



AERO & VISTA

Demonstrating HF Radio Interferometry with Vector Sensors

Presenting author: Michael Hecht

(On behalf of VISTA PI Frank Lind and AERO PI Phil Erickson)

Co-authors:

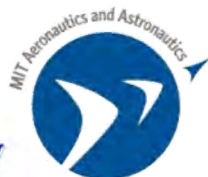
Mary Knapp, Geoff Crew, Ryan Volz, John Swoboda (MIT Haystack Observatory)

Frank Robey, Mark Silver, Alan Fenn (MIT Lincoln Laboratory)

Ben Malphrus (Morehead State University)

Kerri Cahoy (MIT AeroAstro Dept.)

1



Target: The Radio Aurora



High-latitude aurora viewed from the ISS (~350 km altitude)
September, 2011 (NASA ISS Crew Earth Observations experiment)



Mission summary



AERO and VISTA are twin 3U CubeSats that will launch & deploy together into a polar orbit, using drag to control separation.

Individually, they will answer key scientific questions about the nature and sources of auroral radio emissions.

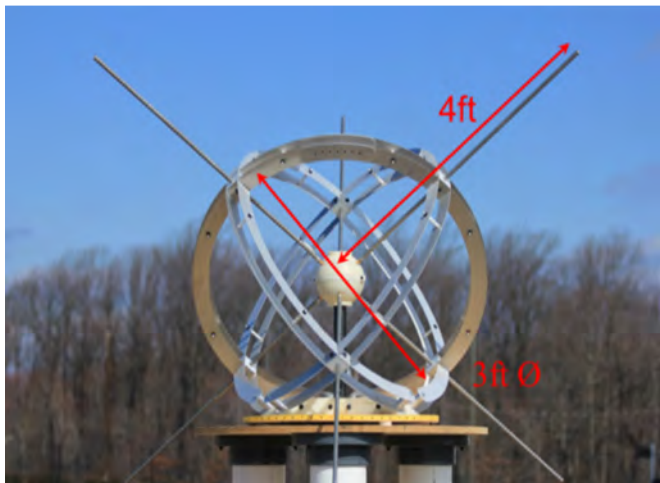
Together, they will demonstrate interferometric imaging, beamforming, and nulling using electromagnetic vector sensors at low frequencies (50 kHz – 5 MHz).

A key goal is to demonstrate that interferometric arrays of VS's will maintain sensitivity in the presence of terrestrial interference.

- This would allow low frequency interferometers be placed in LEO, reducing cost and increasing data volume.



