



Electro-Optical Lyot Filter Automation

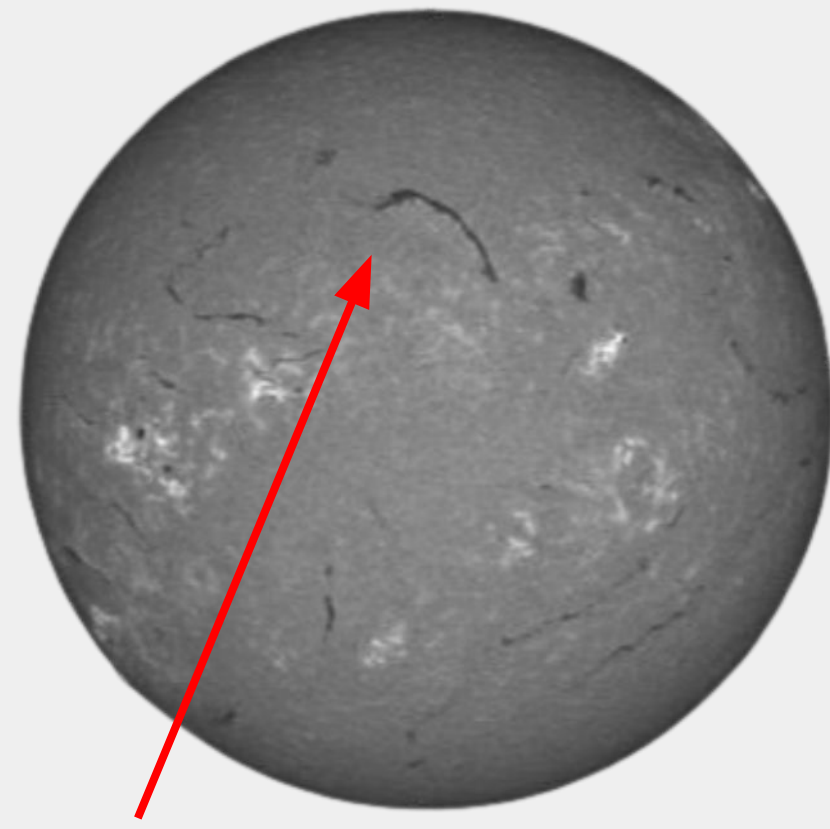
Lyot Filter Demonstration Instrument (LFDI) – HESTO

Mitchell Jeffers, Elizabeth Bernhardt, Nir Patel

mjeffers@ucar.edu, embern@ucar.edu, nirpatel@ucar.edu High Altitude Observatory (HAO), National Center for Atmospheric Research (NCAR), University Corporation for Atmospheric Research (UCAR)

