



Using Lunar Observations to Radiometrically Assess the CUMULOS' Visible Camera

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Outline



- CUMULOS Sensor
- Objective and Lunar Cal Overview
- Observed Anomalies
 - *DN Dips Along FPA*
 - *Ghost Imaging*
- Lunar Calibration
 - *Explanation of the ROLO Model*
 - *Application and Results of the ROLO Model*
- Summary

The CubeSat Multispectral Observation System (CUMULOS)

- The NASA payload (ISARA), Aerospace bus, was launched from NASA Wallops on Nov 12, 2017 ISARA goal is to demonstrate downlink data rates for CubeSats to >100 Mbps. NASA's ISARA will transmit a Ka tone to ground station. Experiment consists on measuring antenna pattern.
- CUMULOS operations commenced on 8 June 2018 following the successful completion of the ISARA mission
- Primary Requirement:
 - To develop **optimal methods for the operation** of passively cooled COTS sensors and cameras and determine their **suitability to perform weather/environmental monitoring** missions
- CUMULOS: 3 COTS cameras (1U) in a 3U CubeSat
 - Visible camera – **0.4 – 0.8 μm** ;
 - CMOS chip (Aptina MT9M001C12STM)
 - 1280 x 1024 pixels
 - SWIR camera – **0.9 - 1.7 μm**
 - LWIR camera – **7.5 - 15.5 μm**
- No requirements on calibration accuracy or precision

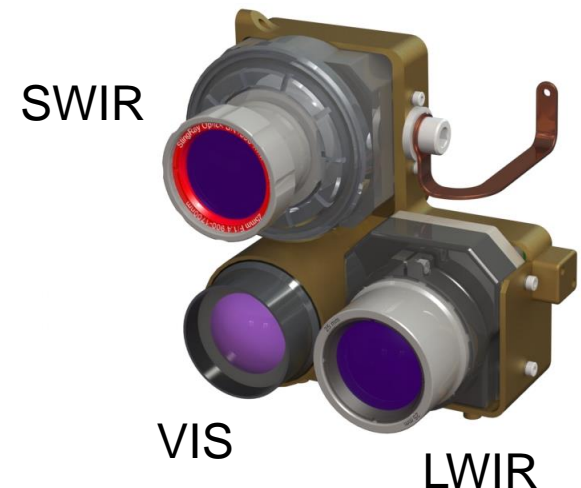
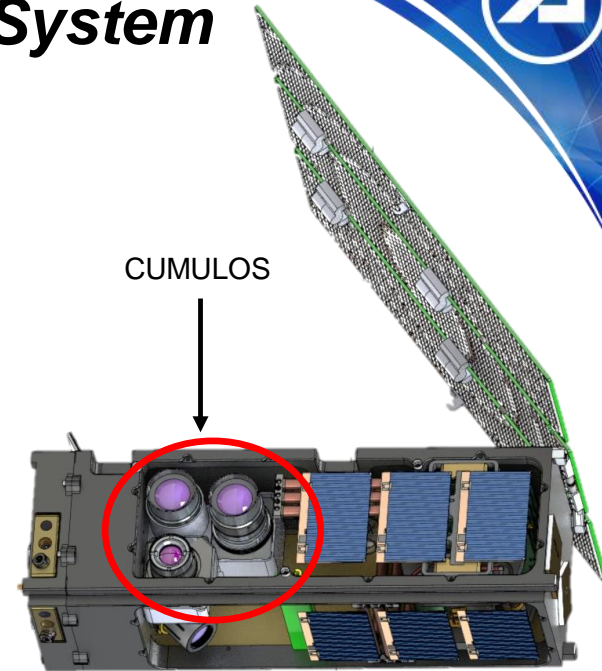
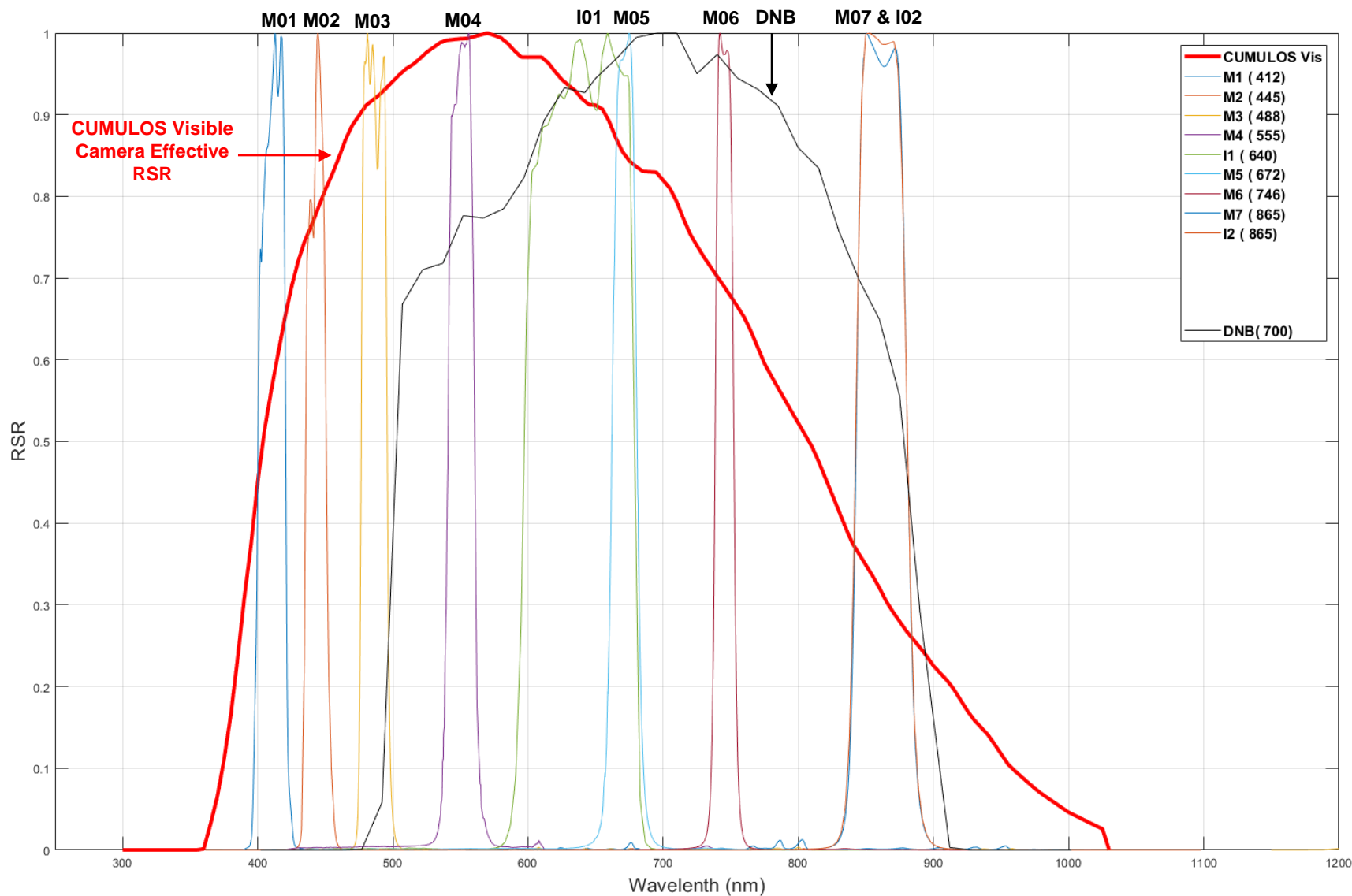