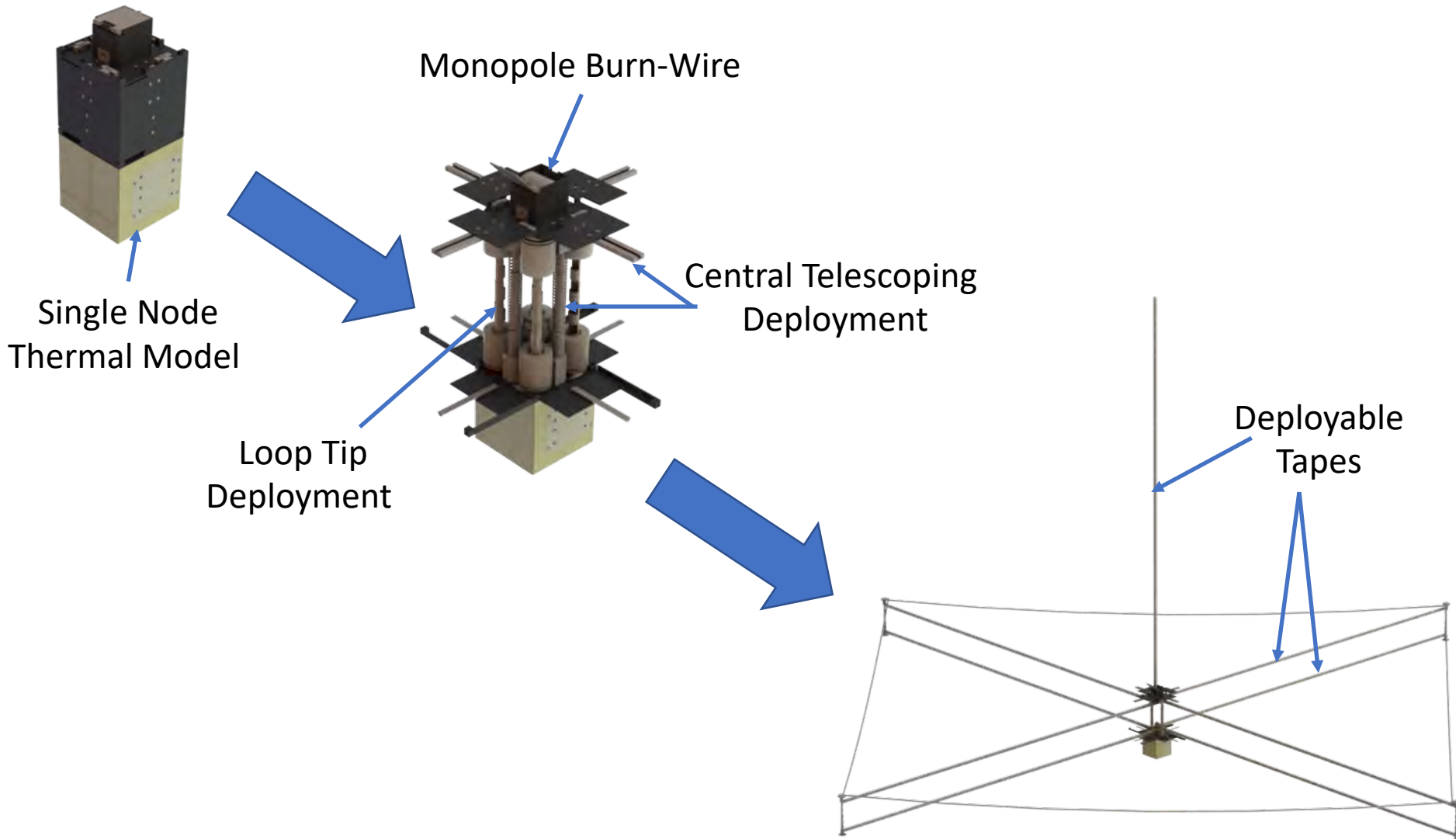
Electromagnetic Vector Sensing



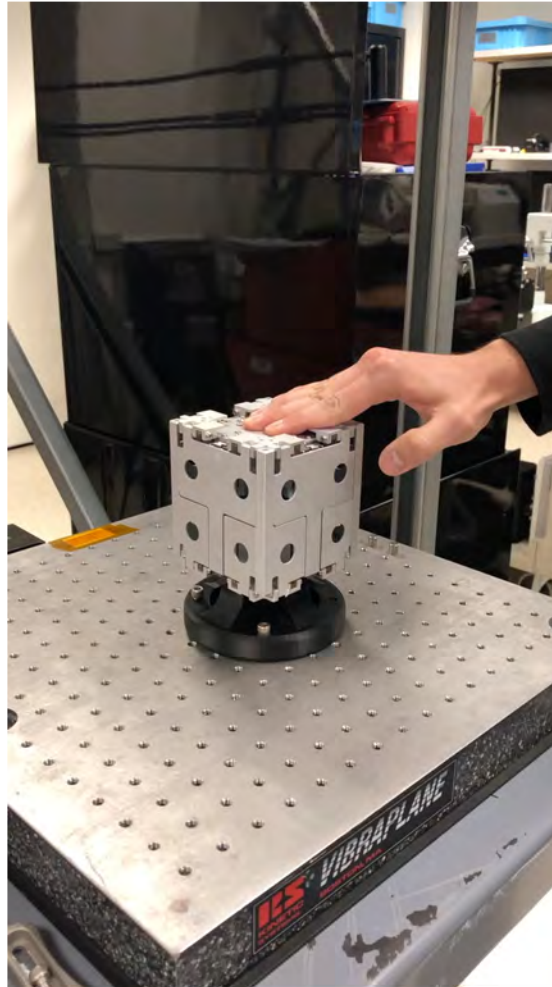
- 3 dipoles + 3 loops (electrically small)
- Measures full E and B field vectors
 - $\mathbf{E} \times \mathbf{B} = \mathbf{S}$ (Poynting vector)
- Determines source intensity, polarization, and direction (to a few degrees)
- Some imaging capability

**Spatially resolved detection
with a single electrically
small sensor!**

Vector antenna deployment



Central Telescoping Deployment



Deployable Vector Antenna Prototype



Mission Objectives

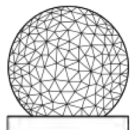


Table 1: Top level objectives for AERO and VISTA missions.