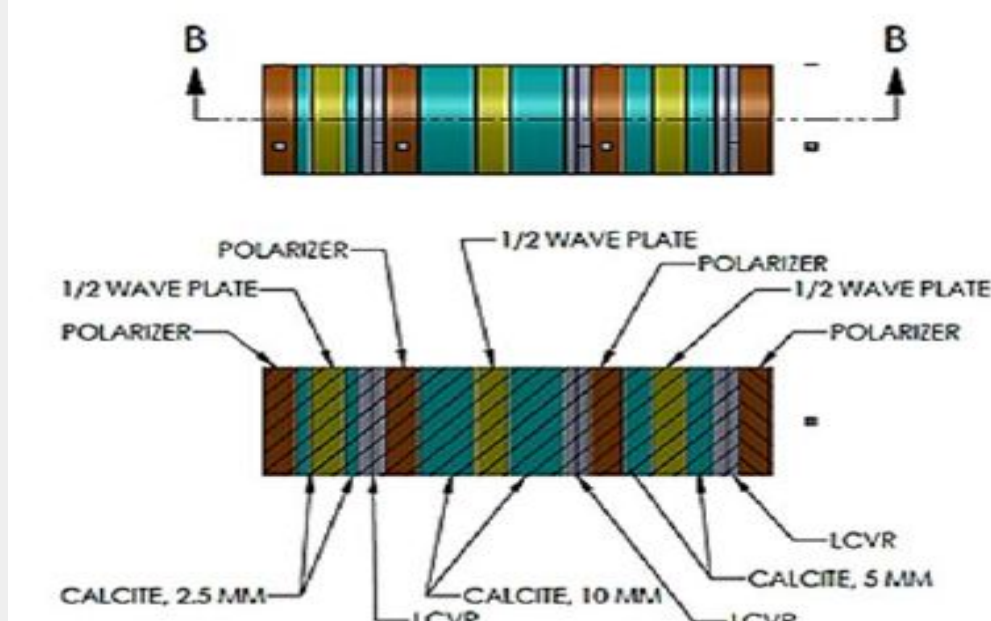


Lyot Type Filters and LFDI Background

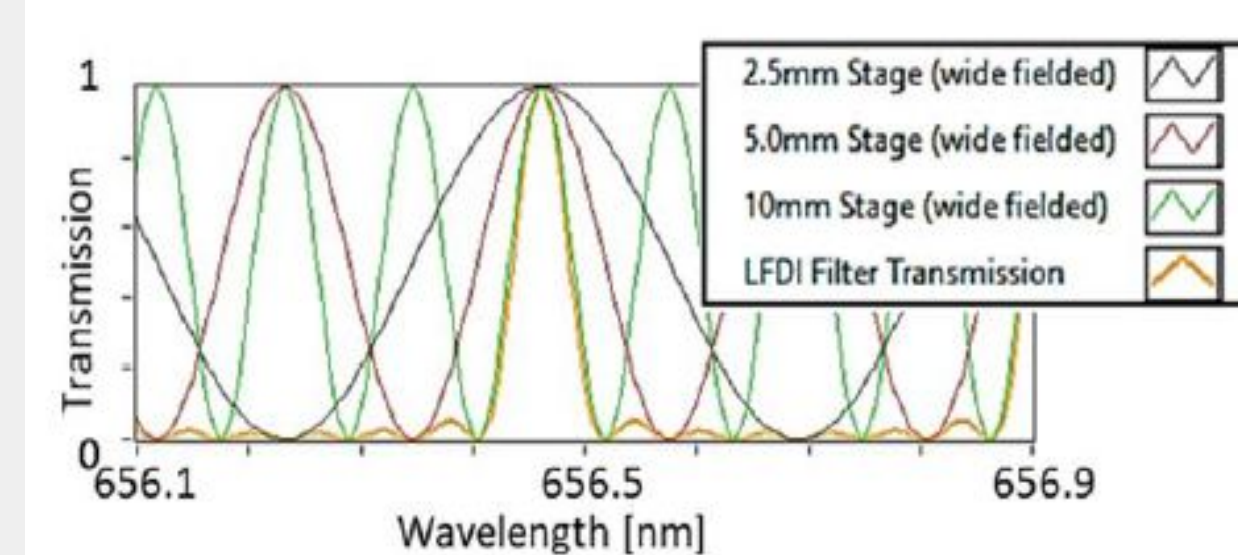


Lyot filters offer solar physicists the advantage of full-disk imagery combined with spectropolarimetry, enabling the observation and analysis of solar activity in detail. This technology not only captures physical interactions on the sun's surface but also provides mapping of spectral characteristics such as Doppler shifts. This capability, along with the extended observational periods afforded by space-based operations, drives the development of the Lyot Filter Demonstration Instrument (LFDI) at NSF-NCAR's High Altitude Observatory.

Filaments are one type of phenomena solar astronomers are interested in observing, as CMEs tend to be associated with these structures.

