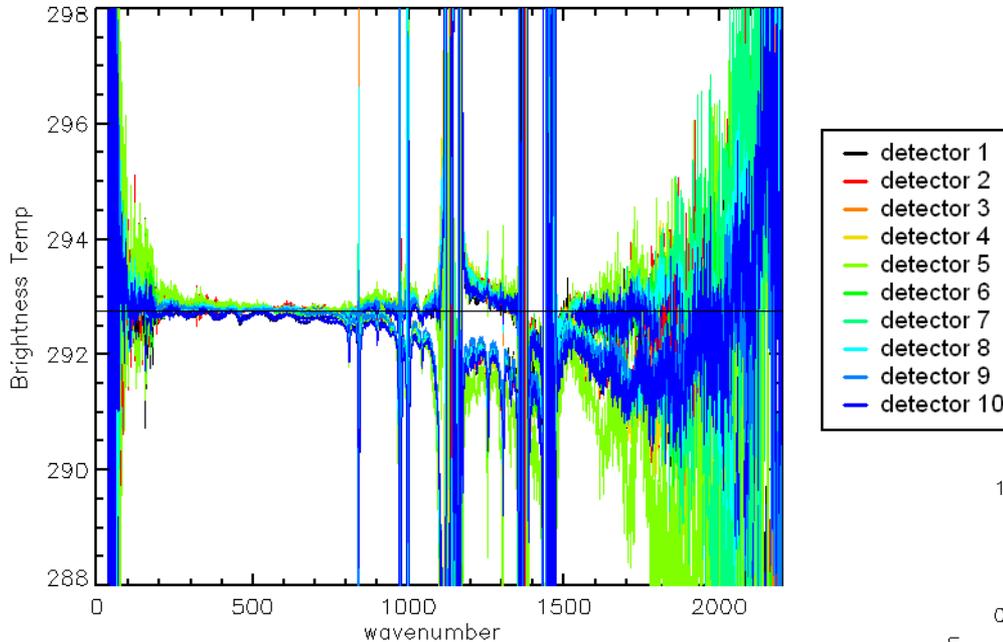
- ▶ Calibrated LWIRCS spectra
 - Imaginary part zero in calibrated spectra



- ▶ Radiances from other detectors overlap these
- ▶ Look like Planck functions on this scale

LWIRCS radiance from detectors 1 and 2 for LWRICS at 324, 310, 293, 271 and 225 K.