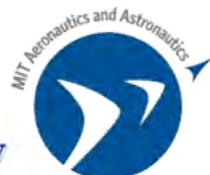
AERO	VISTA
AO1: Characterize auroral radio emissions in the ionosphere	VO1: Demonstrate vector sensor interferometry (VSI) in space
AO2: Connect radio emissions to overall auroral geospace system	VO2: Apply VSI to auroral radio emissions
AO3: Demonstrate polarimetric HF radio signal detection [Tech validation]	VO3: Characterize low Earth orbit (LEO) radio frequency interference (RFI) environment [at HF frequencies]

Nominal launch date in 1st Quarter of 2022

Backup



MIT
HAYSTACK
OBSERVATORY



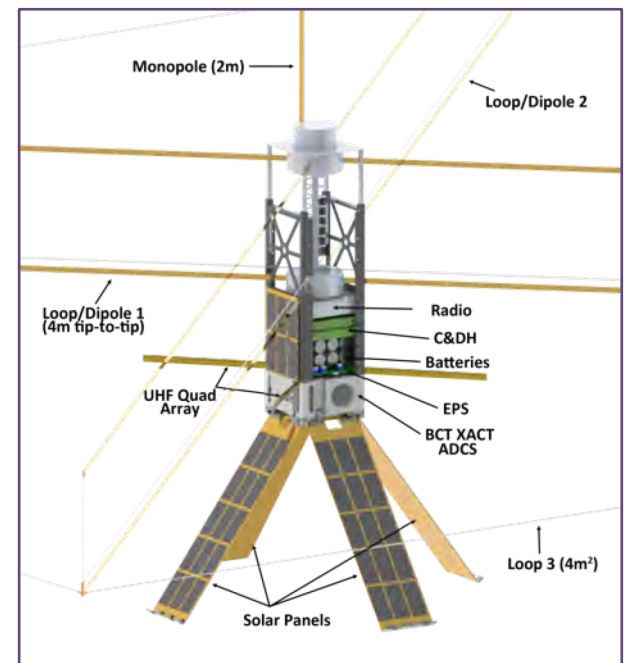
AERO AURORAL EMISSIONS RADIO OBSERVER

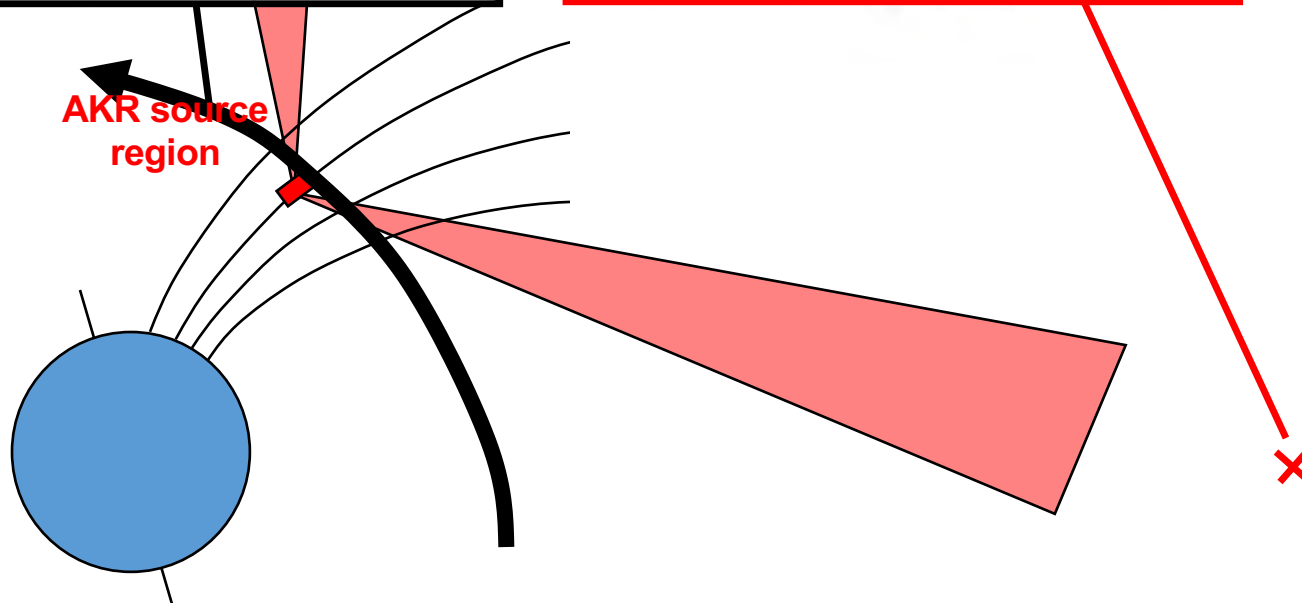
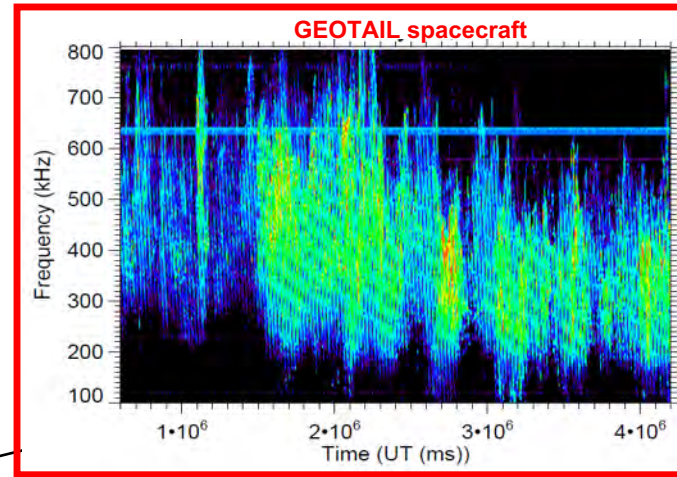
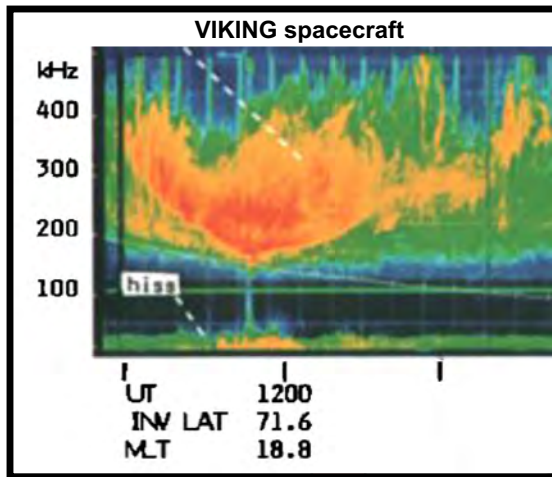
VISTA VECTOR INTERFEROMETRY SPACE TECHNOLOGY USING AERO

Twin 3U CubeSats will be placed in polar orbit for a nominal 90-day mission to:

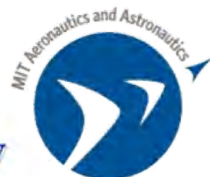
- Qualify and validate a novel electromagnetic Vector Sensor (VS)
- Answer key scientific questions about the nature and sources of auroral radio emissions.
- **Demonstrate low frequency interferometry**

- Orbiting at >400 km, above the ionospheric peak, allows sensing of radiation not visible from Earth
- Each deployable *vector sensor* (50 kHz – 5 MHz) can determine direction of arrival of auroral radio emission
- Together, they demonstrate how interferometry enables high resolution imaging
- An Auxiliary Sensor Package includes magnetometers and optical aurora sensors





MIT
HAYSTACK
OBSERVATORY



Developments in auroral radio emissions

Type	frequency	polariz'n	outstanding problems
Auroral Hiss	<1 MHz (below f_{ce})	Right (W-mode)	Structure: hissers, LF cutoff etc. Dayside/nightside Source altitudes
AKR	50-750 kHz (below f_{ce})	Right	Confirm connection How W-mode generated? Ducted? Area illuminated? Remote sensing application?
Auroral Roar	$2f_{ce}$, $3f_{ce}$, $4f_{ce}$, $5f_{ce}$	Left (sometimes Right?)	Nonlinear mode conversion?
Medium Freq Burst	1.5-4.5 MHz (above f_{ce})	Left (L-mode)	Generation mechanism? Use for substorm onset timing? Connected to Langmuir cavitation?

CubeSat Deployable Vector Sensor