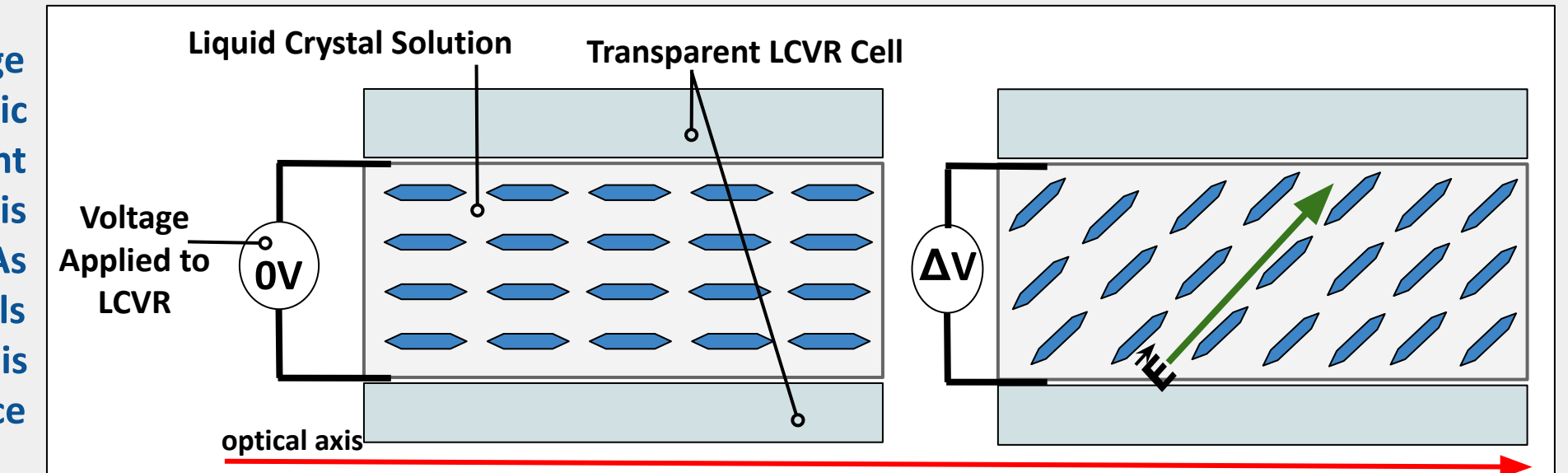


Complete stack of LFDI Filter components (left), which is made up with the 3 different stages. Each stage has a different thickness which when combined transmits at the central wavelength of 656.46 nm commonly known as H-alpha (right).