Calibration results at 292.76 K

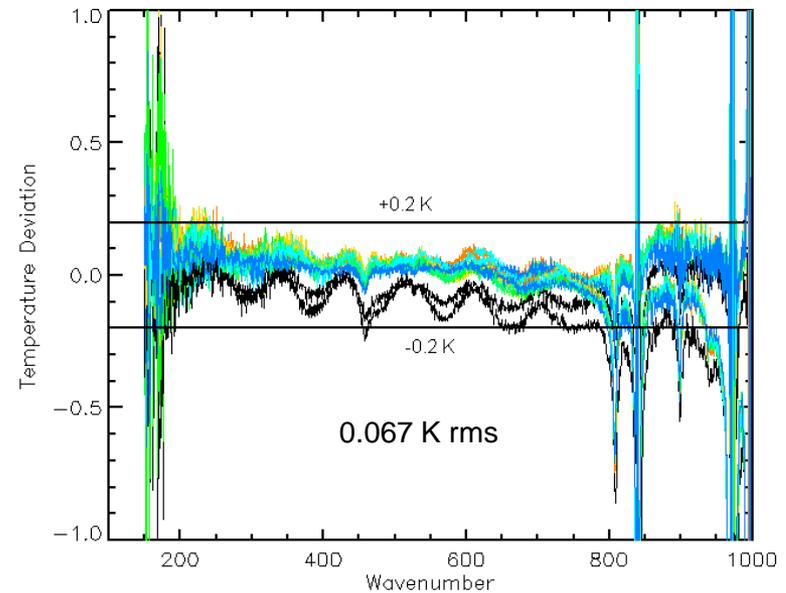


LWIRCS brightness temperature at 292.76 K

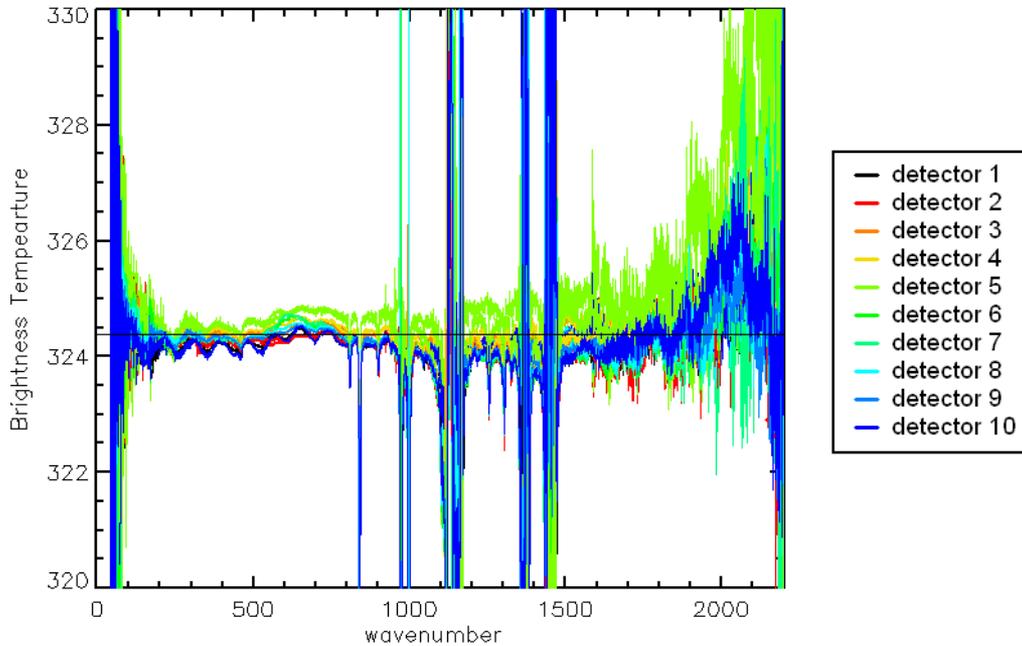
- ▶ Brightness temperature within 0.2 K of LWIRCS sensors 200 to 1000 cm^{-1}
- ▶ Vibration noise above 900 cm^{-1}
 - Different for two scan directions
 - Same for all detectors

- ▶ LWIRCS at room temperature
 - Directly compares LWIRCS, ABB temp sensors
 - Temp sensors within 0.1 K of each other