Image and slide content from

RSR Comparisons

SNPP VIIRS and CUMULOS Visible Camera



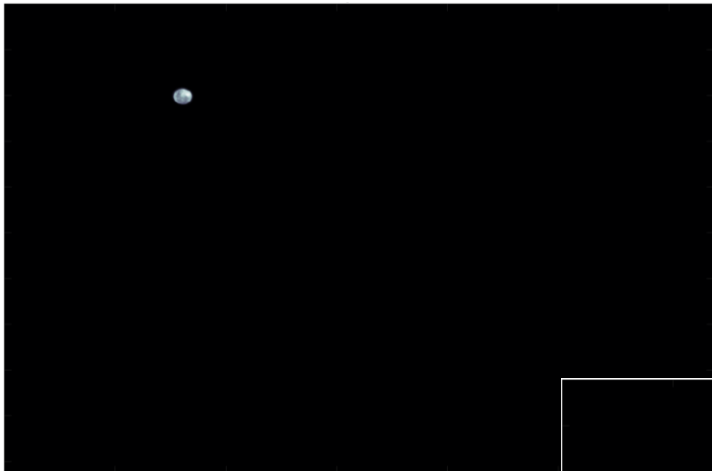
Objective and Lunar Cal Overview



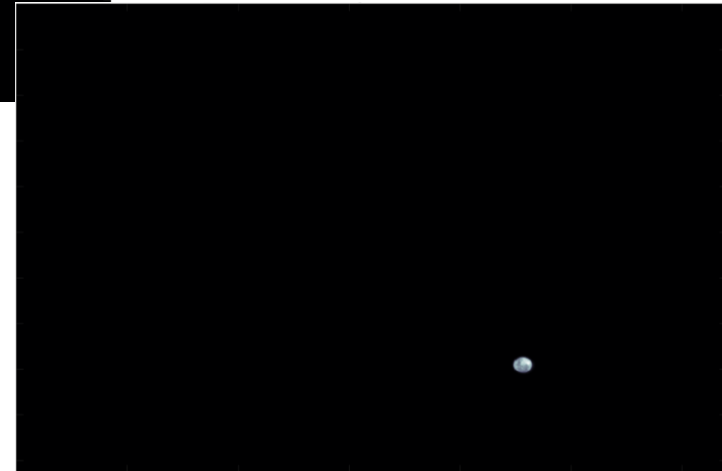
- Motivation: COTS sensors without onboard calibrators will only increase given the pLEO/CubeSat mentality and using the Moon to radiometrically assess said sensors is of absolute importance
- This effort was funded by The Aerospace Corporation's FY19 Independent Research and Development (IRAD) program with a goal of utilizing the Moon to radiometrically assess the CUMULOS Visible Camera
- Process:
 - *Sampling CUMULOS' FPA:*
 - Given the data downlink limitations it was near impossible to sample the entire Visible Camera with the Moon. Camera does not have zooming capability so to image entire FPA.
 - Observe Moon at center and four quadrants of the FPA
 - *Apply Flat Field Correction derived from lab tests*
 - *Dark Current Offsets were determined from neighboring (before and after) dark space collects about lunar collects*
 - Dark Current Offset Trending was assessed and determined negligible in change (fraction of a DN)