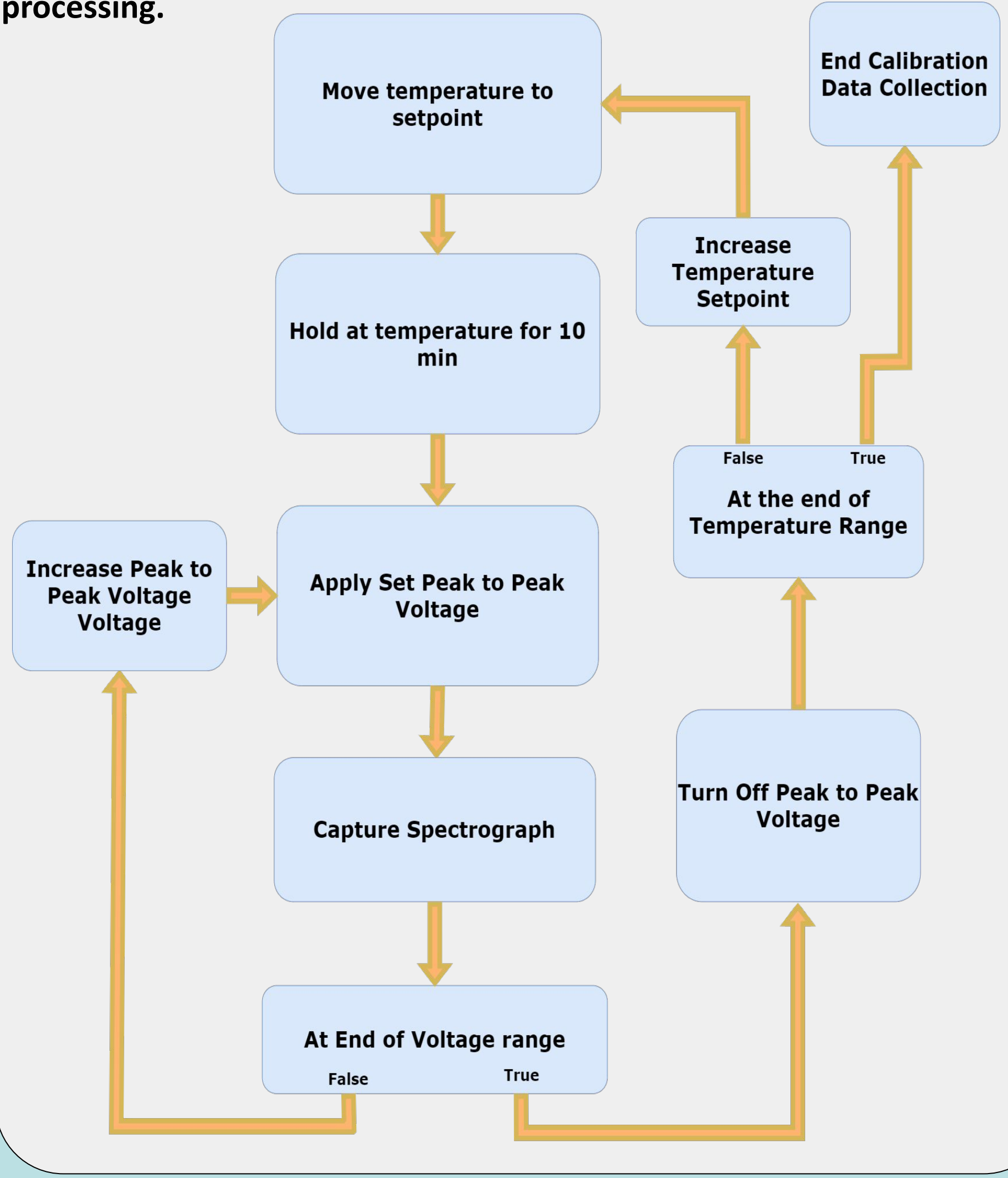
Orientation of Liquid Crystals as voltage is applied. When there is no electric field the crystal alignment is dependent on the alignment layer which is determined during manufacturing. As the peak voltage changes the crystals align with the electric field that is induced, giving a level of retardance (right)

LCVRs consist of a transparent cell filled with a liquid crystal solution that acts as a waveplate. When voltage is applied, the liquid crystals align with the electric field, altering the polarization state of the light and therefore the transmission pattern, a phenomenon known as retardance. Similarly the polarization state changes as the LCVR is subjected to temperature changes. The LFDI design leverages this electro-optical tuning to overcome thermal and tuning challenges faced by previous Lyot filters. Calibrating the LCVRs' retardance is essential for optimal performance of each stage. A significant advancement in the LFDI project is the development of software to automate this calibration process via the Tuning Control Board (TCB). Once calibrated, the Lyot filter can be controlled, to tune to user defined wavelength and to compensate for temperature variations during operation.