LWIRCS temperature deviation excluding detectors 2, 5 and 10



Calibration results at 324.38 K

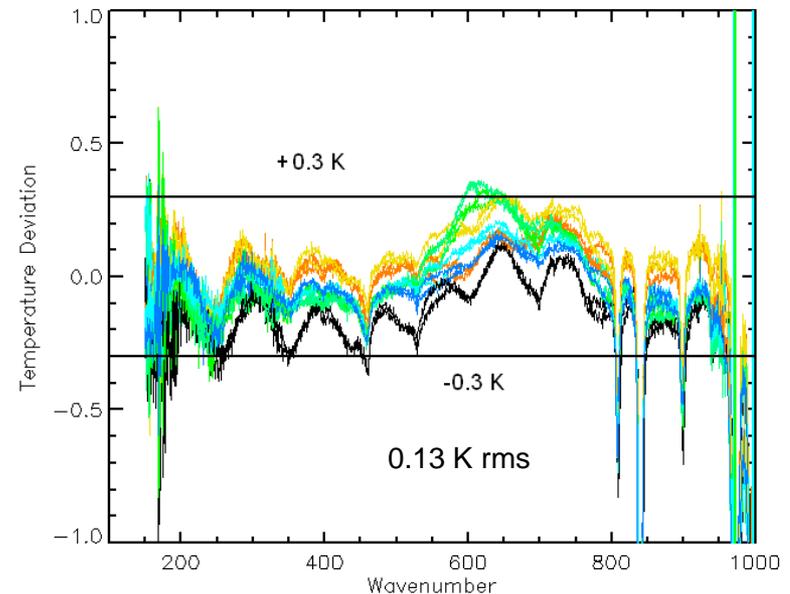


LWIRCS brightness temperature at 324.38 K

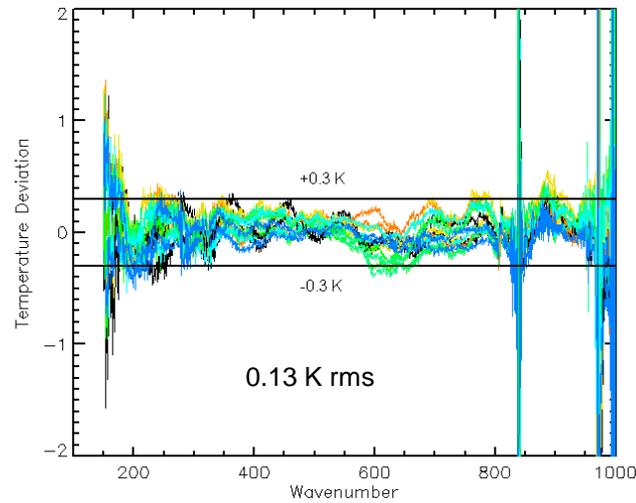
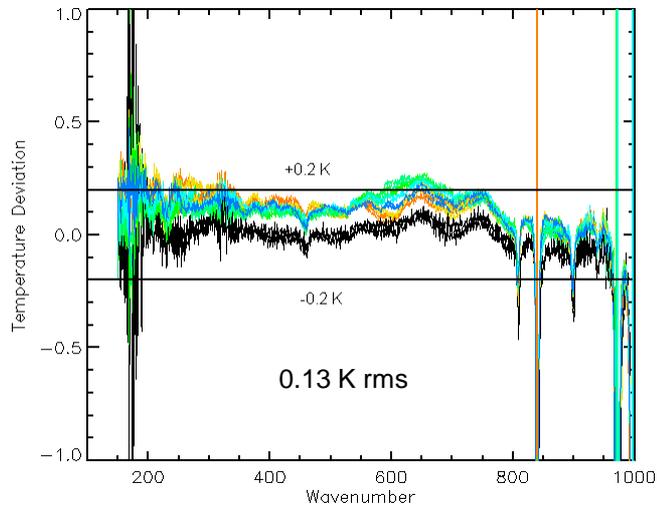
- ▶ Brightness temperature within 0.3 K of LWIRCS
- ▶ Exclude detectors 2, 5, 10
 - 2, 10 excess noise in some cases
 - 5, 10 stray light response

- ▶ LWIRCS at WBB temperature
 - LWIRCS and WBB agree to within 0.1 K
- ▶ Noise mainly systematic
 - Same in both scan directions
 - Less vibration here

LWIRCS temperature deviation excluding detectors 2, 5 and 10

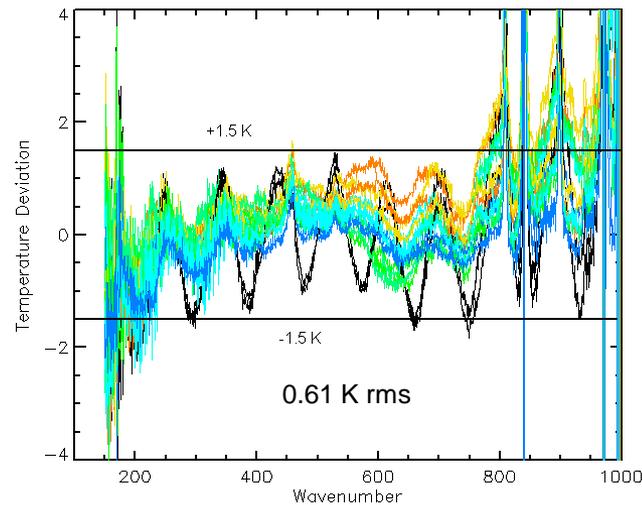
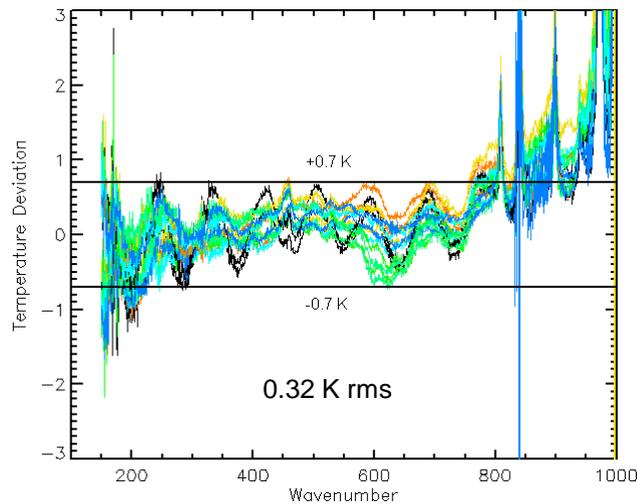