CUMULOS Lunar Collect

Feb 19, 2019 at 116 us Integration



1280 x 1024 pixels

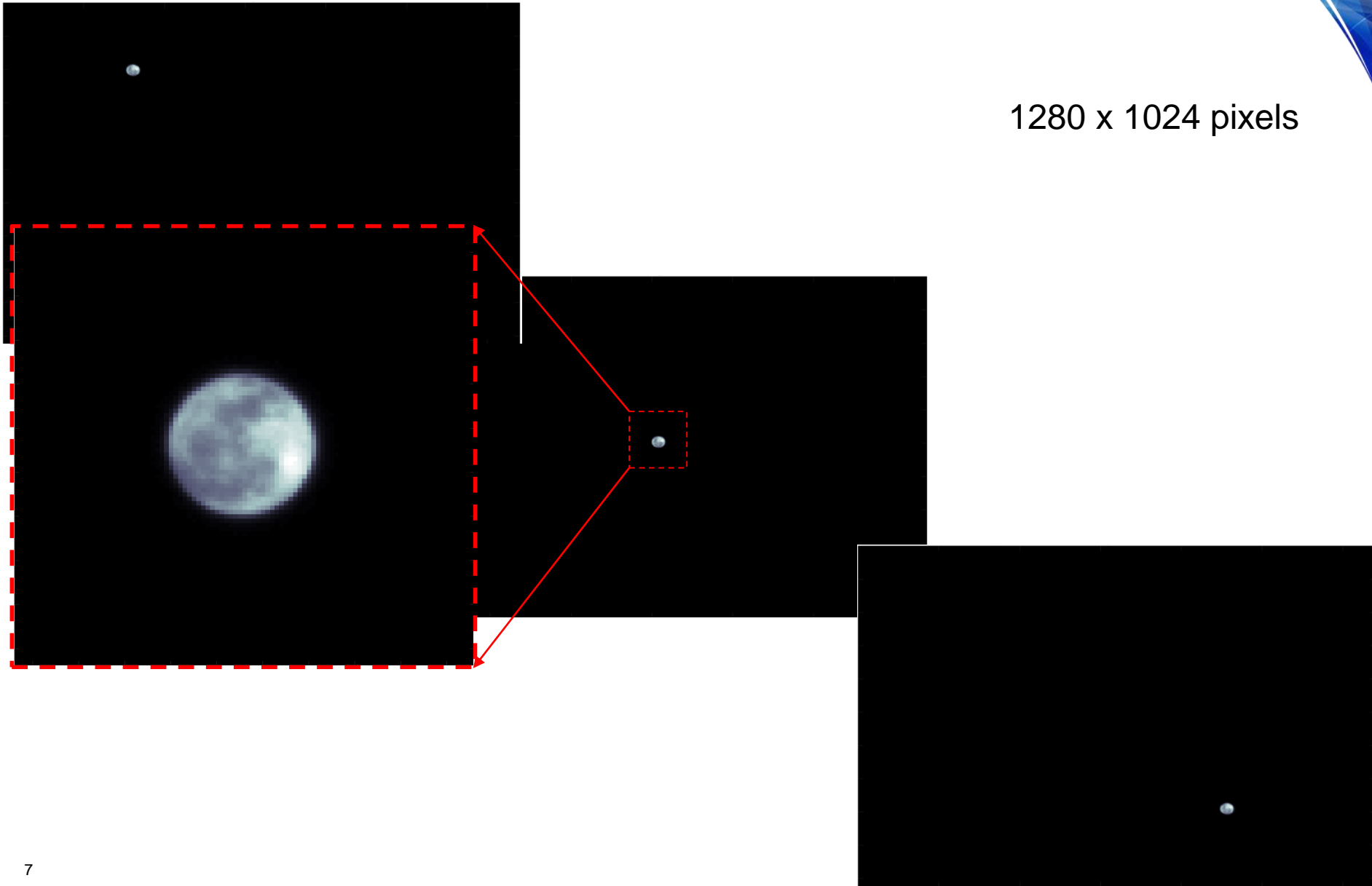


CUMULOS Lunar Collect

Feb 19, 2019 at 116 us Integration



1280 x 1024 pixels



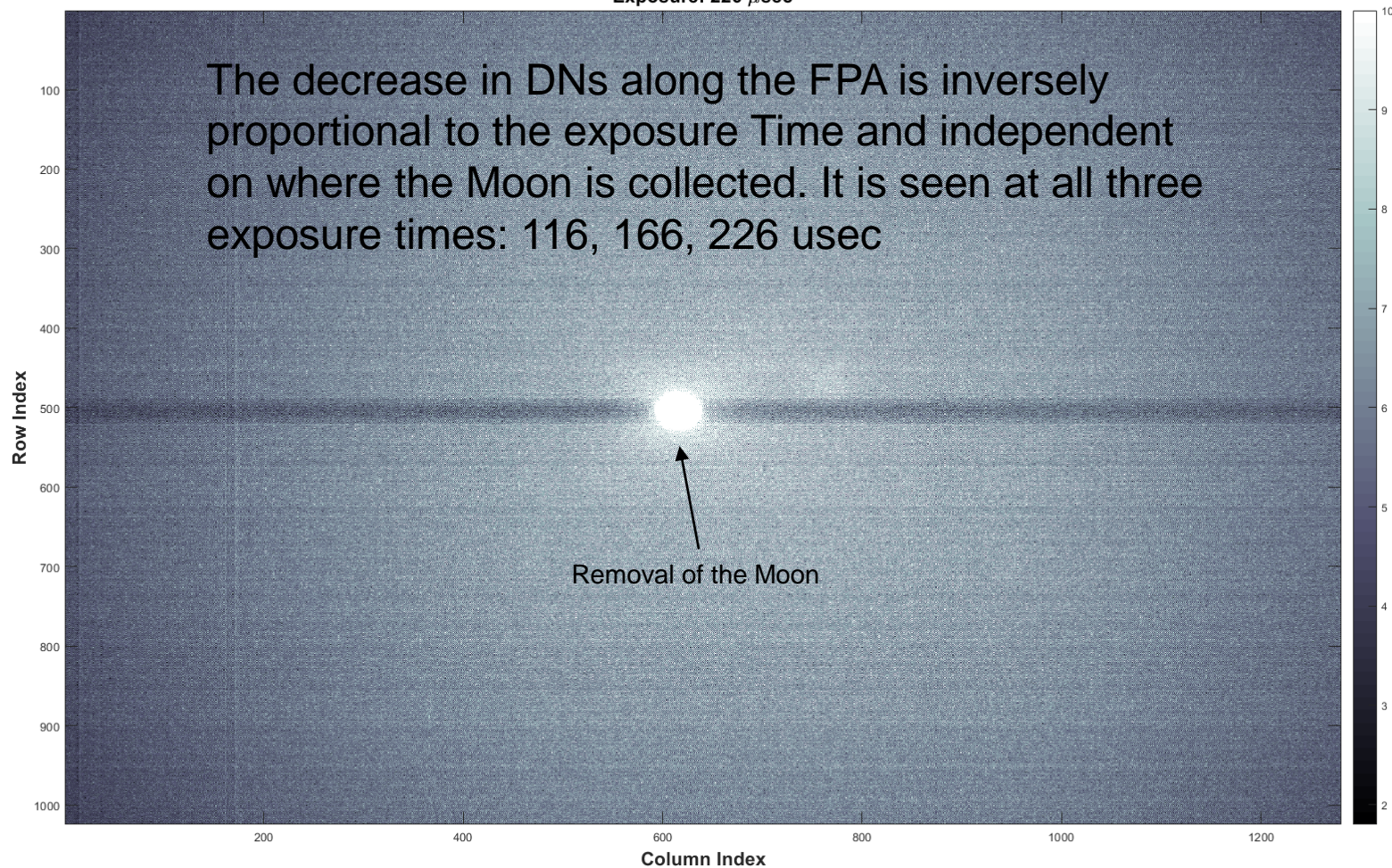


Anomalies: Dip in DNs

Dip in DNs



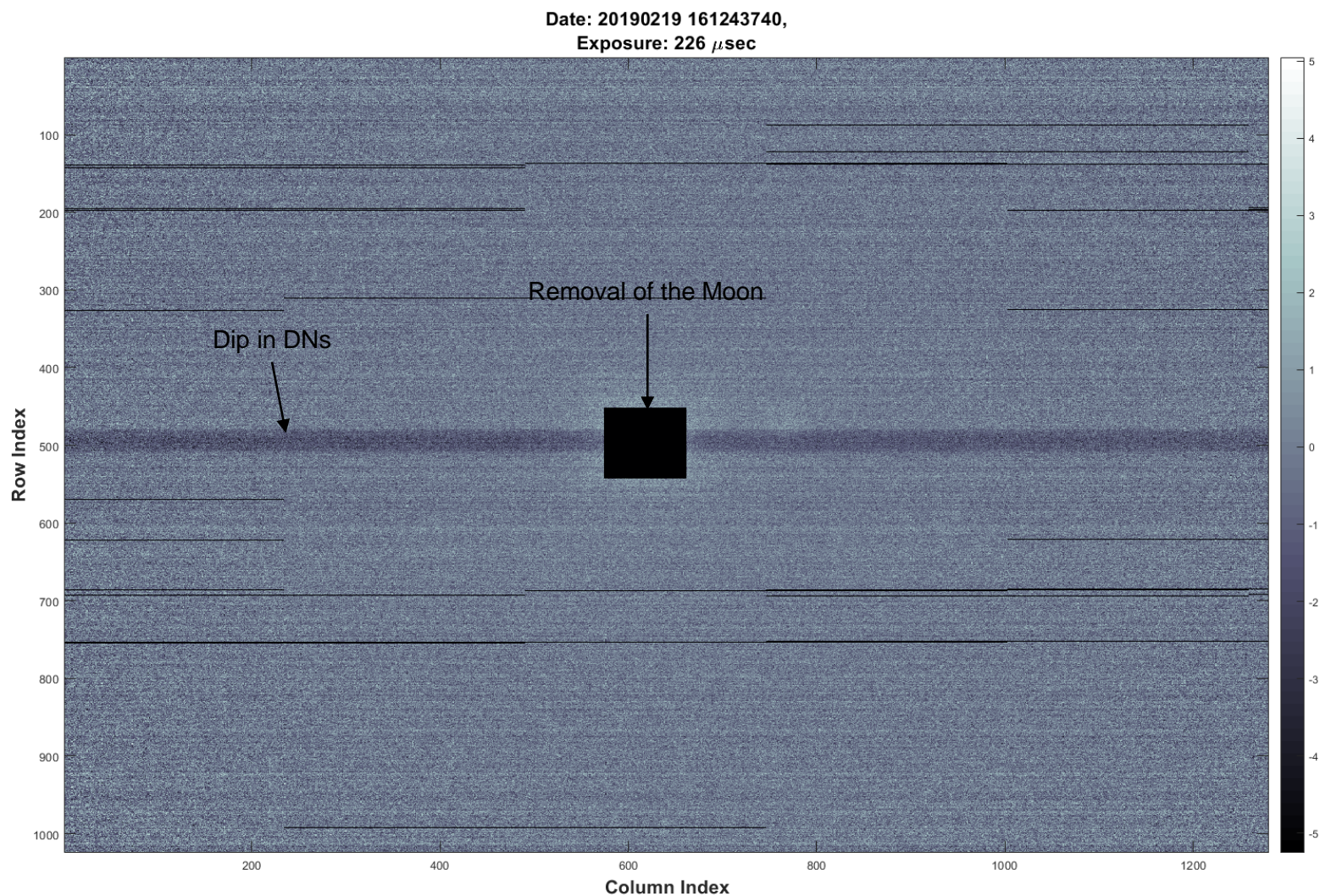
Date: 20181123 053018740,
Exposure: 226 μsec