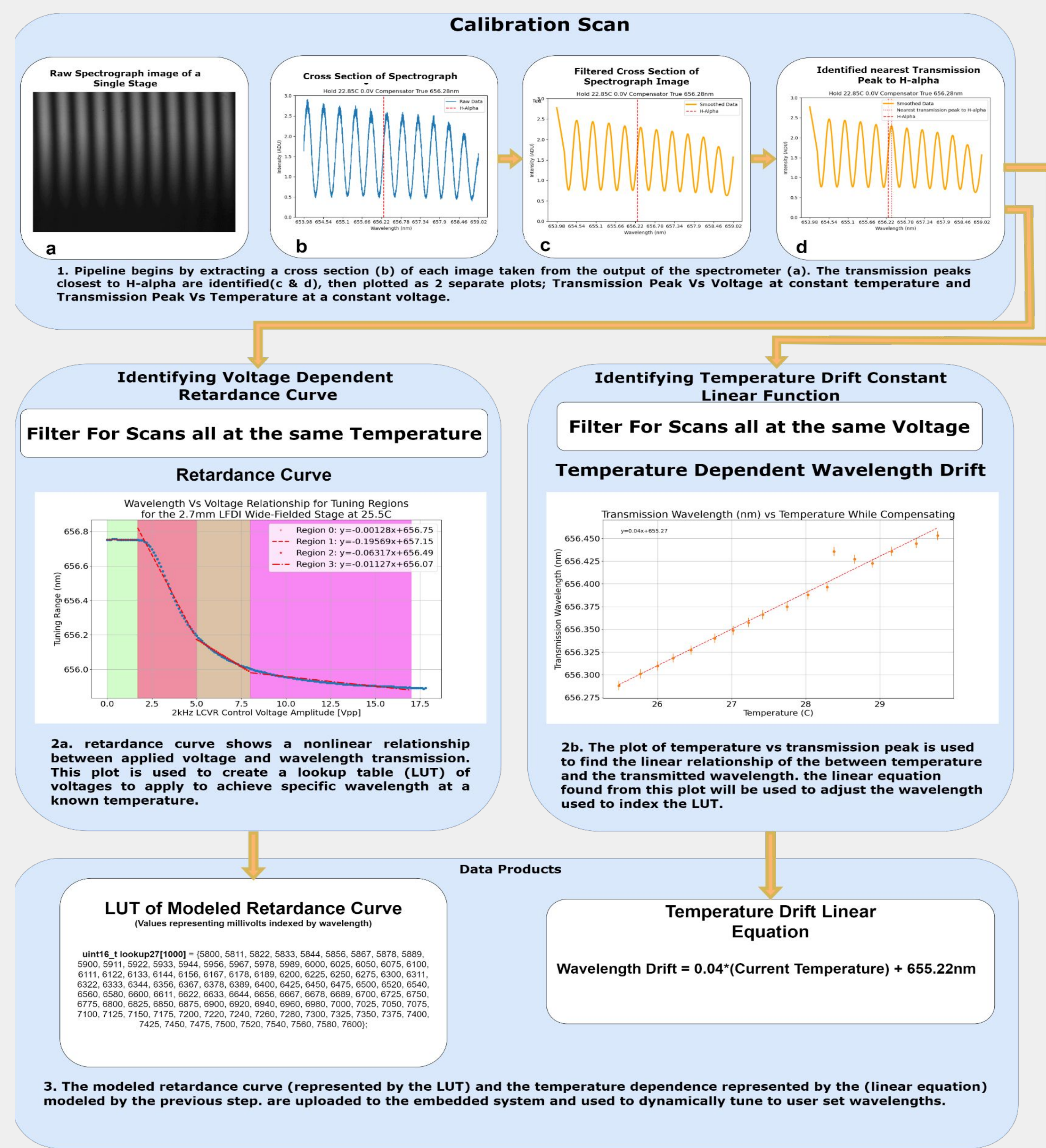


Calibration Data Collection

To tune a Lyot filter, the characteristics of each stage—dependent on calcite thickness, LCVR temperature, and voltage—must be identified. This is done by placing a single stage in front of a spectrometer, warming it to a set temperature, applying a square wave to the LCVR, and capturing images at different voltage-temperature pairs for processing.