More Calibration Results



LWIRCS
temperature
deviation
270.55 K

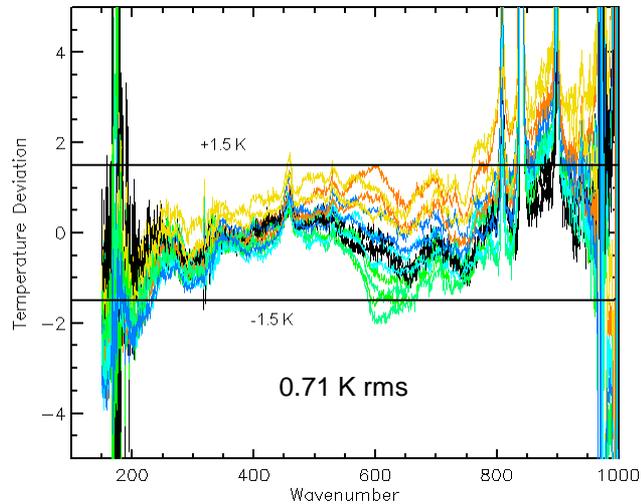
LWIRCS temperature deviation 310.34 K



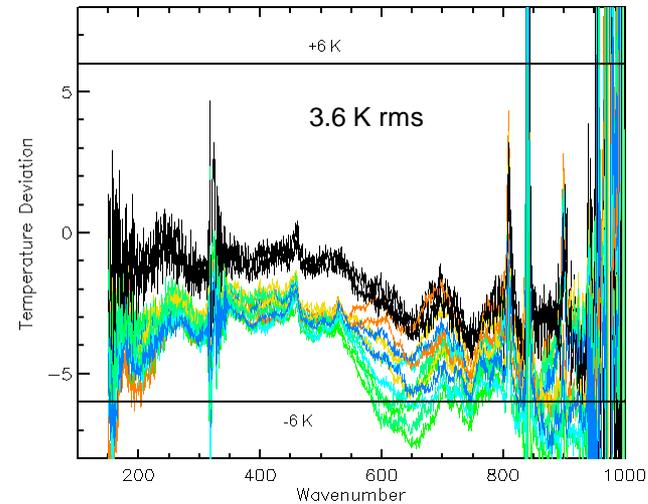
LWIRCS
temperature
deviation
225.18 K

LWIRCS temperature deviation 247.42 K

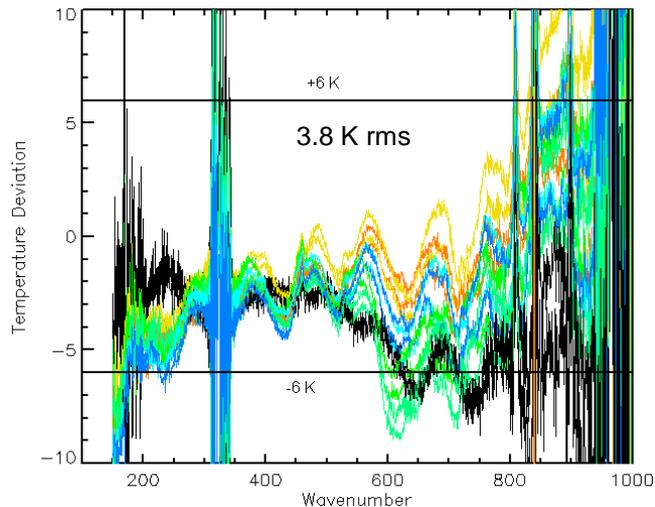
More Calibration results



LWIRCS temperature deviation 209.41 K



LWIRCS temperature deviation 189.33 K



LWIRCS temperature deviation 169.06 K

- ▶ Noise increases with falling temperature, reduces high end of range
- ▶ Deviations larger below 200 K
- ▶ Deviations mainly systematic

Error propagation effects

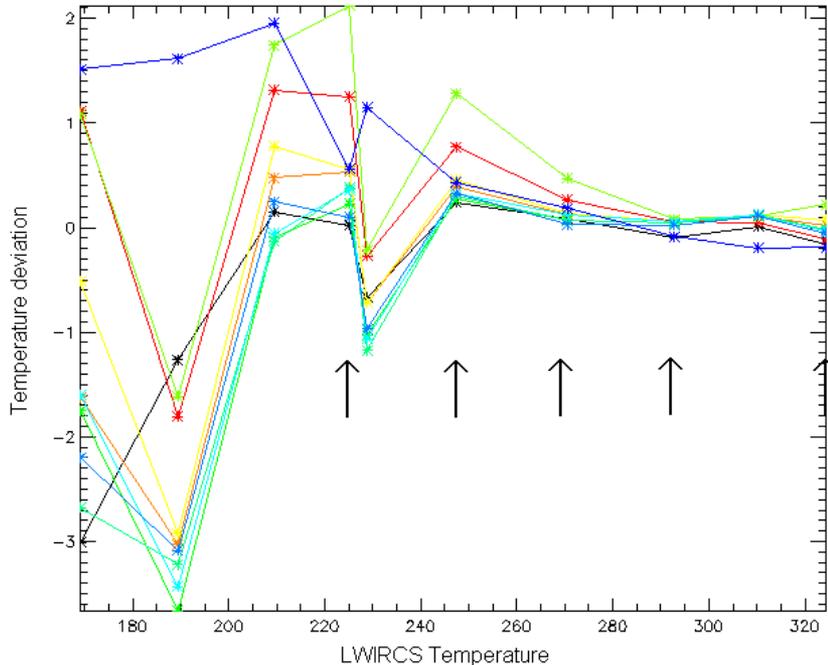
- ▶ Error from ABB, WBB spectra propagate into errors in target spectra
 - Amount rises significantly with temperature

Target Temp	200 cm ⁻¹	500 cm ⁻¹	800 cm ⁻¹
225 K	0.9 K	1.1 K	1.4 K
169 K	1.7 K	2.7 K	5.4 K

Propagated error in target assuming
324.5 K WBB with 0.3 K error
293 K ABB with 0.2 K error

- ▶ Some of increased low temperature deviation is simply an effect of error propagation

More Calibration Results



LWIRCS temperature deviation (average from 459.6 to 559.9 cm^{-1}) forward direction only. Data taken during two vacuum cycles, arrows show data from one cycle

- ▶ Deviation at low temperatures shows no consistent trend with temperature
- ▶ Source probably systematic effect that varies over time combined with error propagation
- ▶ Deviation at low temperature may vary with vacuum cycle

Conclusions

- ▶ FIRST absolute accuracy: 1.5 K or better (peak deviation) for temperatures >200 K from 200 to 800 cm^{-1}
- ▶ From 270 to 330 K (near ABB, WBB temperatures), FIRST meets design accuracy goals
 - 0.2 K ($k=1$) 170 to 1000 cm^{-1}
- ▶ No additional corrections needed in calibration equation
 - No significant non-linearity
 - ABB, WBB consistent with LWIRCS
- ▶ Observed deviations probably due to small systematic effect that changes over time combined with error propagation