No Dark field nor Flat field applied

Dip in DNs

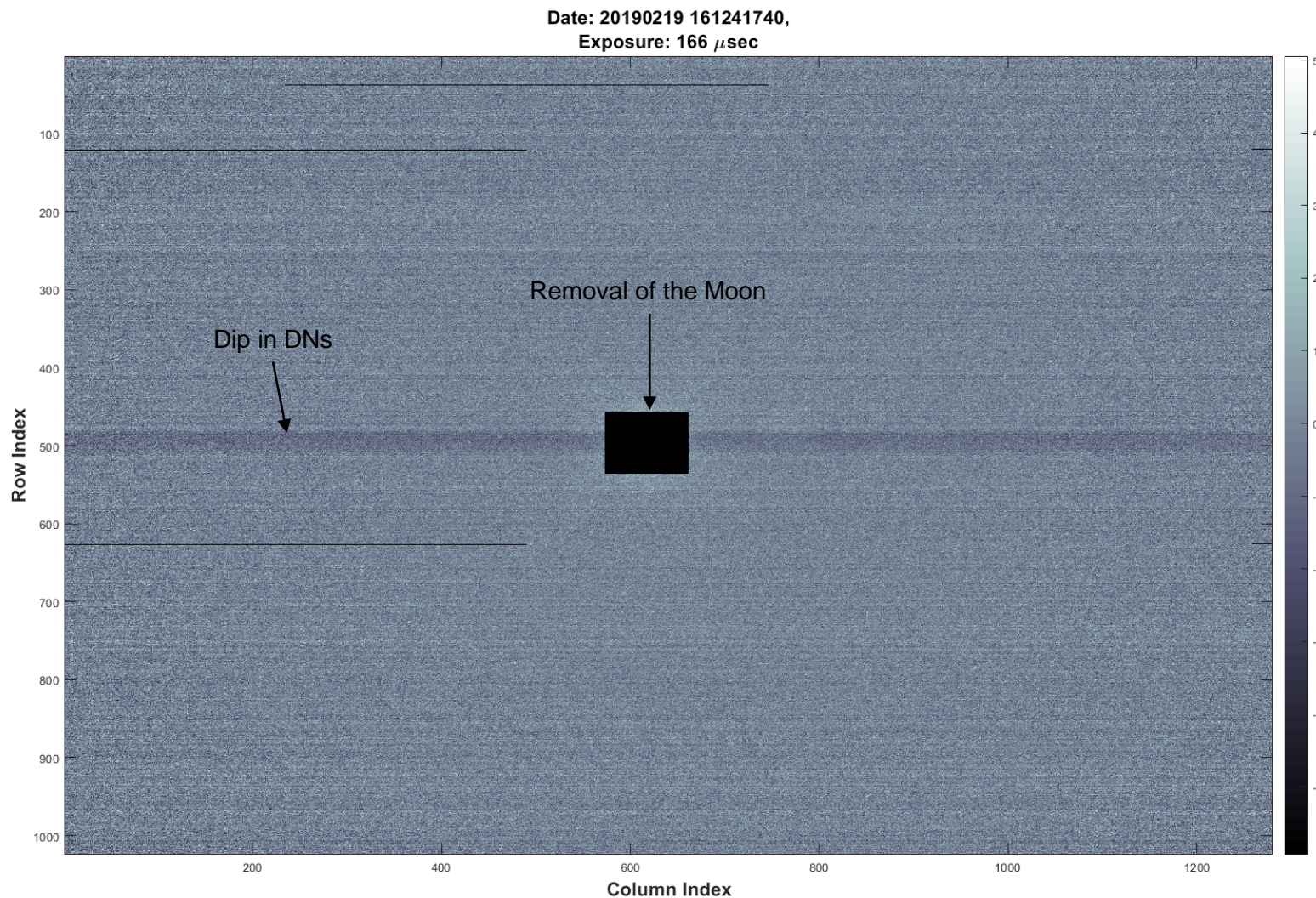
Exposure Time: 226 μ s



Dark field and Flat field applied

Dip in DNs

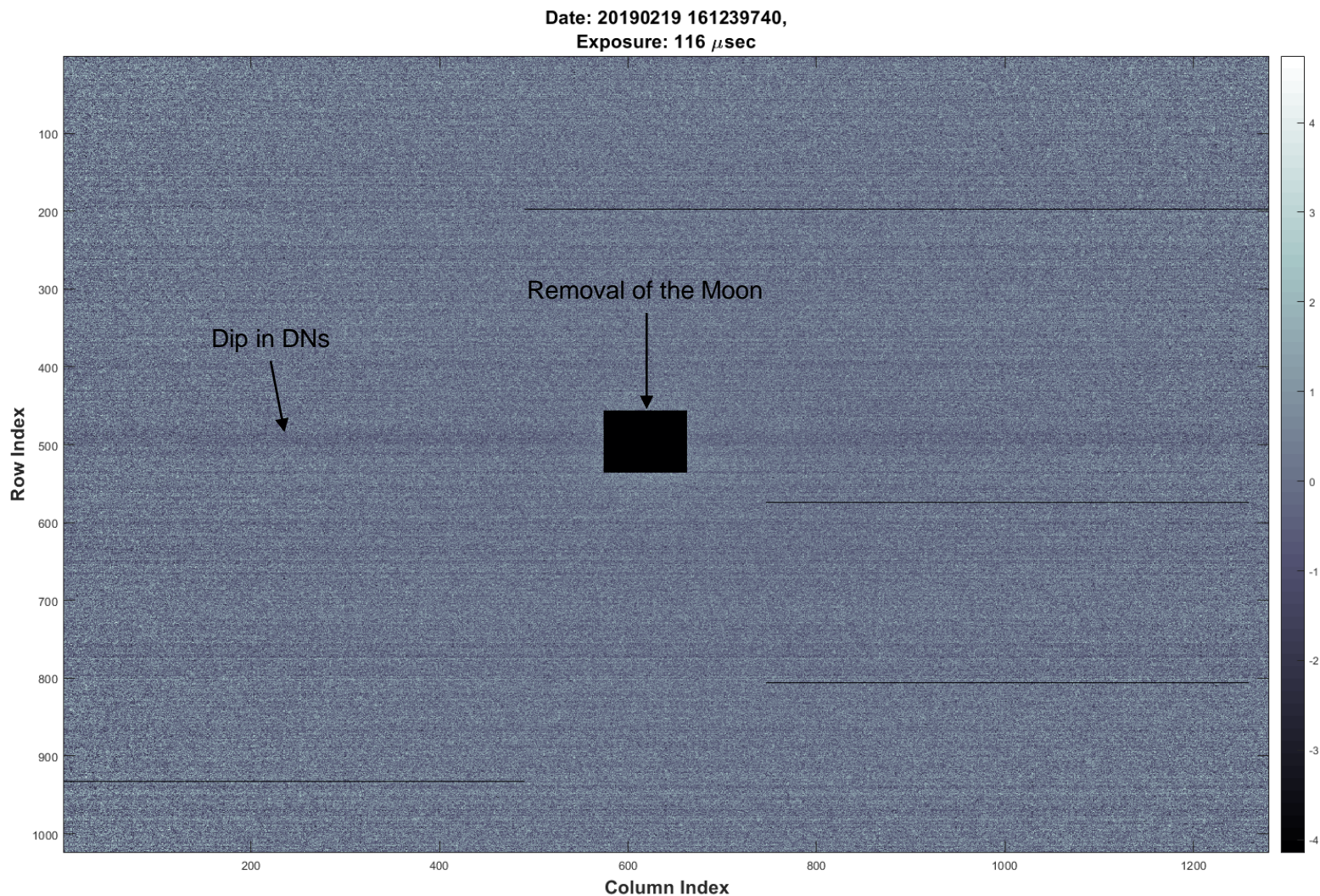
Exposure Time: 166 μ s



Dark field and Flat field applied

Dip in DNs

Exposure Time: 116 μ s



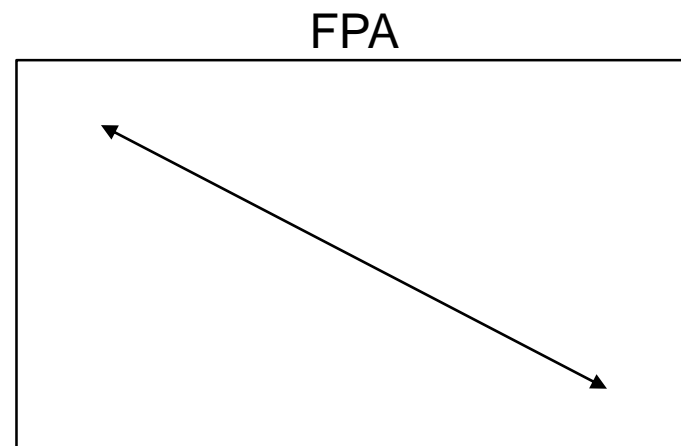


Anomalies: Ghosting / Retro-Reflection

Ghosting or Retro-Reflection

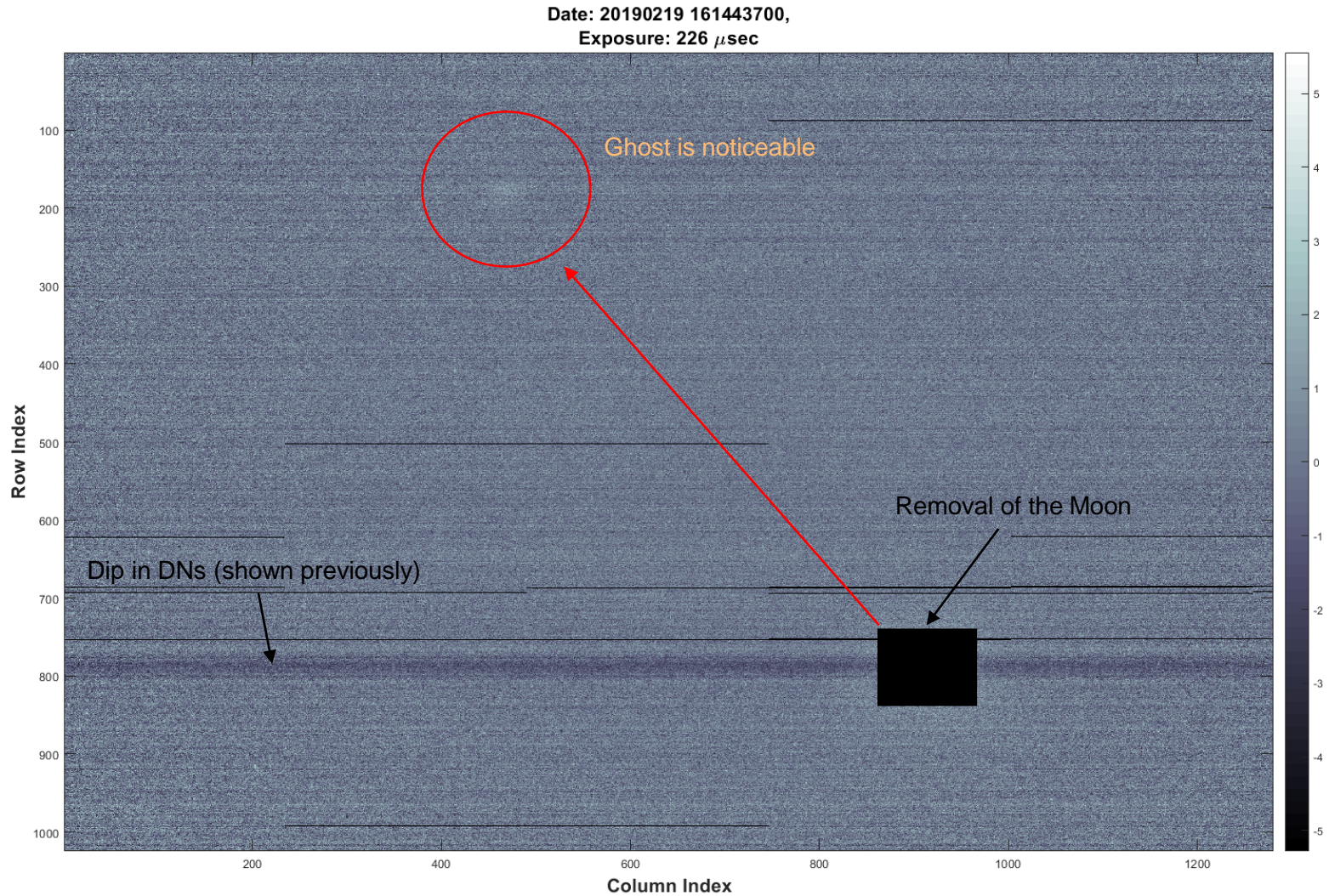


- There are unexpected patterns during Lunar collects that are away from the Moon on a certain part of the FPA for Exposure times 116 and 226 usec.
- These patterns are seen before and after flat-field and dark subtraction is applied and are believed to be ghost images
- What is a ghost image?
 - *A feature or shape at the focal plane of a camera or other **optical** instrument that is not present in an actual scene, or an unfocused duplicate image that is overlaid upon a desired image. **Ghost** images, or ghosts, are caused by reflections from the surfaces of lenses.*
- Ghost image is only noticed when the Moon is observed in the center, upper left, or lower right quadrants