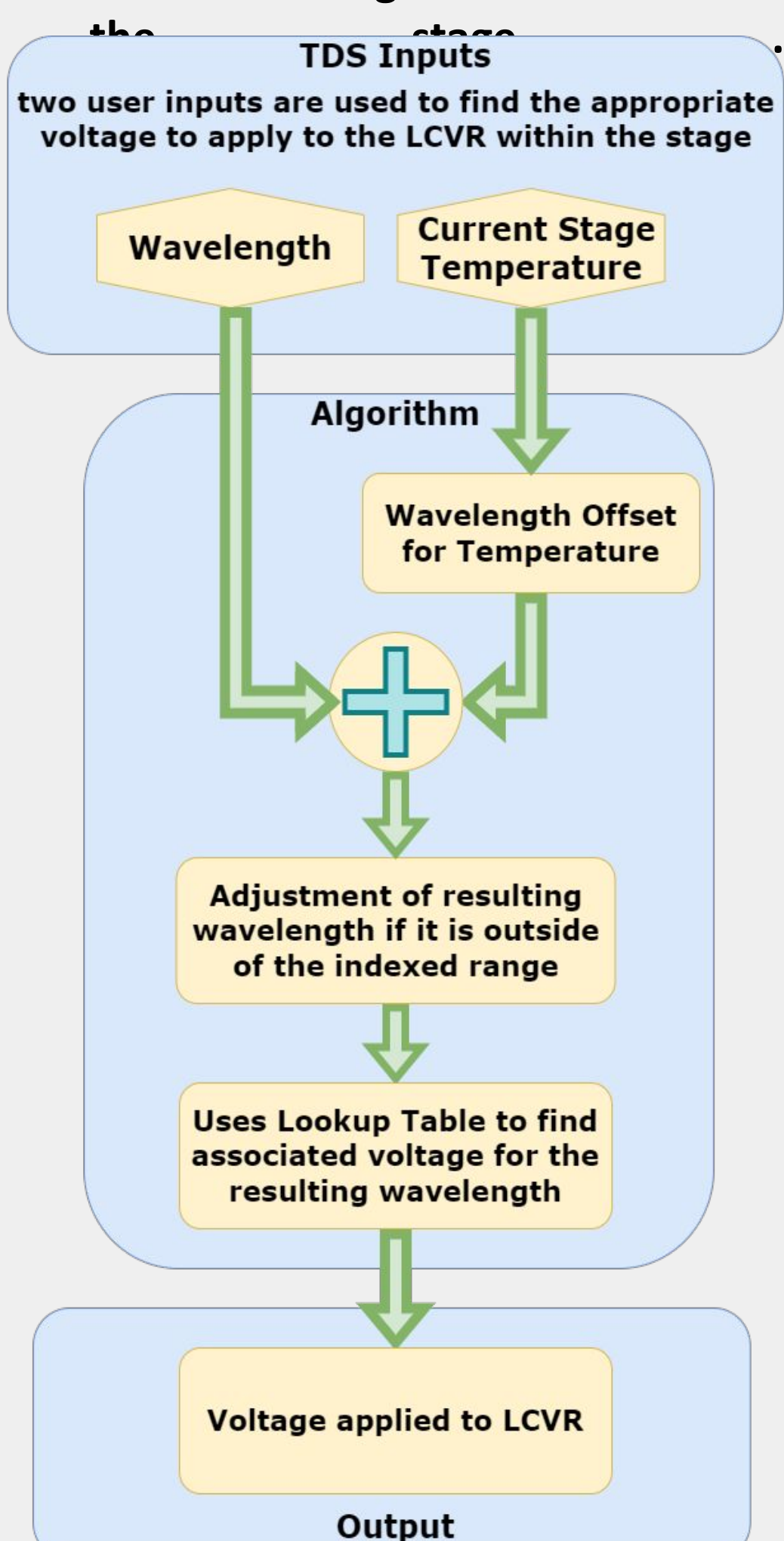
Calibration Data Processing Pipeline



In-Flight Tuning Determination System

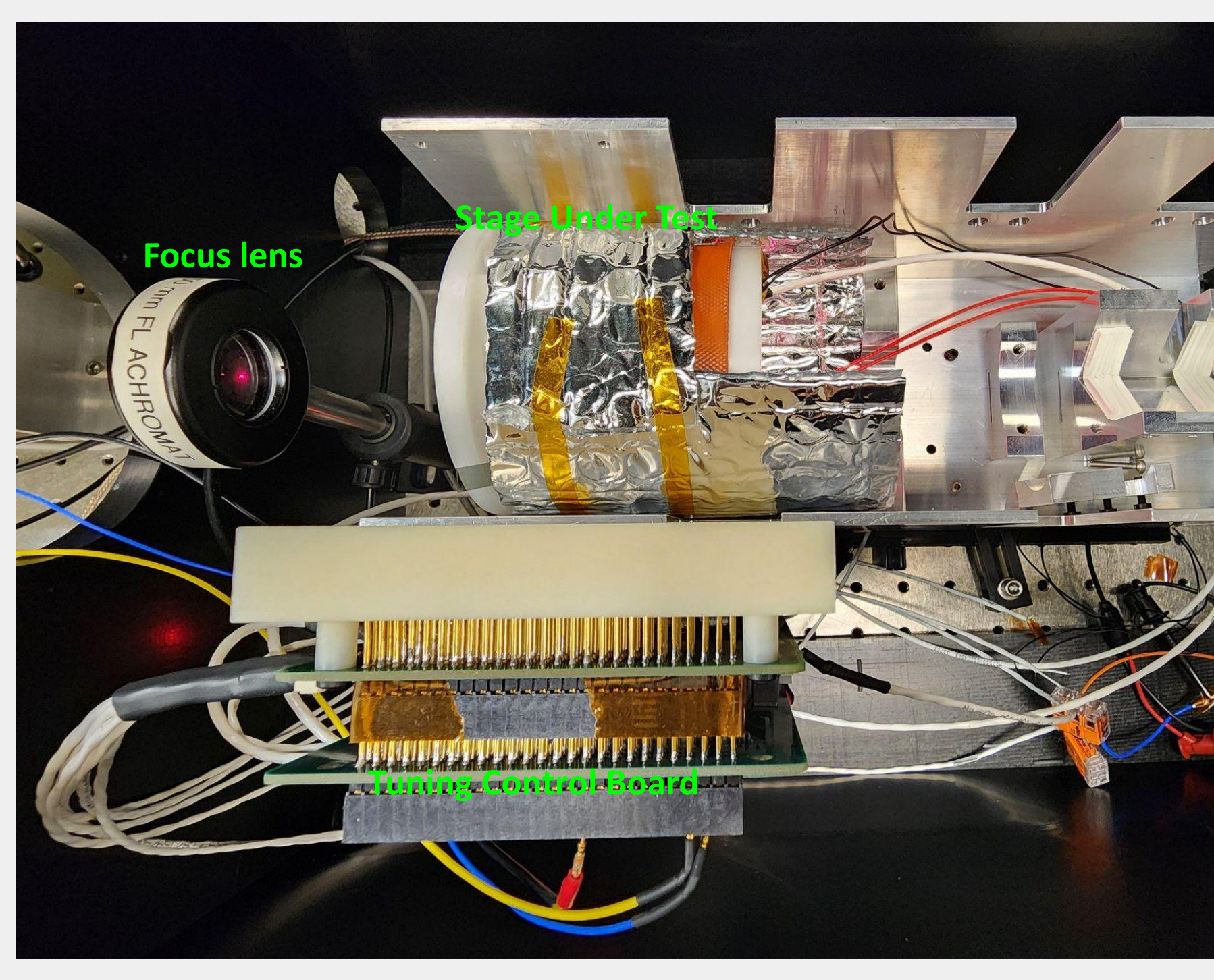
During the operation of the Lyot Filter Demonstration Instrument (LFDI) a tuning determination system is used to find the appropriate voltage to apply to the LCVR to achieve a desired wavelength at the current temperature of the stage.

- 1. Wavelength defined by user which the Lyot filter will be tuned for. The current stage temperature is measured on the outside of the stage with a (TMP117) digital temperature sensor.
2. a) The TDS uses the current temperature to calculate the wavelength offset, which is then added to the desired wavelength to get an adjusted wavelength.
b.) The adjusted wavelength is used to index the LUT finding the correct voltage to apply to the LCVR to generate the desired wavelength.
3. The associated voltage is applied to the LCVR for tuning.