Ghost Image

Exposure 226 usec

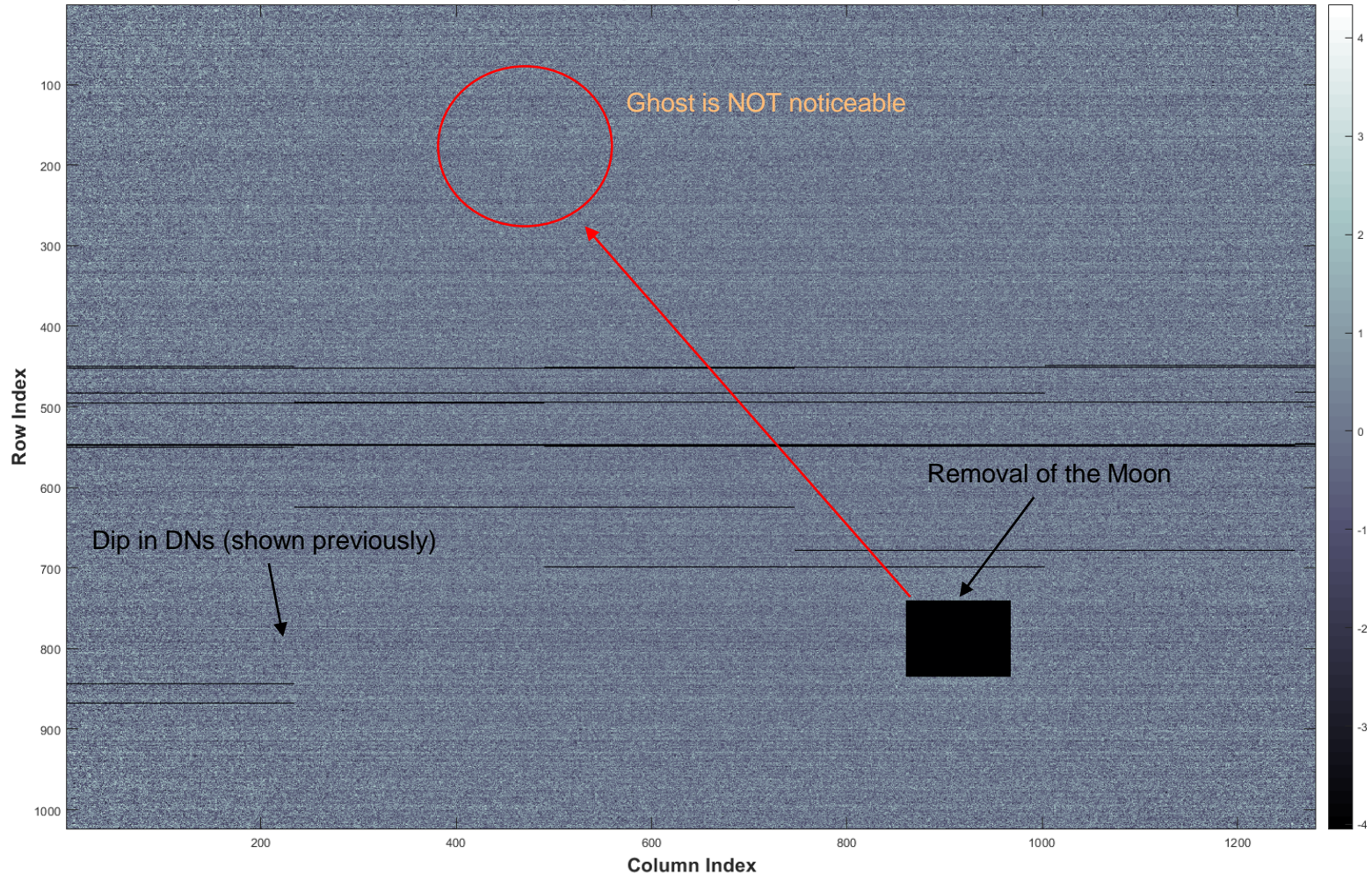


Ghost Image

Exposure 116 usec

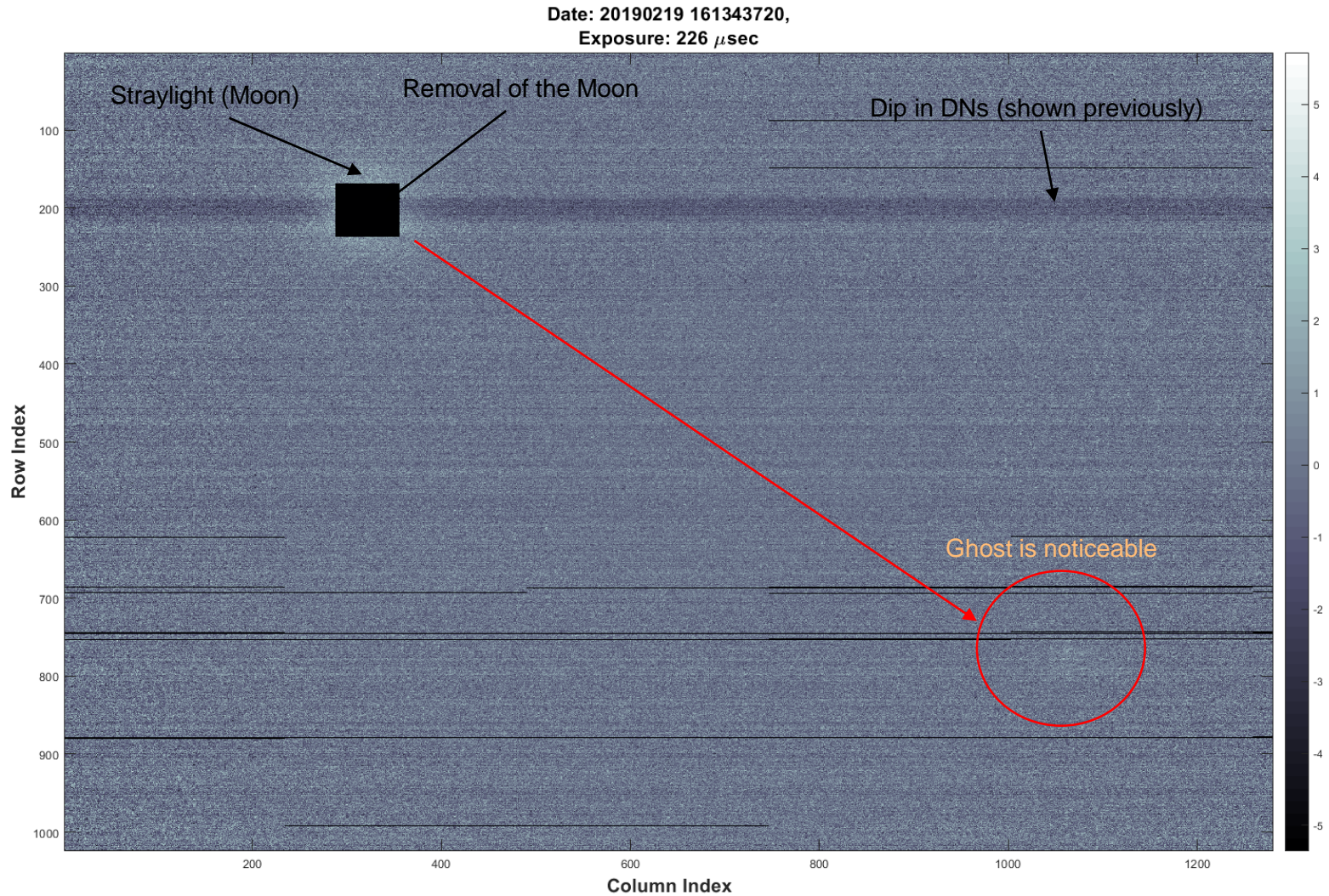


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Exposure: 116 μ sec



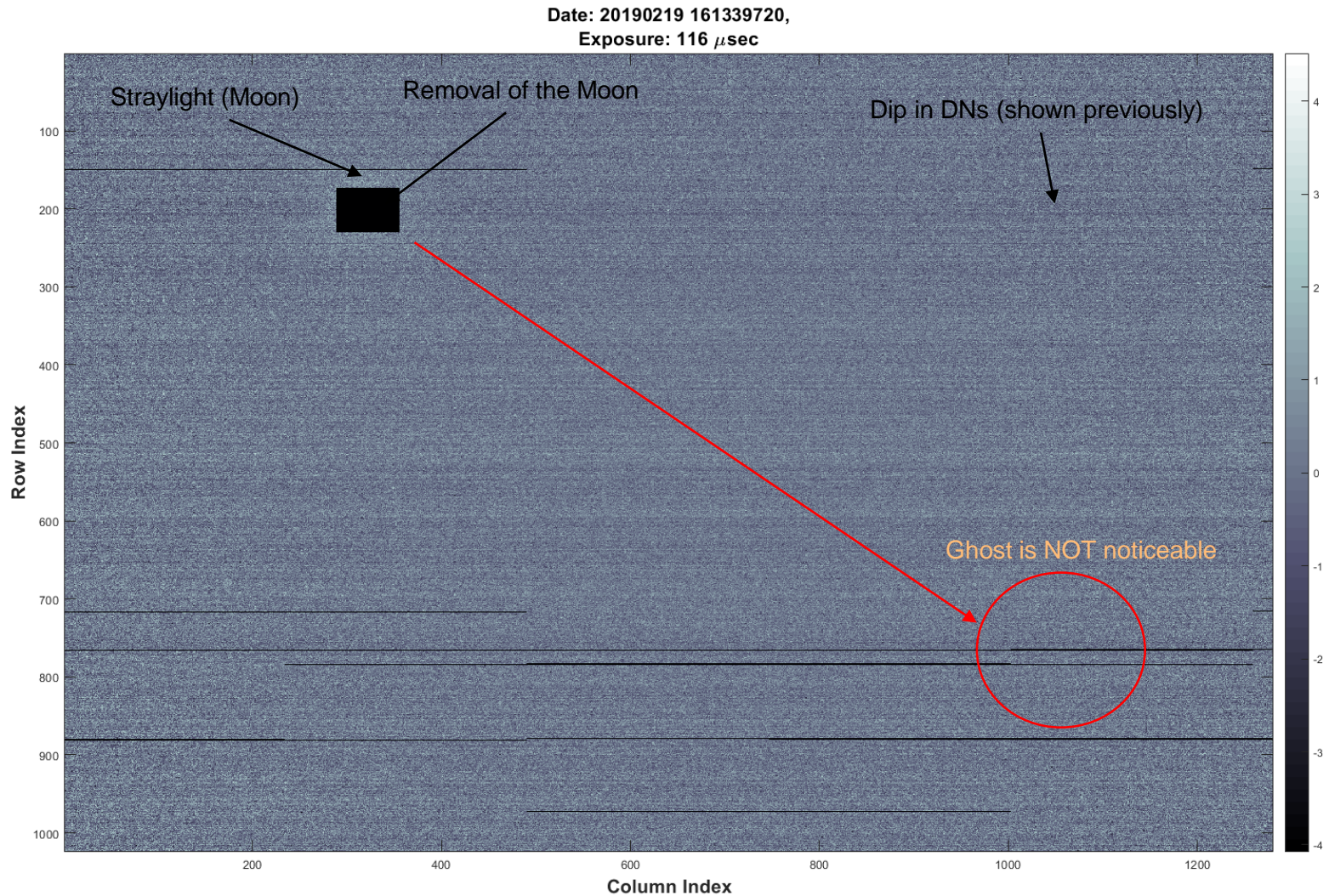
Ghost Image

Exposure 226 usec



Ghost Image

Exposure 116 usec





Lunar Observations and Model

Using the Moon for Radiometric Cal Assessment

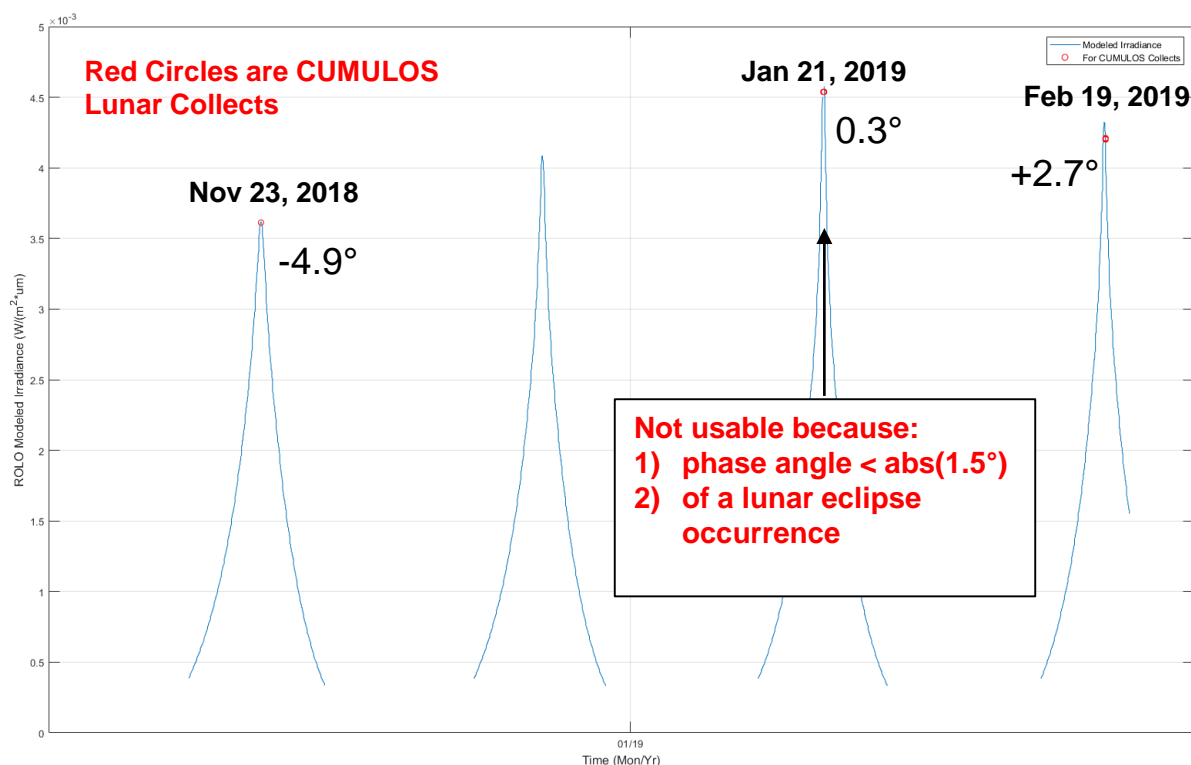


- The Moon is a stable calibration source
 - *Tidal lock so does not rotate with respect to the observer*
 - *Most significant variation in Lunar Disk irradiance is due to phase change*
- NOAA and NASA uses observations of the Moon in conjunction with a Lunar Model (ROLO) for calibration purposes to determine long-term radiometric stability for VIIRS and MODIS, respectively, even though these sensors have on-board calibrators
- The USGS Robotic Lunar Observatory (ROLO) model
 - Kieffer, H., Stone, T., “The Spectral Irradiance Of The Moon” June 2005