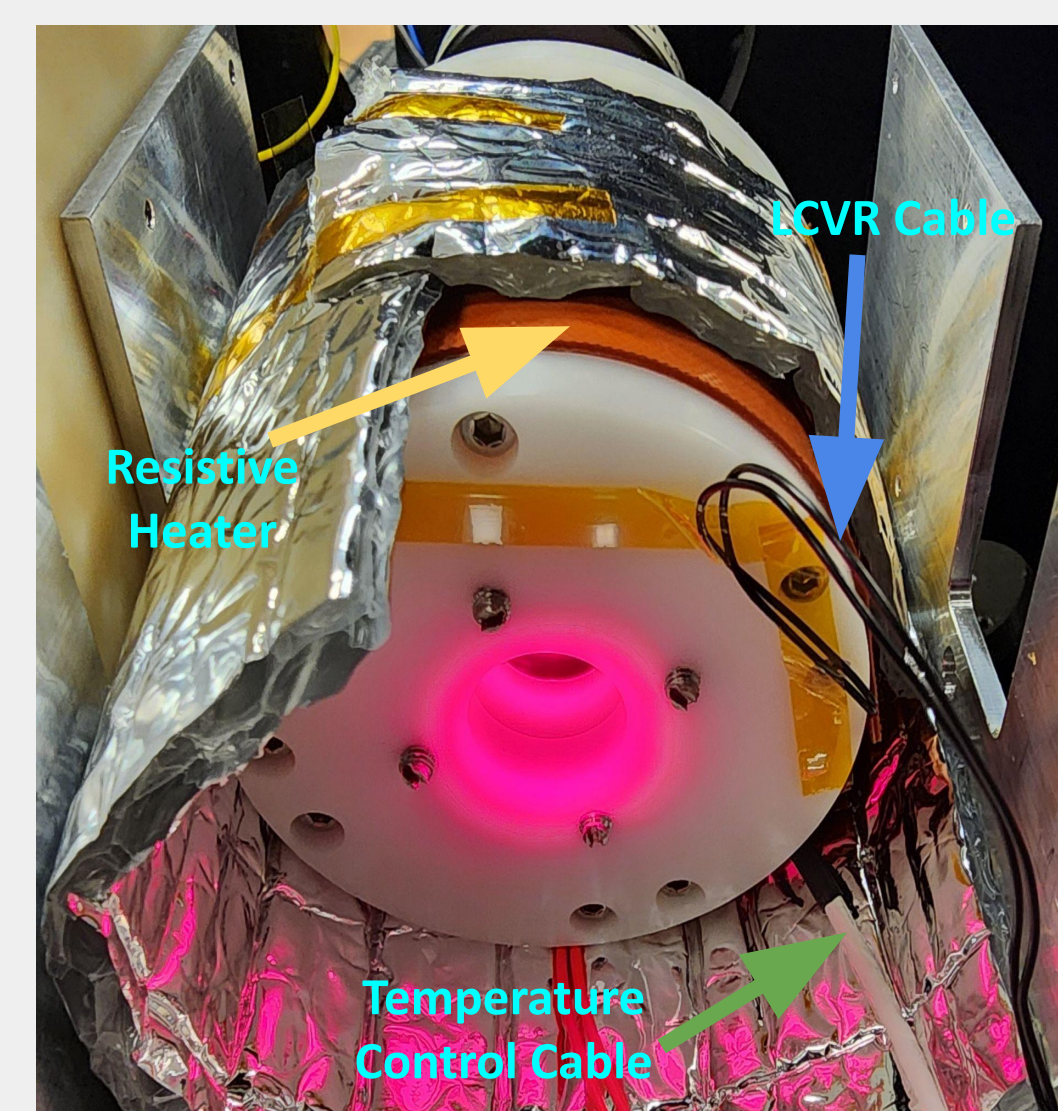


For LFDI three instances of the TDS are ran simultaneously (one for each stage) to achieve full tuning of the assembly.

Lab setup for the calibration of each LFDI stage



Lyot Filter Stage



Acknowledgements

This work was supported by NSF/NCAR and NASA Heliophysics Research Program grant 20-HTIDS20-0004. The authors gratefully acknowledge the support of Principle Investigator Scott Sewell, Opto-mechanical support from Andrew Carlile and Patrick Zamarly, and Electronic Support From Damon Burke.

Scan QR code to access open source embedded software, calibration pipeline printed circuit board files.



Results

Using the calibration process outlined in this poster the system is able to maintain a tuning stability of +/- 0.01nm RMS over a temperature delta of 10C.