 - *was developed using 6+ years of ground-based observations*
 - *applicable 0.35 – 2.45 um (bounds current CUMULOS spectrum response)*
 - *absolute uncertainty: 5-10% (irradiance scale)*
- Process:
 - *Utilize implementation of USGS Robotic Lunar Observatory (ROLO) model (previous IRAD)*
 - *Aerospace Satellite Orbit Analysis Program (SOAP) used for orbital and body geometry and image processing techniques to determine Moon-illuminated Pixels*
 - *Irradiance of CUMULOS is calculated using the DN coefficient provided by Dee Pack’s Alpha-Lyra star calibration (methodology presented during this Conference)*

CUMULOS Lunar Collects



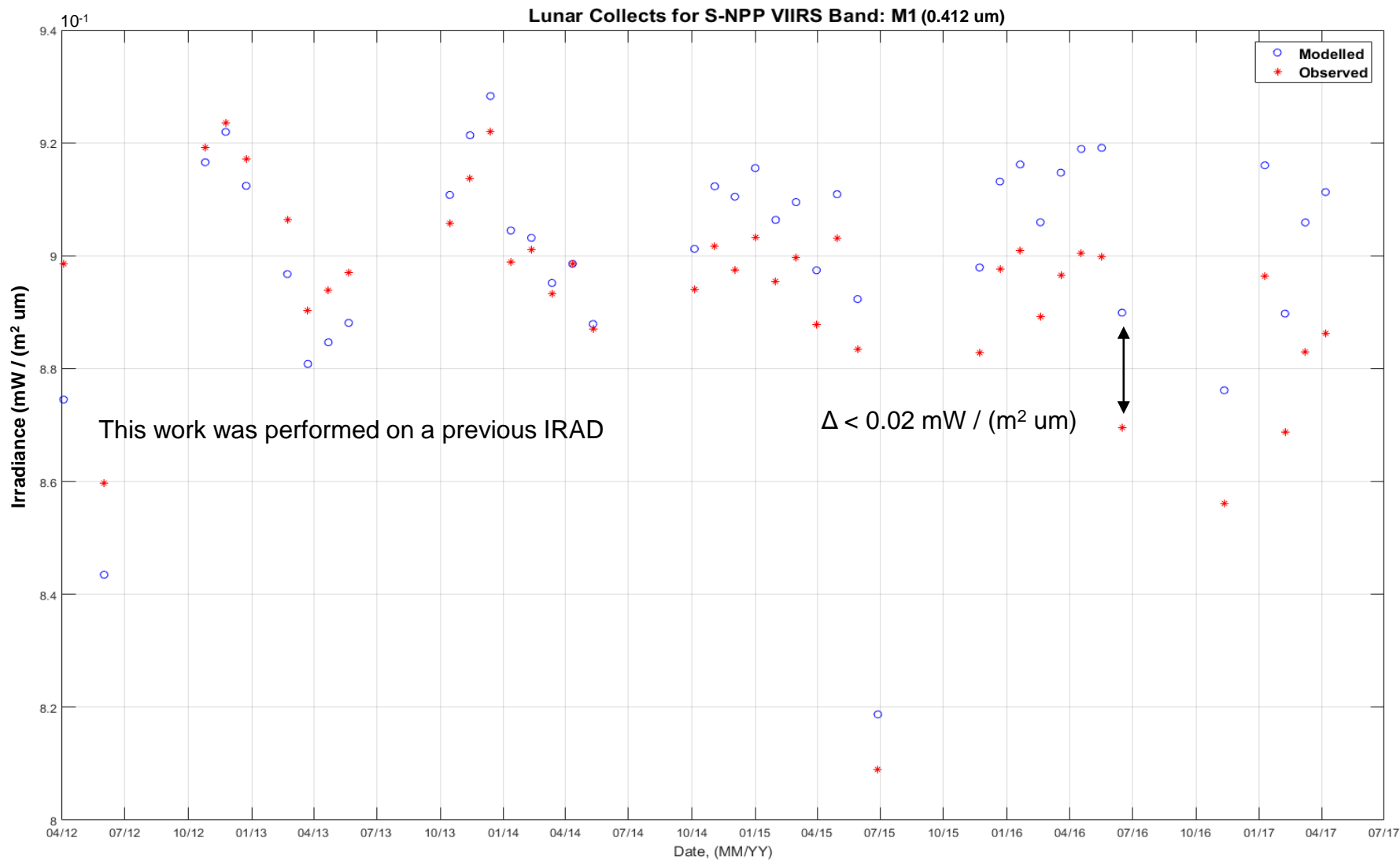
- Even though there have been three exposure times to collect the Moon (116,166, and 226 us) the latter 2 saturate when capturing the Moon
- To Date 116 us Exposure Time Lunar Collects
 - 2 Collects: 23-Nov-2018, Phase Angle: -4.9°
 - 6 Collects: 21-Jan-2019, Phase Angle: 0.3°
 - 6 Collects: 19-Feb-2019, Phase Angle: 2.7°



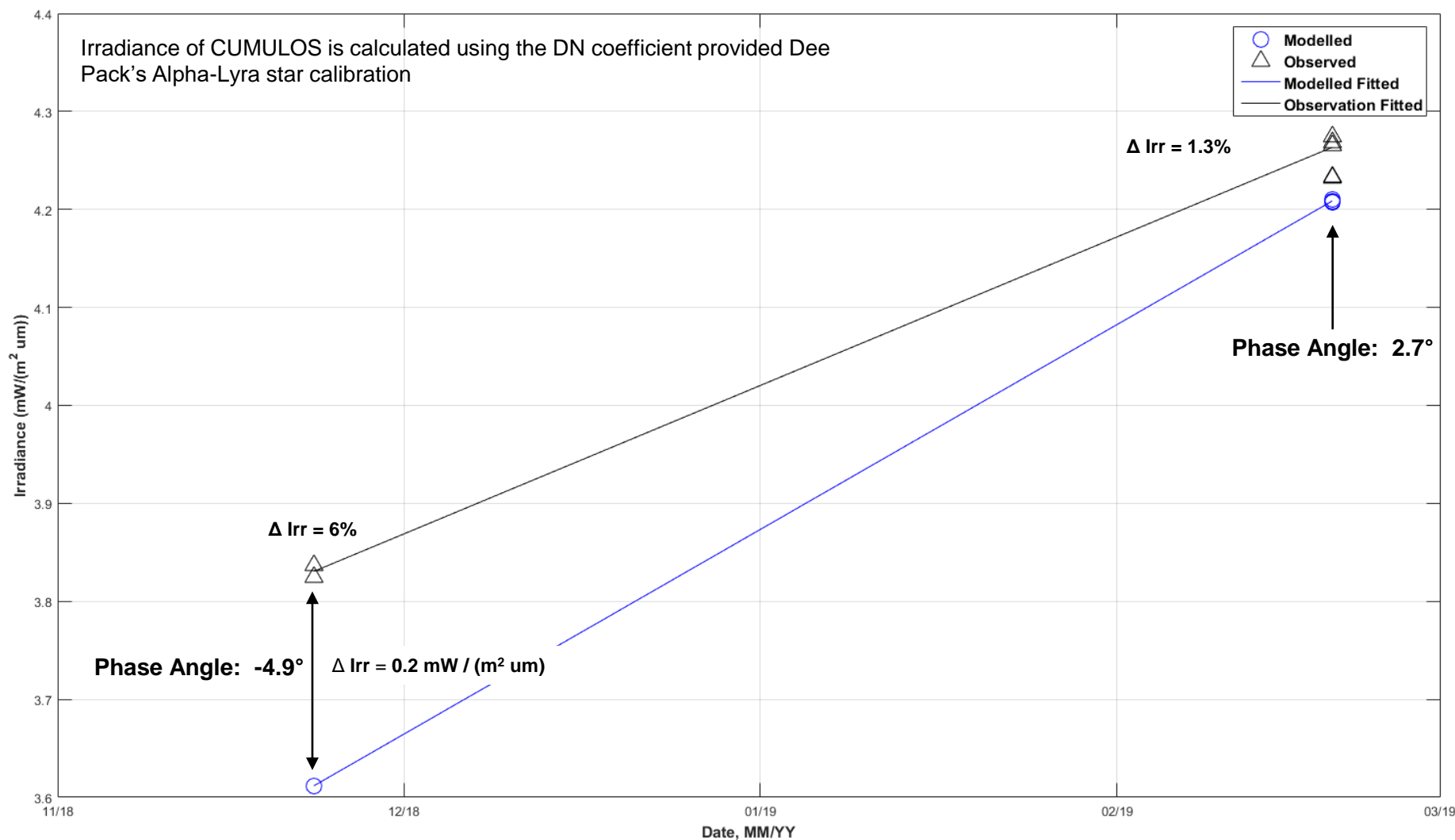
Comparing VIIRS Model to Observations

S-NPP VIIRS Band M1

Phase Angle Range: -50.4° to -55.4° (40 Collects over 5 years)

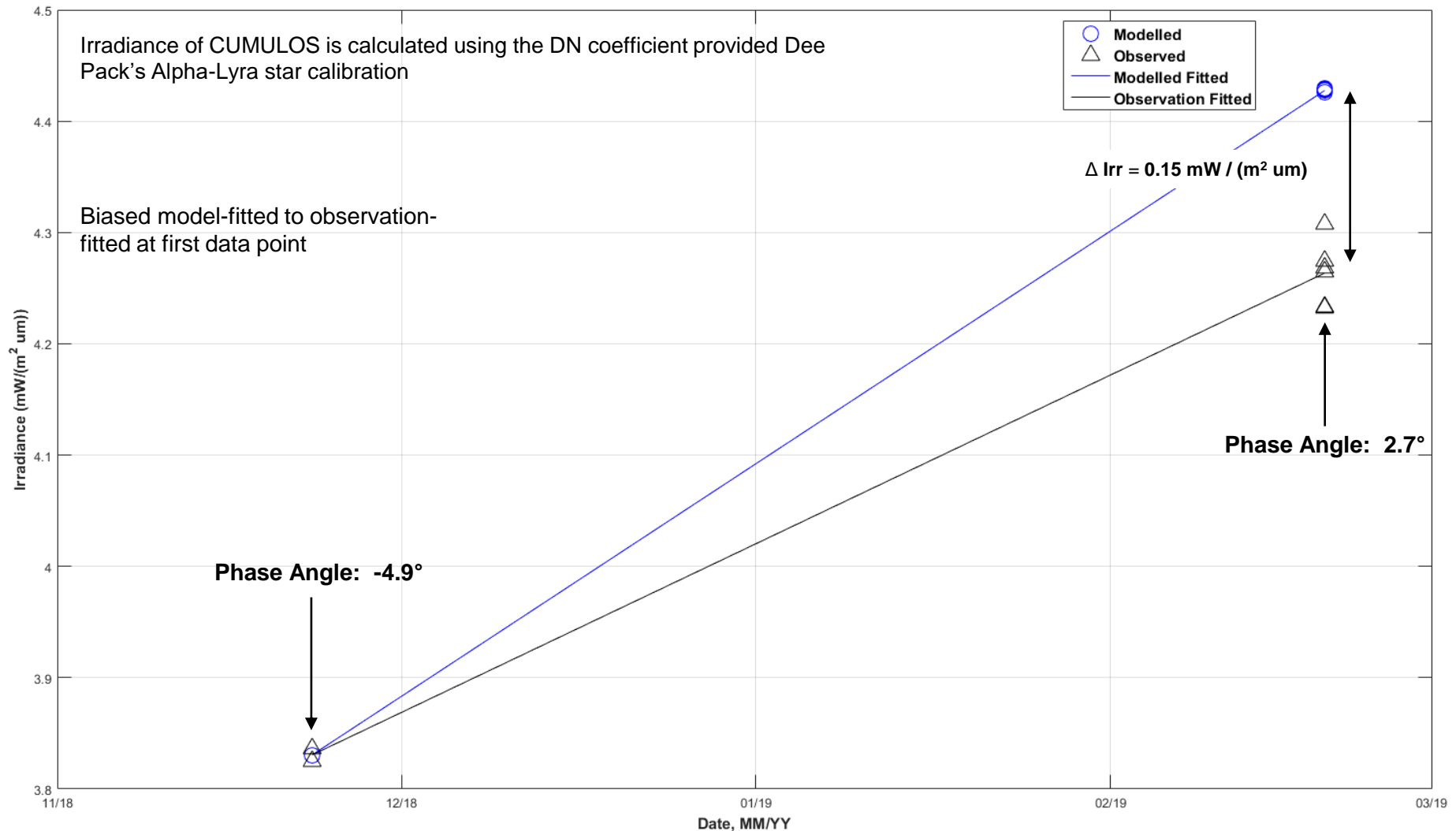


Comparing CUMULOS Model to Observations



Comparing CUMULOS Model to Observations

Biasing Model towards Observation



Summary



- The Moon was used to assess the CUMULOS Visible Camera upon ISARA CubeSat
 - *There have been unexpected anomalies using the Moon*
 - Ghost Imaging
 - Light Sensitivity correlation to FPA detectors (DN Dip)
 - *Based on the ROLO model it has been determined that the CUMULOS Visible camera requires additional radiometric biasing for 116 us exposure time*
- Desired Future:
 - *Perform a more extensive validation of our implementation of the ROLO Model with GSICS Implementation of the ROLO (GIRO) Model values*
 - *Apply this ROLO work to future Aerospace Corp and Gov/Private industry sensors*
- Lessons Learned:
 - *More Lunar Collects are always good (we already knew this)*
 - *In the absence of using ROLO one can still use the Moon for more than just radiometric trending (Expect the unexpected)*
- COTS sensors without onboard calibrators will only increase given the pLEO/ CubeSat mentality and using the Moon to radiometrically assess said sensors is of absolute importance



Thank You